# UCLA

## Exaflop, Petawatt, and Terabar physics

B. Polllock et al. PRL 2015

**W.B.Mori** University of California Los Angeles (UCLA)

Departments of Physics and Astronomy and of Electrical Engineering

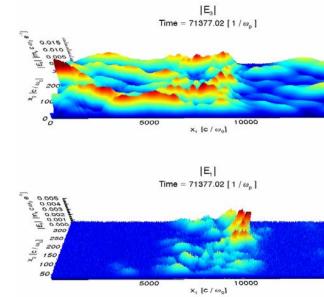
Particle-in-Cell and Kinetic Simulation Software Center (PICKSC)

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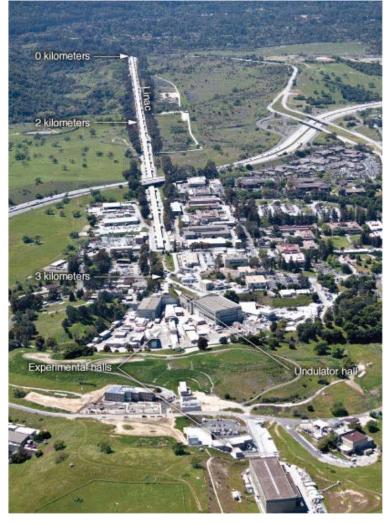






Today's tools for scientific discovery can be large, complex, and expensive: Accelerators

#### LCLS: 4th generation



#### LHC: "Last" generation?

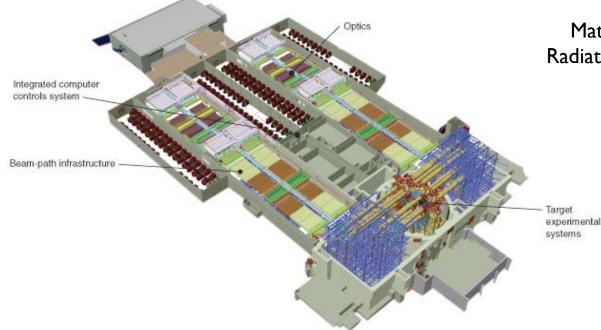


#### Today's tools for scientific discovery can be large, complex, and expensive: National Ignition Facility: 1.8 MJ, ~4ns, 192 beams









Matter Temperature > 10^8 K Radiation Temperature >3.5 10^6 K Densities >10^3 g/cm3 Pressures >.1Tbar .5PW

# NSF, DOE, NASA, NNSA invests in computers that cost ~\$250,000: Large, complex, and expensive

Blue Waters - Cray XE/XK hybrid

INT

24140 XE Compute Nodes 2× 16 core AMD 6276 @ 2.3 GHz R<sub>peak</sub> 7.1 PFlop/s

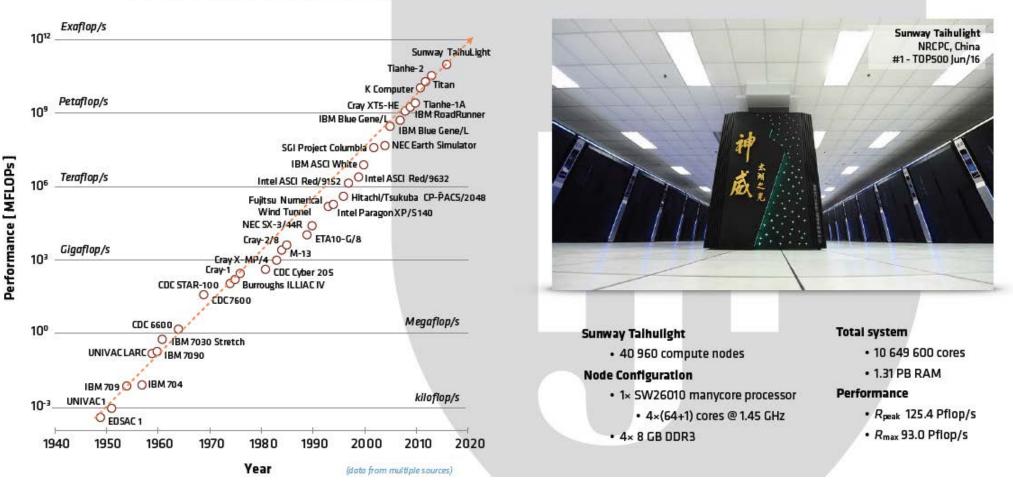
3072 XK Compute Nodes 1× 16 core AMD 6276 @ 2.3 GHz 1 × Nvidia Tesla K20 GPU R<sub>peak</sub> 4.51 PFlop/s

R<sub>peak aggr</sub> 11.61 Pflop/s

Progress in science is often driven by riding up a Moore's Law curve. Need to be using the previous generation of that tool.

You just don't wake up one day and say you want to use the LHC, NIF, or a leadership class computer.

# **Computing Power (R)Evolution**



**High Performance Computing Power Evolution** 

High Performance Computing Power Evolution (data from multiple sources)

Progress in science is often driven by grand challenge questions coupled with discovery driven research and advances in tools for discovery.

Let me show show two in high energy density plasma physics.

Large Hadron Collider: 30 km in circumference, \$10 Billion +

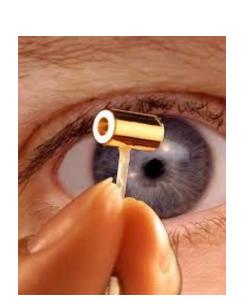
# What is next? Use plasma waves? PLASMA BASED ACCELERATION

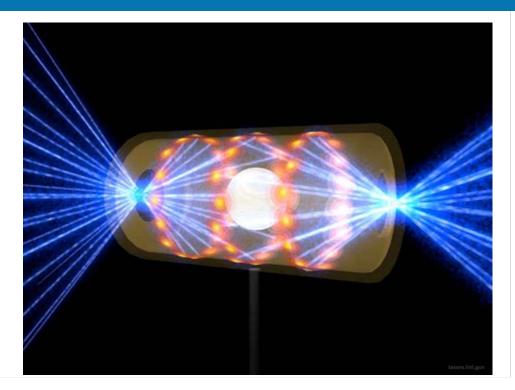


#### Can fusion ignition be achieved?

National Ignition Facility (NIF) is a 1.8 MJ, 192 laser facility that was built to demonstrate fusion ignition, \$3.5 Billion + Can the laser-plasma interactions be mitigated or controlled?

#### NONLINEAR OPTICS OF PLASMAS





#### What is high energy density plasma physics?

Why are both plasma-based acceleration and the nonlinear optics of plasmas considered high-energy density plasma research?

- High energy density, means high pressure
  - What is a high pressure?
    - MBar? GBar?
  - Need a dimensionless parameter
    - In plasma physics an important parameter is the number of particles in a Debye sphere (which is directly related to the ratio of the kinetic energy of an electron to the potential energy between particles). It measures the discreteness of the plasma.

$$\frac{4\pi}{3}n\lambda_d^3 \equiv N_D = 2.1 \times 10^3 \frac{T_{keV}^2}{P_{MBar}^{1/2}}$$

- When the pressure exceeds ~1 MBar then the discrete nature of the plasma becomes important:
  - ND is not "infinite"
- Discreteness makes developing computational methods difficult

#### What is high energy density plasma physics?

Why are both plasma-based acceleration and the nonlinear optics considered high-energy density plasma research?

- High energy density means high pressure
  - What is a high pressure?
    - MBar? GBar?
- An intense laser or particle beam can have a high energy density (high intensity).

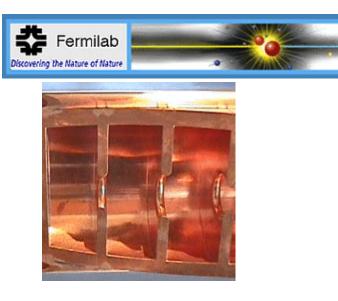
$$\mathcal{E} = \frac{E^2 + B^2}{8\pi} \qquad \mathcal{I} = c\mathcal{E} = c\frac{E^2}{4\pi}$$

- It turns out that for radiation pressures corresponding to ~GBar that a laser (or particle beam) causes individual electrons to move at relativistic energies.
- Relativistic particles and trajectory crossing make computational modeling difficult

#### A major driver for HEDP remains the goal of reducing the size and cost of expensive particle accelerators









Energy, efficiency, charge, beam quality are the important metrics.



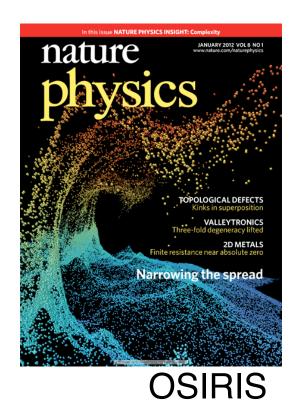
Plasma based acceleration has been a driver for the field of short-pulse laser and beam plasma interactions which is at the forefront of basic science



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# The synergy between simulation and experiment has "accelerated" the rate of discovery







Each article contained experimental results whose interpretation was supported from simulations.

#### QUICKPIC

#### Particle Accelerators Why Plasmas?

## **Conventional Accelerators**

## <u>Plasma</u>

- Limited by peak power and breakdown
- 20-100 MeV/m
   20km /0.8 TeV

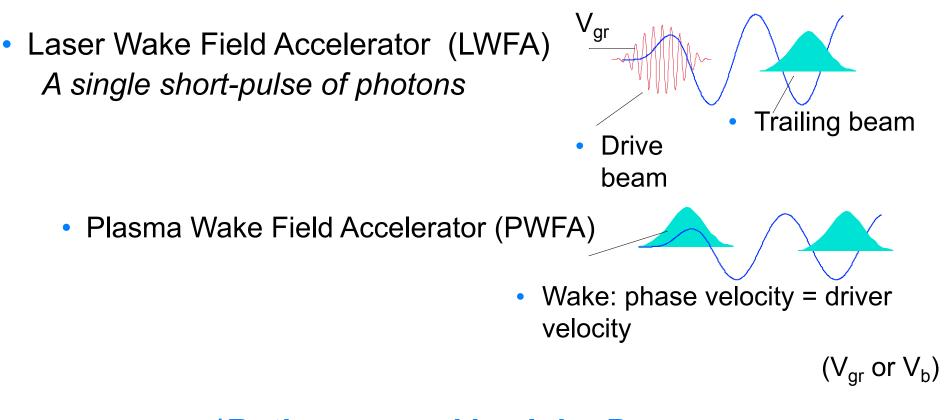
- No breakdown limit
- 10-100 GeV/m

Based on plasma wave wakefields ("longitudinal plasma waves"

#### Create relativistic plasma waves as wakefields: "Fast boats"



#### Create relativistic wakefield using lasers or particle beams: Concepts for plasma based accelerators\*



\*Both proposed by John Dawson LWFA: Tajima and Dawson 1979 PWFA: Chen, Dawson et al., 1985

#### Use waveframe or quasi-static variables Sprange, Esarey, and Ting 1990

For a fixed driver shape the wake can be calculated. The wake only changes if the driver shape changes. The driver's shape changes very slowly.

#### Use appropriate variables

• Transform from:

• Transform to:

$$(\xi = z - v_{\phi}t, x, y; s = z)$$

#### Meaning of new variables

- .  $\xi = z v_\phi t_{\rm ~~is~the~distance}$  from front of the driver
- $S = \mathcal{Z}$  is the distance the driver has propagated into the plasma

Mathematical meaning of quasistatic approximation

$$\partial_s << \partial_{\xi}$$

Let the wake move at c and make the quasi-static approximation

$$E_{z} = -\partial_{z}\phi - 1/c \ \partial_{t}A_{z}$$
$$F_{z} \approx -q\partial_{\xi}(\phi - A_{z})$$
$$\vec{F}_{\perp} = q\left(\vec{E}_{\perp} + (\vec{v}_{b} \times \vec{B})_{\perp}\right)$$
$$\vec{v}_{b} = \hat{z}c$$
$$F_{\perp} \approx q(-\nabla_{\perp}(\phi - A_{z}))$$

**Psuedo-potenial** 

$$\psi = (\phi - A_z)$$

Don't choose a gauge where

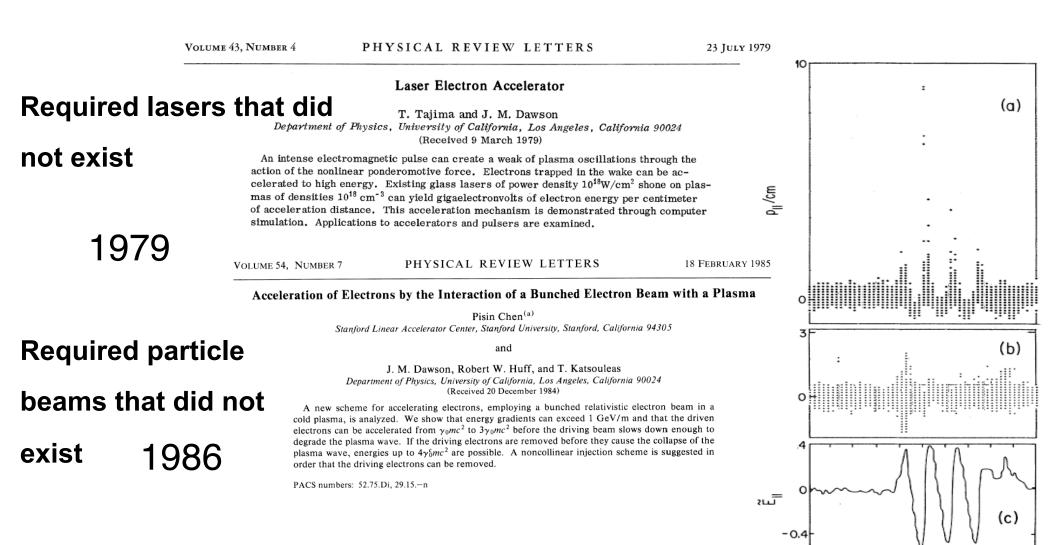
$$\phi = A_z$$

Forces on relativistic particle

$$F_z = -\partial_\xi \psi$$

$$F_{\perp} = -\nabla_{\perp}\psi$$

#### Humble beginning of a new subject area!

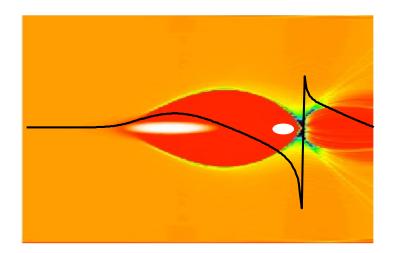


512

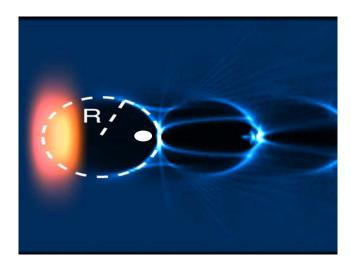
The simulations of Tajima and Dawson would take <1 second on my laptop! ~5000 particles for 500 time steps

#### Today: short pulse drivers and nonlinear 3D wakes

#### Driven by an electron beam



Driven by a laser pulse



Called blowout or bubble (Rosenzwieg et al. 1990, Mori et al. 1992, Puhkov and Meyer-ter-Vehn 2002, Lu et al. 2006, 2007)

Need a nonlinear description of these wakes (Not till Lu et al. 2006)

Ideal for accelerating electrons/not for positrons

Very stable wakes!

Experimental progress in the last decade has been in this regime.

#### Wake is described by an equation for the radius of the bubble:

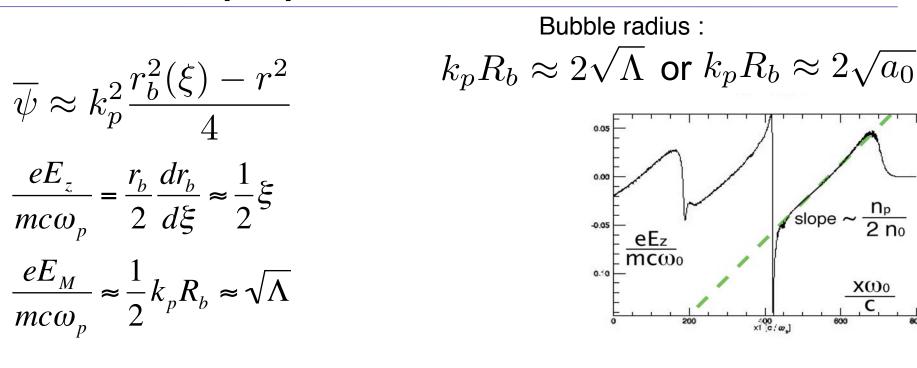
 $dr_b$  $d\xi$ 

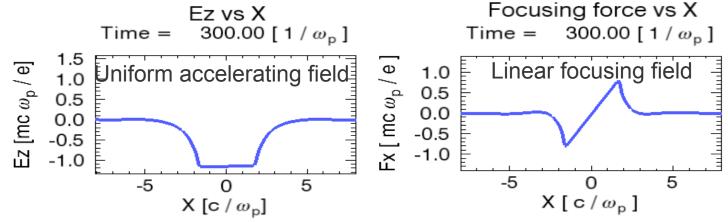
<u>n</u>p 2 no

XWo C

slope

Relativistic blowout regime for blowout radius and for large maximum radius the trajectory of rb is a circle: Bubble Lu et al.PRL 16, 16500 [2006]





#### **Transverse Dynamics and Beam Quality**

• Emittance  $\epsilon_n$  = phase space area and a measure of its ability to get focused:

px

σ

Inside a plasma wake a single particle oscillates as:

The spot size of a beam in vacuum evolves as:

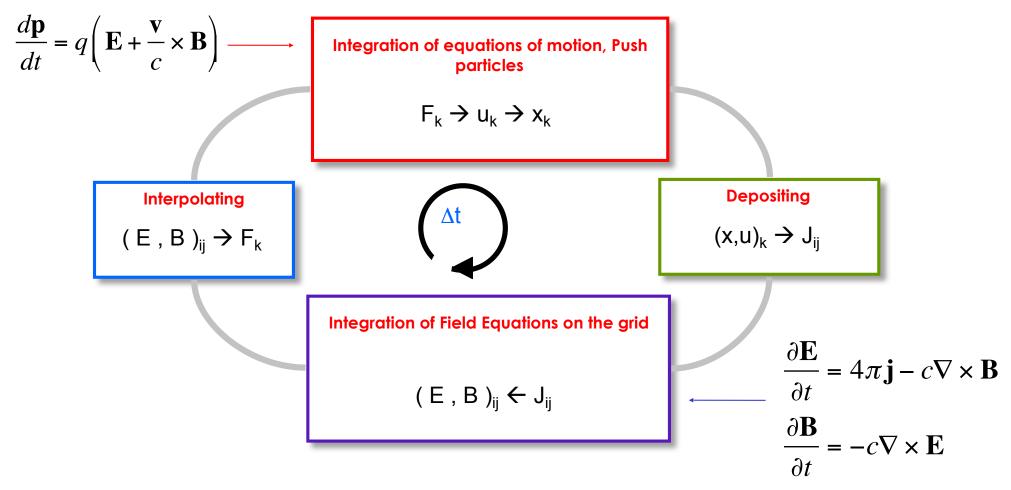
$$\frac{dP_{\perp}}{dt} = q(-\nabla_{\perp}\psi)$$

 $\overrightarrow{\mathbf{x}} \ \ \sigma_r = \sqrt{\left(1 + (\frac{z}{\beta^*})^2\right)} \ \ \text{where} \ \ \beta^* = \frac{\sigma_r^2}{\epsilon_n} \gamma$ 

• If the focusing force is "linear" AND radial in the transverse coordinates then

$$\frac{d^2 x_{\perp}}{dt^2} + \omega_{\beta}^2 x_{\perp} = 0$$
$$k_{\beta} \equiv \frac{\omega_{\beta}}{c} \qquad k_{\beta} = \alpha \frac{k_p}{\sqrt{2\gamma}}$$

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#### What is the PIC model?

Is it an efficient way of modeling the Vlasov equation? No, it is a Klimontovich description for finite size (macro-particles)

Mathematical model for PIC

Klimontovich equation of macro-particles

Maxwell's equations

$$\begin{split} &\frac{D}{Dt}F = 0\\ F(\vec{x}, \vec{v}; t) = \sum_{i}^{N} S_{p}(\vec{x} - \vec{x}_{i}(t))\delta(\vec{v} - \vec{v}_{i}(t))\\ &\left[\frac{D}{Dt} \equiv \partial_{t} + \vec{v} \cdot \nabla_{x} + \vec{a} \cdot \nabla_{v}\right]\\ &\vec{a} \equiv \frac{d}{dt}\vec{v} = \frac{q}{m}\left(\vec{E} + \frac{\vec{v}}{c} \times \vec{B}\right) \end{split}$$

$$\begin{split} \vec{J}(\vec{x},t) &= \int d\vec{v} \; q\vec{v} \; F(\vec{x},\vec{v},t) \\ & \frac{\partial}{\partial t} \vec{B} = -\nabla \times \vec{E} \\ & \frac{\partial}{\partial t} \vec{E} = \nabla \times \vec{B} - \frac{4\pi}{c} \vec{J} \end{split}$$

## OSIRIS 4.0 (began in late 1990s from LLNL funding)

Massivelly Parallel, Fully Relativistic

Particle-in-Cell (PIC) Code

Visualization and Data Analysis

osiris framework

Infrastructure

Developed by the

osiris.consortium

 $\Rightarrow$  UCLA + IST





#### Ricardo Fonseca:

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epp.tecnico.ulisboa.pt/ http://picks.idre.ucla.edu/

accessible through MoU

# Speedup on Sequoia

#### 1,000

#### code features

• Scalability to ~ 1.6 M cores

100,000

# Cores

10,000,000

- SIMD hardware optimized
- Parallel I/O
- Dynamic Load Balancing
- Collisions
- Field ionization
- · QED module
- Particle splitting/merging
- · Quasi-3D
- Boosted frame
- GPGPU support
- Xeon Phi support

#### UCLA QuickPIC: A PIC code based on quasistatic approximation

$$\vec{E}_{\perp} + \hat{z} \times \vec{B}_{\perp} = -\nabla_{\perp} \cdot \psi$$

$$\nabla_{\perp}^{2} \psi = -(\rho - J_{z})$$

$$\nabla_{\perp}^{2} \vec{B}_{\perp} = \hat{z} \times (\frac{\partial}{\partial \xi} \vec{J}_{\perp} + \nabla_{\perp} \cdot J_{z})$$

$$\nabla_{\perp}^{2} B_{z} = -\nabla_{\perp} \times \vec{J}_{\perp}$$

$$\nabla_{\perp}^{2} E_{z} = \nabla_{\perp} \cdot \vec{J}_{\perp}$$

$$plasma: \frac{d\vec{p}}{d\xi} = \frac{q/m}{1 - v_{z}} \left[\vec{E} + \vec{v} \times \vec{B}\right]$$

$$\frac{\partial}{\partial \xi} (\rho - J_{z}) + \nabla_{\perp} \cdot \vec{J}_{\perp} = 0$$

$$\frac{\partial}{\partial \xi} \int (\rho - J_{z}) d\vec{x}_{\perp} + \int \nabla_{\perp} \cdot \vec{J}_{\perp} d\vec{x}_{\perp} = 0$$
\* P. Mora and T. Antonsen, Phys. Plasmas 4, 217 (1996)



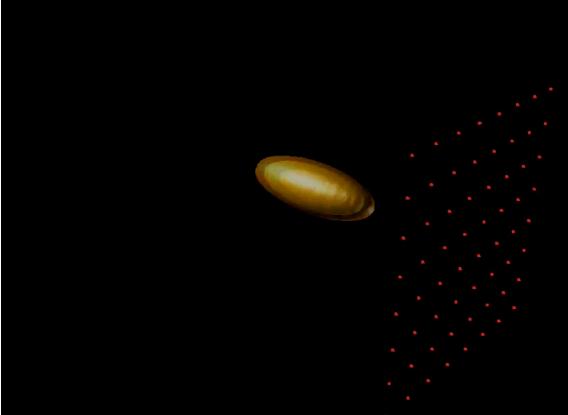
#### QuickPIC: 3D Quasi-static PIC Opensource

Fully parallelized and scaled to 100,000+ cores

Requires predictor corrector, has some similarities with a Darwin code.

C-K. Huang et al., 2006 W. An et al., 2014

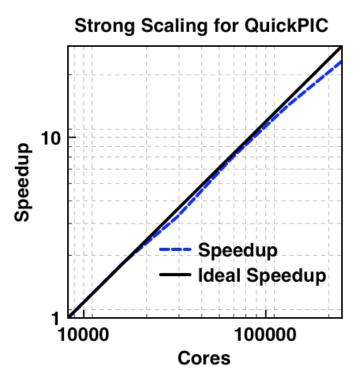
**Recently HIPACE** 



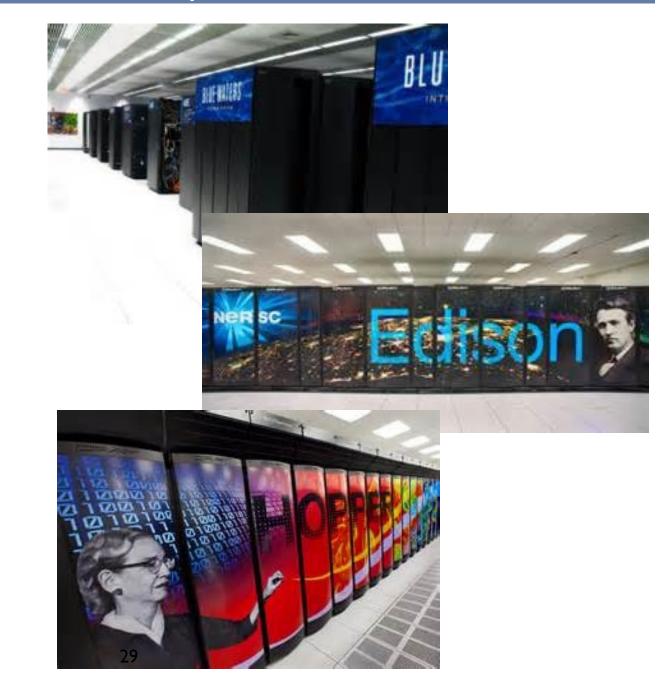
Embeds a parallelized 2D PIC code inside a 3D PIC code based on UPIC Framework.



#### Current Status of QuickPIC Opensource



Time for pushing one particle for one step u s i n g a s i n g l e processor (double precision): ~770 ns





## Example of synergy at FFTB: 42 GeV energy gain in less than I meter!

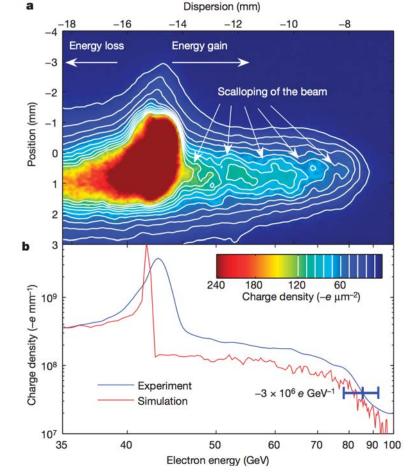


#### Can high gradients be sustained over 1 meter?

Blumenfeld et al., 2007

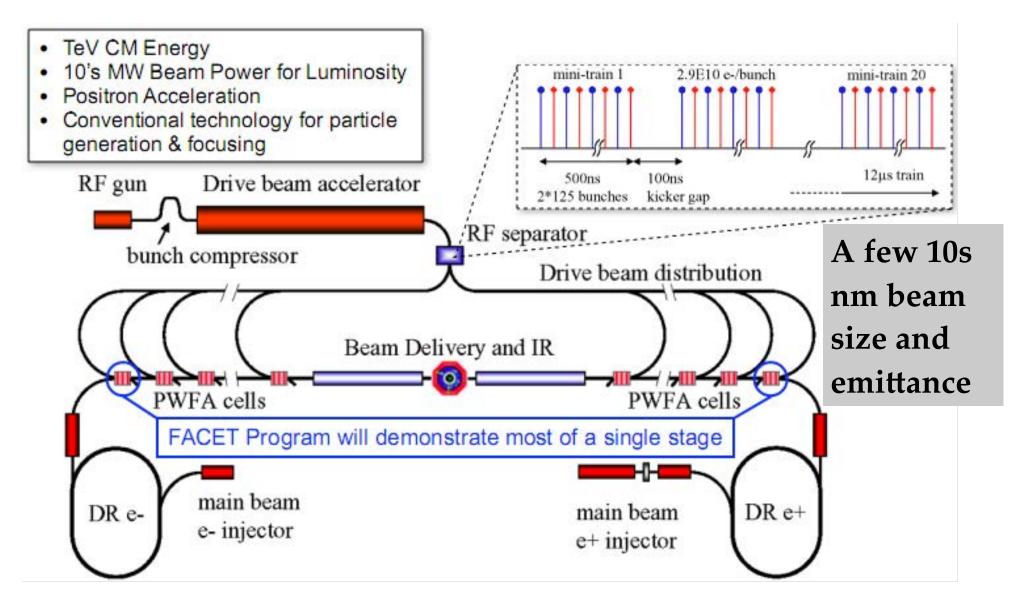
43 GeV electron beam 1.8 x 10^10 particles ~25kA 10^17cm-3 plasma

QuickPIC simulations explained what limited the acceleration length





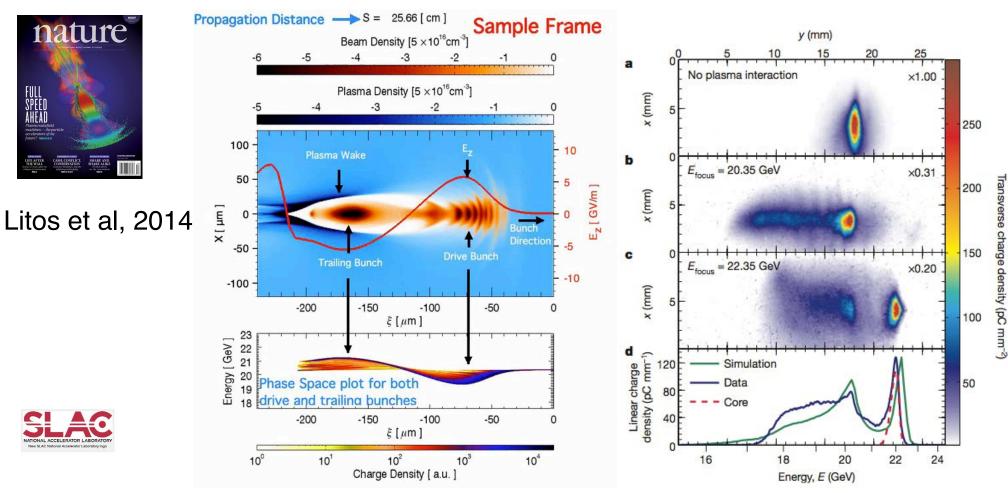
#### PWFA-based collider concept (no ILC)



a 19 Stages PWFA-LC with 25GeV energy gain per stage

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#### Another example at FACET: Synergy demonstrated efficient beam loading of wake



QuickPIC and OSIRIS simulations helped to design and interpret the experiment.

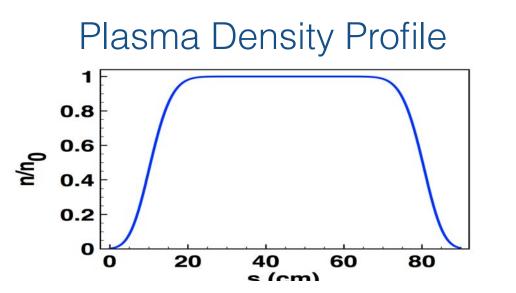
# UCLA FACET II (proposed) with QuickPIC

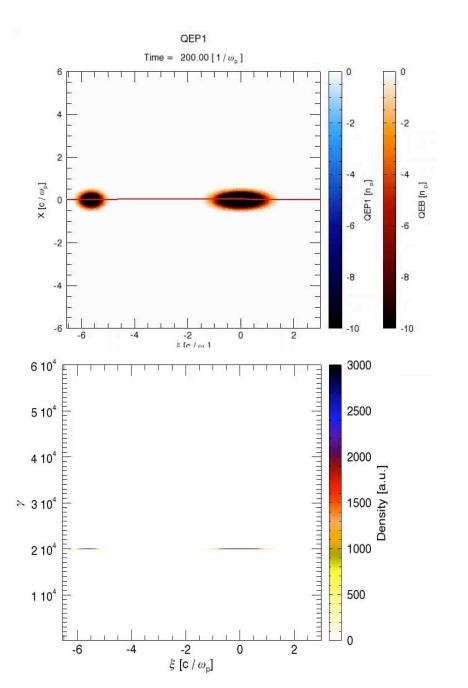
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Drive Beam: E = 10 GeV, I<sub>peak</sub>=15 kA  $\sigma_r$  = 21.17 μm,  $\sigma_z$  = 12.77 μm , N =1.0 x 10<sup>10</sup> (1.6 nC),  $\epsilon_N$  = 10 μm

Trailing Beam: E = 10 GeV, I<sub>peak</sub>=9 kA  $\sigma_r$  = 21.17 μm,  $\sigma_z$  = 6.38 μm , N =0.3 x 10<sup>10</sup> (0.48 nC),  $\epsilon_N$  = 10 μm

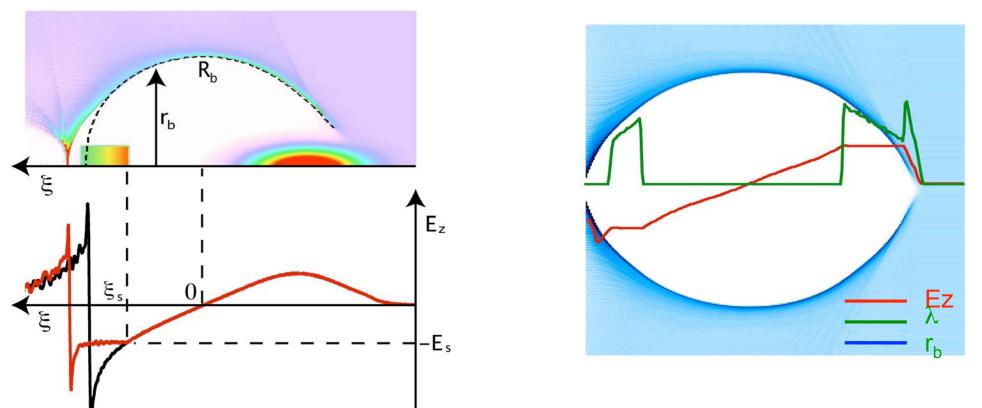
Distance between two bunches: 150 μm Plasma Density: 4.0 x 10<sup>16</sup> cm<sup>-3</sup> (with ramps)





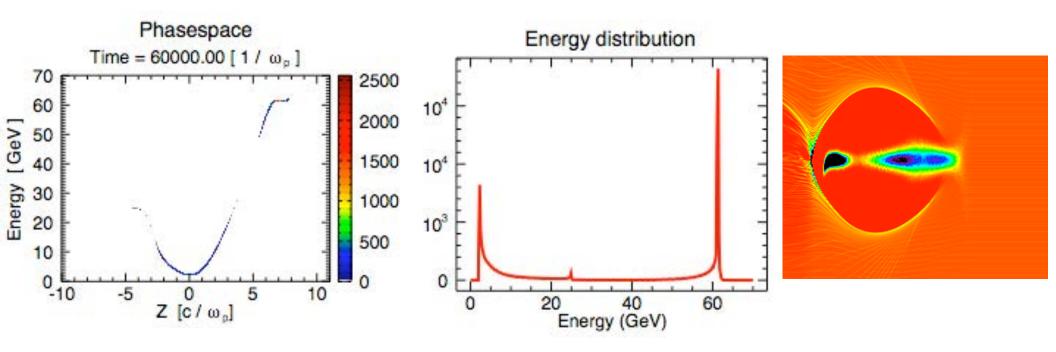
Grand challenge research problem to develop self-consistent beam loading scenarios: Stable high gradient acceleration while maintaining beam quality

#### Nonlinear beam loading and shaping bunches M. Tzoufras et al. PRL 2008



- Theory allows for designing highly efficient stages that maintain excellent beam quality.
- Theory allows for understanding how standard beams absorb energy of nonlinear wakes.
- This regime cannot work for positrons.

# Model a single stage of a PWFA-LC design including beam loading



Drive beam energy is 25 GeV Output bean energy is ~60GeV

1% Energy spread Efficiency from drive to trailing bunch ~48%!

Trailing beam is very tightly focused. Electric field in trailing beam ~10 TeV/cm.

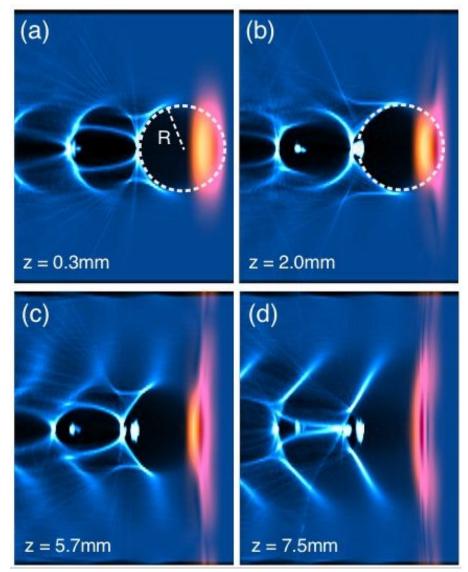
### What about laser drivers?

## Nonlinear self-guided blowout regime for LWFA

- The ponderomotive force of the laser pushes the electrons out of the way.
- The ion channel supports huge and ideal accelerating and focusing fields.
- Electrons are self or externally injected at the tail of the ion channel.
- Beam loading flattens wake.
- → The laser's spotsize is "matched":
- Local pump depletion: The front of the laser etches back:

$$k_p w_0 \simeq k_p R_b \simeq 2\sqrt{a_0}$$
 $\upsilon_{etch} \simeq c \frac{n_p}{n_c}$ 

Lu et al. 2007

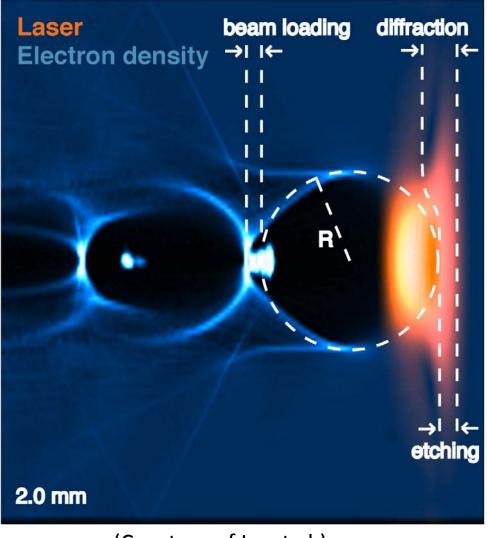


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# The Laser Wakefield Accelerator



### Phenomena that are Relevant to the Study of LWFAs



<sup>(</sup>Courtesy of Lu et al.)

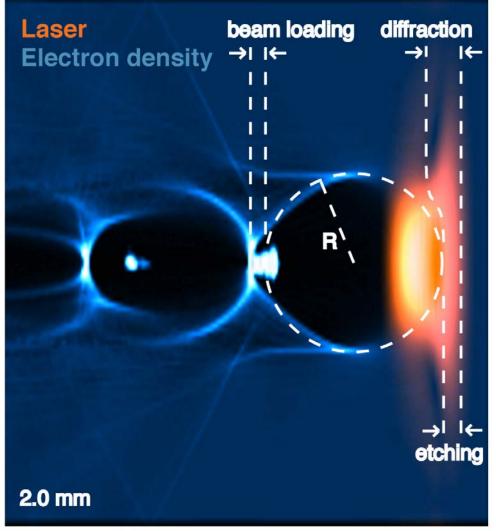
- The front of the laser pulse loses energy to the plasma and etches back (pump depletion)
- Electrons are self or externally injected in the back of the ion channel, slightly distorting the wake in their region (beam loading)
- The front of the laser, once depleted of most of it's energy, diffracts



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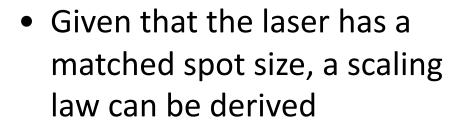
# The Laser Wakefield Accelerator

### Phenomenological Scaling Law<sup>+</sup>



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(Courtesy of Lu et al.)



• The maximum accelerating distance is estimated as

$$L_d \simeq \frac{2}{3} \left(\frac{\omega_0}{\omega_p}\right)^2 W_0$$

 The particle energy estimated as

$$\Delta E = \frac{2}{3}mc^2 \left(\frac{\omega_0}{\omega_p}\right)^2 a_0$$



ſ

$$\Delta E[\text{GeV}] \simeq 1.7 \left(\frac{P[\text{TW}]}{100}\right)^{1/3} \left(\frac{10^{18}}{n_p[\text{cm}^{-3}]}\right)^{2/3} \left(\frac{0.8}{\lambda_0[\mu\text{m}]}\right)^{4/3}$$

• How far does it scale?

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**UCLA** Useful to rewrite in terms of the laser energy



### We Explore the Implications of the Scaling Laws Given a Fixed Energy Laser

Optimal density and pulse length for a laser of fixed energy:

$$E_L = \alpha P \tau,$$

Assuming a matched spot size, we can adjust the relative pulse length as a free parameter:

$$\tau = \mathcal{F}2\sqrt{a_0}\omega_p^{-1}$$

We may recast the scaling laws equations as a function of the laser energy, pulse length, and amplitude:

$$\Delta E = \frac{2}{3} \frac{m_e c^2}{\alpha^{2/3}} \left[ \frac{4\omega_0}{\mathcal{A}} \right]^{2/3} \frac{E_L^{2/3}}{\mathcal{F}^{2/3} a_0^{4/3}}.$$

[*A* = 17 GW]

Logically, there is a lower bound to the pulse length that can be determined empirically

A. Davidson, PhD Disseration UCLA 2016

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# Pushing the Theory Further

### by Scaling to Higher Acceleration Distances

estimated CPU hours (3D)	P (TW)	n <sub>p</sub> (cm⁻³)	W <sub>0</sub> (μm)	L <sub>d</sub> (cm)	a <sub>o</sub>	∆E (GeV) Estimated	∆E (GeV) Simulated
100,000	200	1.5 x 10 <sup>18</sup>	19.5	1.5†	4.0	1.58	1.55†
430,000	324	1.0 x 10 <sup>18</sup>	22.0	2.62	4.44	2.52	???
3,200,000	649	5.0 x 10 <sup>17</sup>	31.7	7.37	4.44	5.28	???
26,000,000	1298	2.5 x 10 <sup>17</sup>	44.8	20.8	4.44	10.57	???
120,000,000	2162	1.5 x 10 <sup>17</sup>	57.8	44.8	4.44	17.6	???
340,000,000	3280	1.0 x 10 <sup>17</sup>	71.2	83.8	4.44	26.7	???

#### We implement the quasi-3D geometry to attain hundreds of times of speedup

<sup>+</sup>Lu et al. Conducted this simulation over 0.75cm, not the entire  $L_d$ 

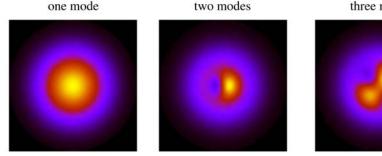
### New reduced models: 3D simulations of LWFA and PWFA (e and p) can be expensive, but quasi-3D simulations are now possible!

- 2D cylindrical r-z simulations can get the geometric scaling correct: Used extensively for PWFA
- Laser pulses are radially polarized in r-z simulations, so cylindrical r-z simulations not used for LWFA studies.
- In many 3D simulations the drivers and wake develop only lower order azimuthal modes.
- Expand in azimuthal mode number and truncate expansion! [1]: This is PIC in r-z and gridless in phi [2]
- We have now incorporated the ability to expand the fields into an arbitrary number of azimuthal modes into OSIRIS. Made improvements to [1] including rigorous charge conserving algorithm [2]. As part of OSIRIS, algorithm scales to 1,000,000+ cores and can model laser, beams, and beam loading. Allows rapid parameter scans.

[1] A.F. Lifshitz et al., JCP 228, pp.1803 (2009).

[2] A. Davidson et al., JCP 281 pp.1063 (2015).

Can reduce simulation time by factors of 100s. For example from 10,000,000s of core hours to 80,000 core hours!

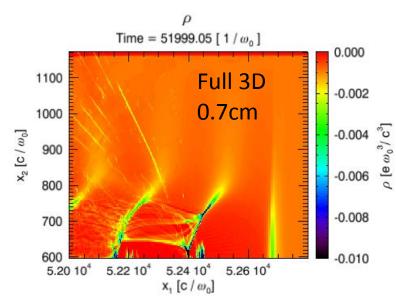


three modes

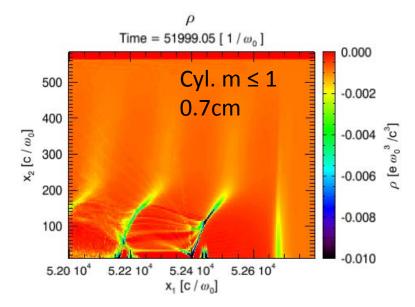
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### Excellent agreement between 3D & quasi-3D OSIRIS for original Lu et al.



 This geometry has been tested against and shown to "reproduce" known 3D Cartesian simulations<sup>†</sup>.



Geometry	Charge [pC] (1cm)	ex [π mm mrad]	ey [π mm mrad]	Max Energy [GeV]
Full 3D Cartesian	340	27	30	1.57
Cyl. Mode ≤ 1	328	43	43	1.55

[†] W. Lu, M. Tzoufras, C. Joshi, F. Tsung, W. Mori, J. Vieira, R. Fonseca, L. Silva, Generating multi-gev electron bunches using single stage laser wakefield acceleration in a 3d nonlinear regime, Physical Review Special Topics - Accelerators and Beams 10 (061301).doi:http://link.aps.org/doi/10.1103/PhysRevSTAB.10.061301.

# **UCLA** Scaling Laws in Nonlinear Regime

### Faster Methods Means Physics in Farther Regimes

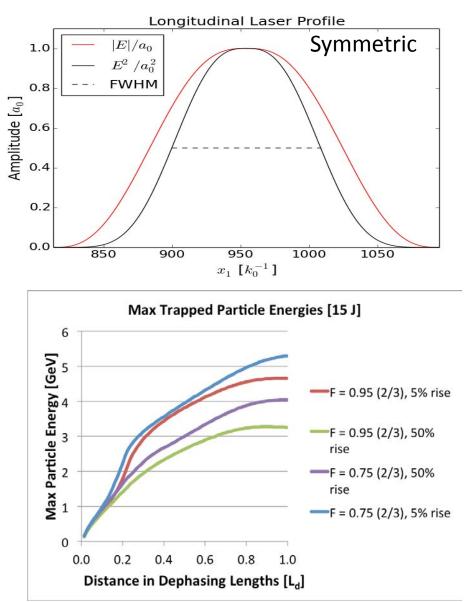
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Laser Energy (J)	P (TW)	n <sub>p</sub> (cm⁻³)	W <sub>0</sub> (μm)	L <sub>d</sub> (cm)	a <sub>0</sub>	ΔE (GeV) Estimated	ΔE (GeV) Simulated
6	200	1.5 x 10 <sup>18</sup>	19.5	1.5†	4.0	1.58	1.55†
16	324	1.0 x 10 <sup>18</sup>	22.0	2.62	4.44	2.52	3.46
47	649	5.0 x 10 <sup>17</sup>	31.7	7.37	4.44	5.28	6.63
133	1298	2.5 x 10 <sup>17</sup>	44.8	20.8	4.44	10.57	13.6
290	2162	1.5 x 10 <sup>17</sup>	57.8	44.8	4.44	17.6	???
542	3280	1.0 x 10 <sup>17</sup>	71.2	83.8	4.44	26.7	???

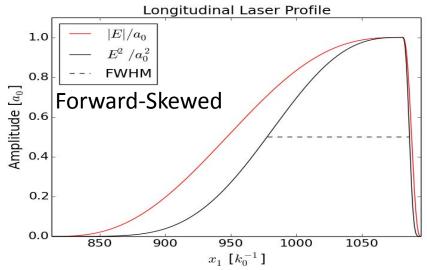


# Capabilities of a Fixed Energy Laser

### Longitudinal Profile Adjustment



UCLA



- Best results I found so far, by combining a skewed profile with optimized pulse length:
  - 15 J Laser : **5.3 GeV**
  - 30 J Laser : 8.1 GeV
- Note that for the 15 J case, this is twice the estimated energy of 2.52 GeV using default parameters

Grand challenge research problem is generation of ultra-low emittance beams and the manipulation of six dimensional phase space

# Easier to inject electrons into wake when they are "born" inside wake.

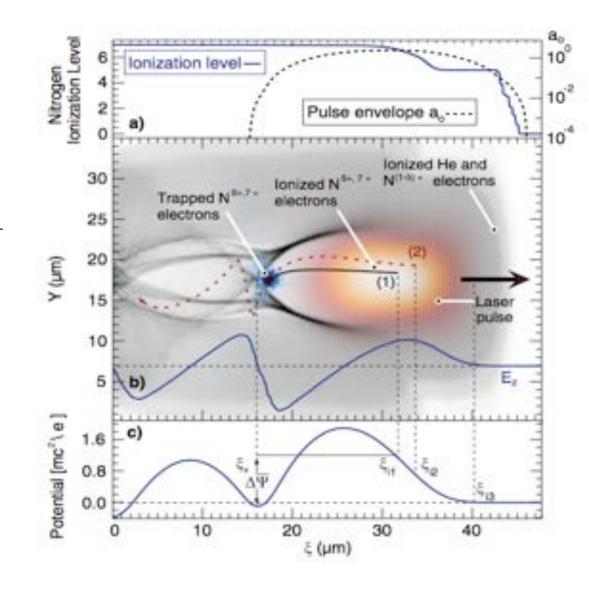


Easier to satisfy trapping condition:

$$\Delta\psi\equiv\psi_{final}-\psi_{init}<-1$$

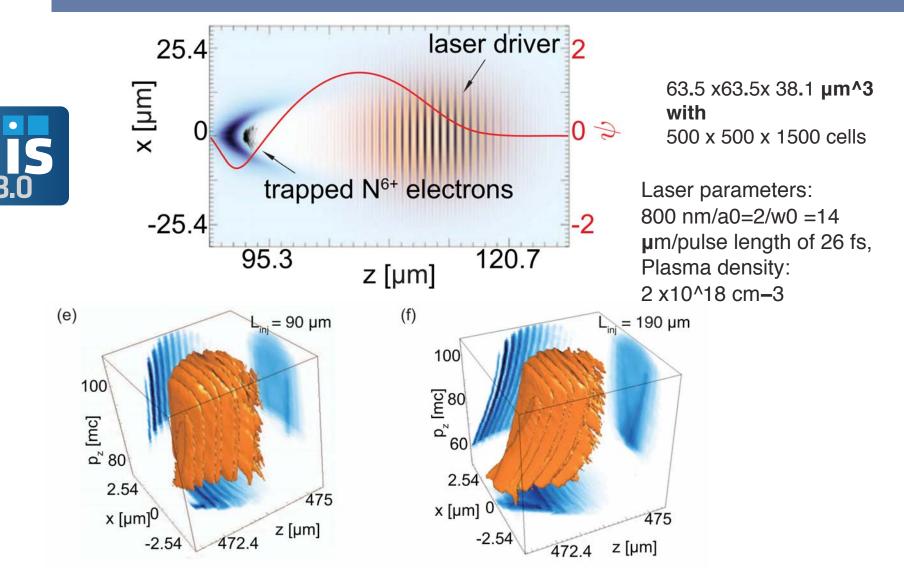
Pak et al. PRL 2009

Many recent papers on using ionization injection





# lonization injection can generate nano-bunched electron beams



X. Xu, et al. Physical Review Letters 2016

#### Ultrabright e<sup>-</sup> bunch generation using down ramp UCLA injection: PWFA or LWFA driven

X. Xu, F. Li, et al. UCLA/Tsinghua submitted 2016

- Could drive a compact XFEL
- Beam-driven plasma acceleration • in blowout regime can produce such beam via density transition: ~1nC

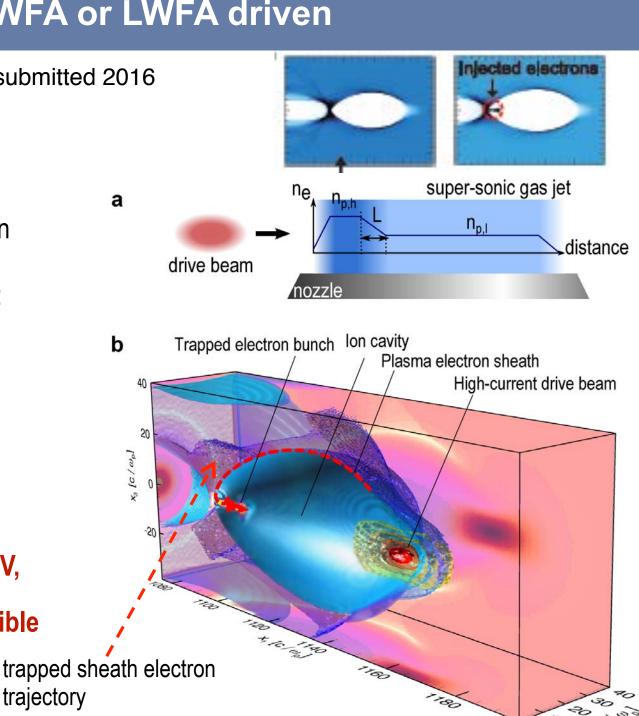


**Output Beam** 

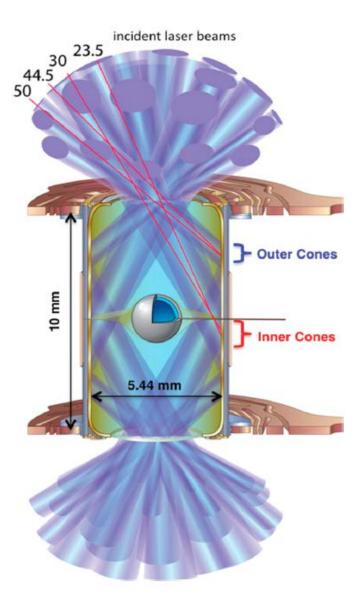
 $\epsilon_n$ <50nm rad, I~10 kA,  $\Delta$ E<3 MeV,

 $B_n > 10^{21}A m^{-2} rad^{-2} may be possible$ 

trajectory



### NIF this is incredibly complicated with much fundamental science



1.8 MJ into holhraum but just ~10s KJ into compressed fuel

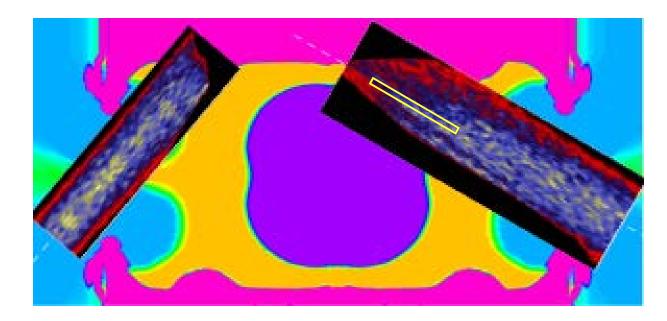
Requires very symmetric compression Requires very symmetric x-ray drive Requires correct time dependence to this drive

Lasers must hit where they are aimed!

Hinkel et al., PoP (2011)

## NONLINEAR OPTICS OF PLASMAS

But lasers propagate through long regions of tenuous on NIF. They can "scatter" into other light waves and electron plasma and ion acoustic waves. This is **very** complicated. Beams cross paths and beams are broken up into "speckles" (Gaussian beamlets) to minimize the instabilities.



Colors correspond to different ranges of density (and material).

Inner Boxes: pF3D simulations of speckled laser beams Inner yellow box is the size that a fully kinetic simulation can model. After 40 years of research, and the evolution of software and computers, it is now possible to carry out full kinetic simulations and data analysis of a meaningful volume and time duration of a NIF (laser fusion) beam propagating through a plasma.

Example: Stimulated Raman Scattering (no magnetic fields) VPIC and OSIRIS

# Multi-speckle SRS

900 µm

2D plasma simulated for 16 ps:

T<sub>e</sub> = 2.75 keV

linear density gradient,  $n_e/n_{cr} = 0.105$  to 0.135

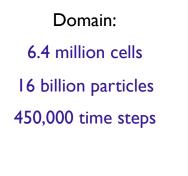
0.14

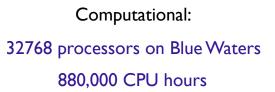
0.12

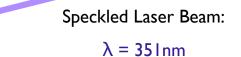
0.10

 $n_{\rm e}/n_{\rm cr}$ 

20 µm



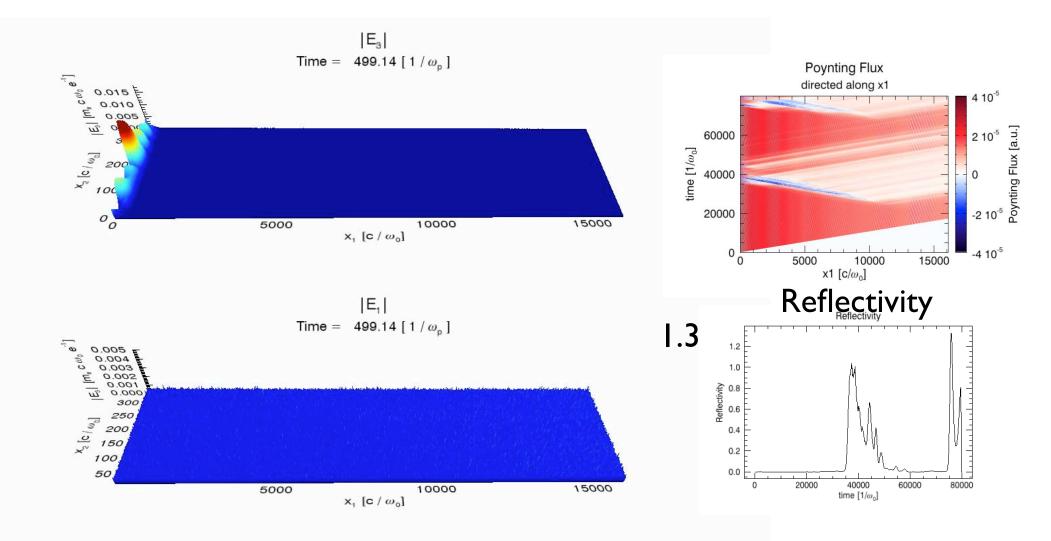




 $I_{avg} = 10^{15} W/cm^2$ 

5 speckles long x 7 speckles wide

### Multi-speckle (~35) fully kinetic simulations: Reflectivity is bursty: ~1,000,000 CPU Hours





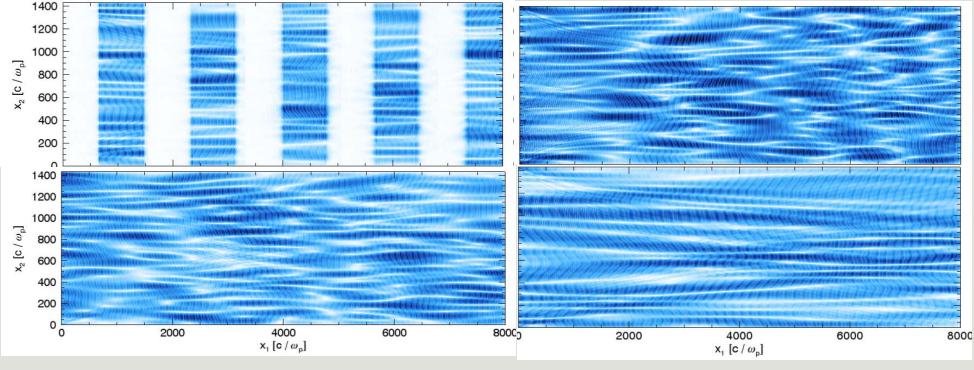




# Examples of speckle patterns generated by OSIRIS with different smoothing techniques

STUD

Multi-FM SSD



ISI

RPP





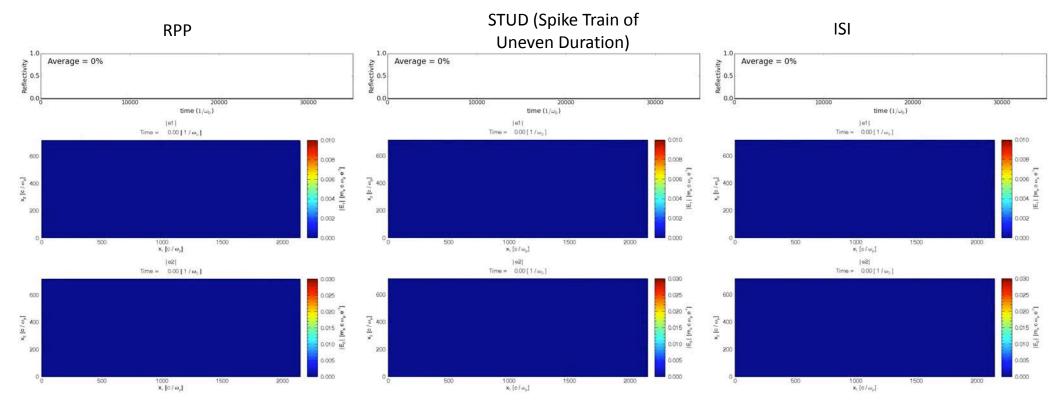
### Simulation parameters

Particle-per-cell: 256 Grid: 44750×4772, 10740×2386 Electron temperature: 3.0 keV Intensity:  $5 \sim 10 \times 10^{14} W/cm^2$   $256^{1}$ 36864 × 4096 2.6 keV 2~5×10<sup>14</sup>W/cm<sup>2</sup>

# LPI Simulation Results — Temporal bandwidth can reduce LPI

### Temporal bandwidth can reduce SRS growth

- Small simulations (90k core-hours each) to identify interesting parameters before starting full simulations (<1 million core-hours each)</li>
  - 15 speckles across and ~120 microns long, I=10^15
  - ~100 million grids and ~10 billion particles each.
- Incorporating polarization smoothing can further reduce SRS reflectivity
- These are very preliminary: Wide parameter space needs to be studied.





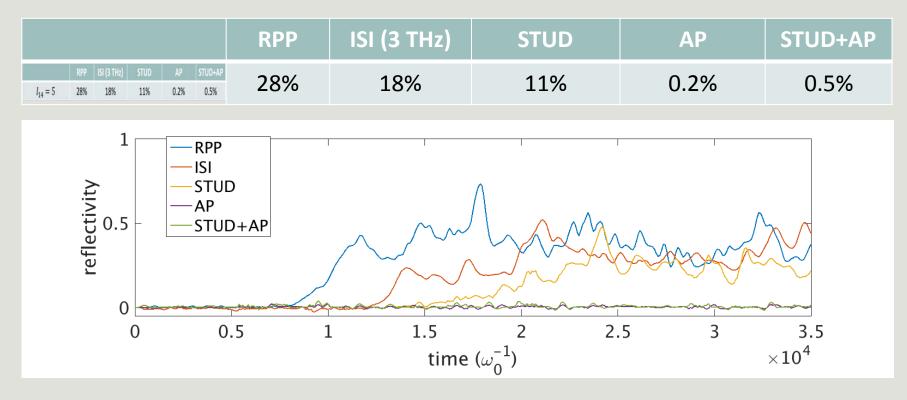




# long-scale-length simulations with $I_{14} = 5 (L_{INT} \gg L_{HS})$

 $n \in [0.125, 0.135] n_c, T_e = 3~keV$ , box size  $500 \times 80 \mu m, L_{HS} = 90 \mu m, L_{spike} = 40 \mu m$ 

•  $I_{14} = 5, L_{INT} = 450 \mu m, G = 56$ 



Exascale may permit 3D PIC simulations of LPI. Will require code development including new numerical methods and algorithms for new hardware

	2D multi-speckle along NIF beam path	3D, 1 speckles	3D, multi-speckle along NIF beam path
Speckle scale	50 x 8	1 x 1 x 1	20 x 20 x 5
Size (microns)	150 x 1500	9 x 9 x 120	56 x 56 x 900
Grids	9,000 x 134,000	500 x 500 x 11,000	1,700 x 1,700 x 80,000
Particles	300 billion (256/cell)	300 billion (64/cell)	10 trillion (64/cell)
Steps	470,000 (15 ps)	540,000 (15 ps)	540,000 (15 ps)
Memory Usage*	1.5 TB	1.5 TB	1 PB
core-hours	8 million	13 million	1 billion (2 months on Blue Waters; Exascale

Estimates are sensitive to resolution and number of particles

# The UCLA Particle-in-Cell and Kinetic Simulation SoftwareCenter (PICKSC)



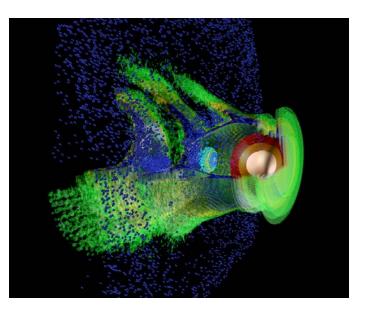
The mission of the Particle-in-Cell and Kinetic Simulation Software Center (PICKSC) at UCLA is to support an international community of PIC and plasma kinetic software users, developers, and educators, and to increase the use of this software for accelerating the rate of scientific discovery.

OSIRIS, QuickPIC, UPIC, OSHUN

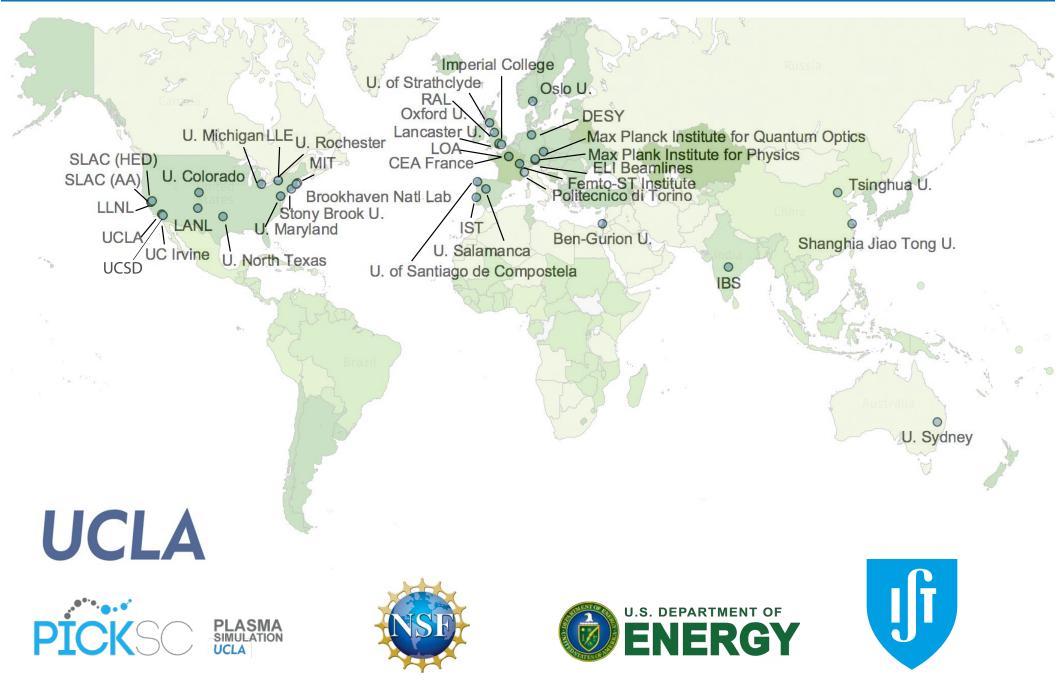
### http://picksc.idre.ucla.edu

E-mail me if you would like more information about available software: Mori@physics.ucla.edu

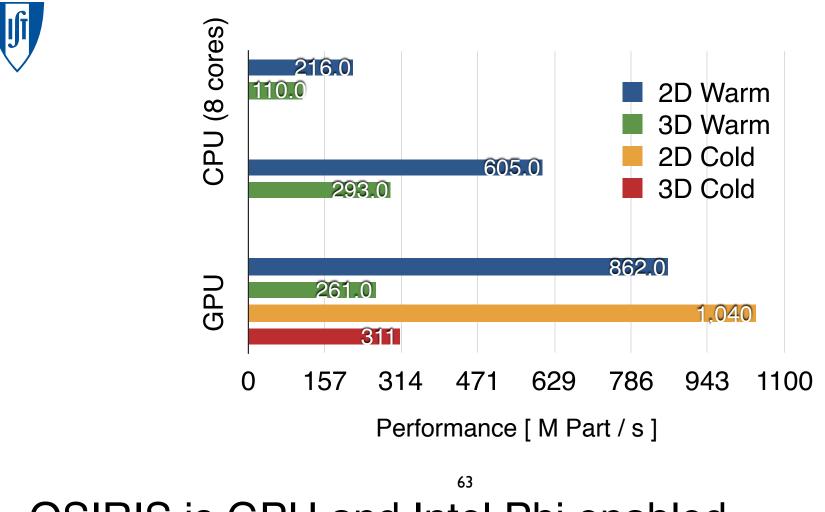




# OSIRIS and QuickPIC access and use is international: Used in AA and HEDP research



# UCLA Take advantage of many core



**OSIRIS** is GPU and Intel Phi enabled

HEDLP is rich in big and discovery driven science.

There is a close synergy between experiment and simulation.

Kinetic software continues to advance forward.

Preparing for exascale requires learning how to run on petascale.

Plasma based acceleration is making rapid progress.

A roadmap for research for a linear collider application is underway. The modeling capability and concepts for a paper study could be available within the next decade.

A compact XFEL is a realistic goal within the next decade.

Ignition on NIF will occur within the next ?????