

Building a Wet Planet from Dry Materials

S.-H. Dan Shim (심상헌)

**School of Earth and Space Exploration
Arizona State University**

Shim Lab

Laser-heated diamond-anvil cell

- H/H₂O-Silicate/Metal reaction
- Pulsed laser heating with gated synchrotron XRD

Multi-anvil press

- Synthesis of dense shock targets
- Chemical interactions in multi-phase systems

Laser-driven shock

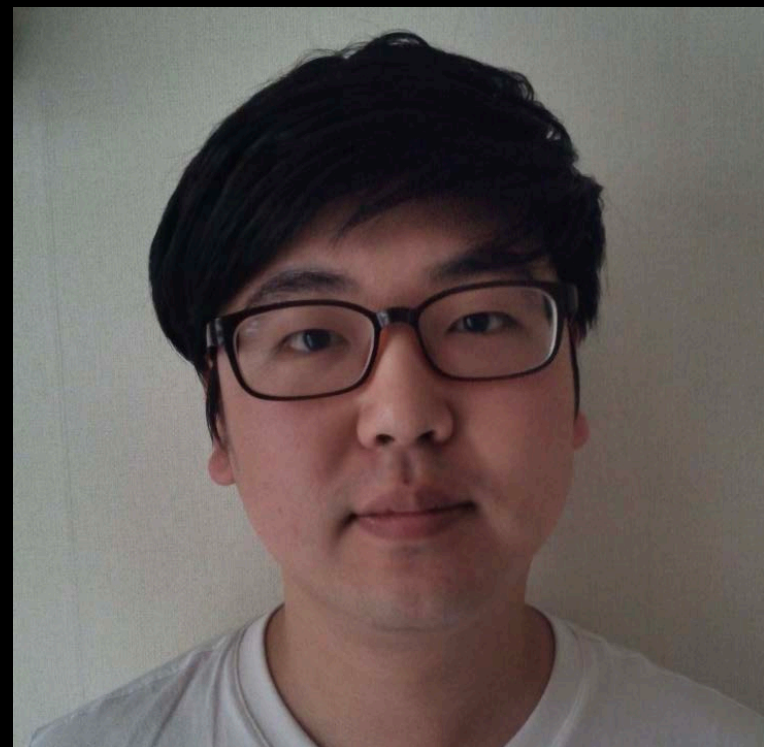
- Atomic-scale structure of planetary melts
- X-ray diffraction and X-ray spectroscopy



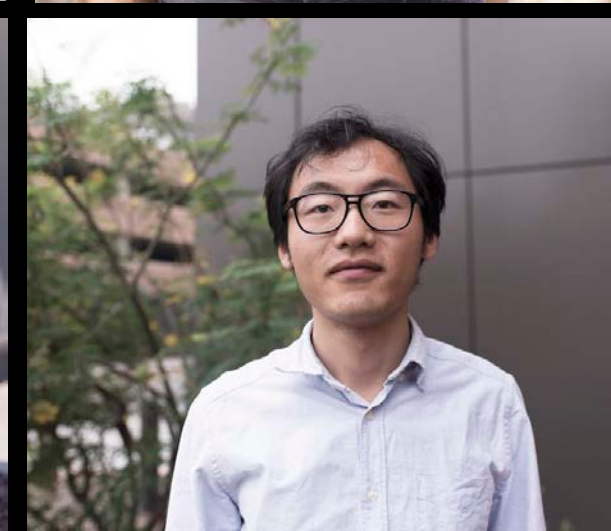
Synchrotron



X-ray free electron laser



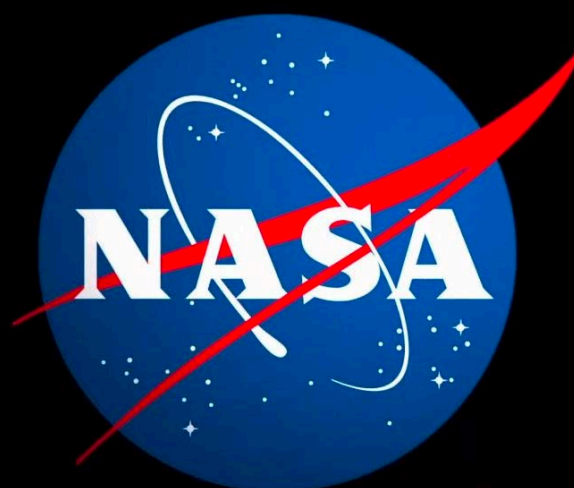
Shim Lab @ASU



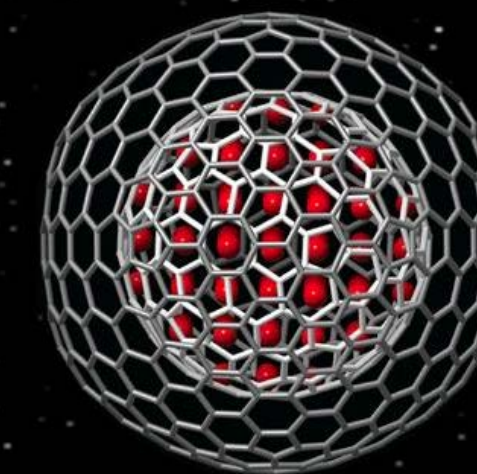
K. Leinenweber, A. Chizmeshya, S. Speziale, S. Chariton, V. Prakapenka, A. Vazan, A. Anbar, S. Desch, M. Li, E. Garnero, P. Buseck



EAR
AST
CSEDI
FESD

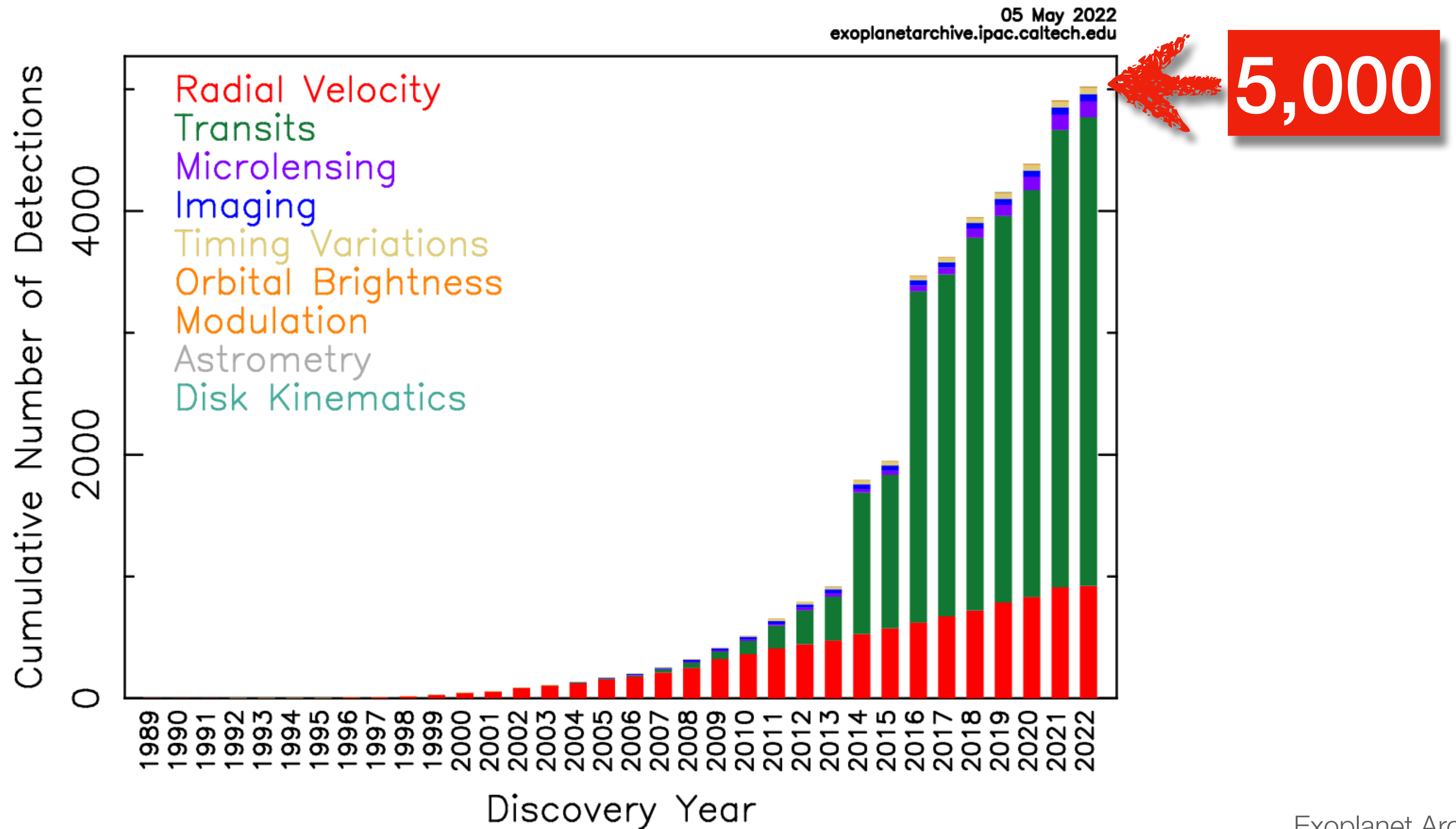


Exoplanet
NExSS
Astrobiology

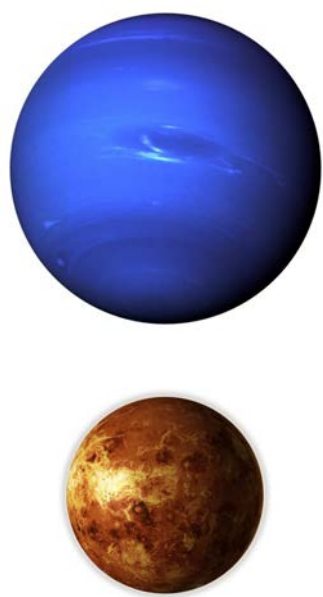


Materials
of the
Universe
@ASU

Exoplanet Discoveries

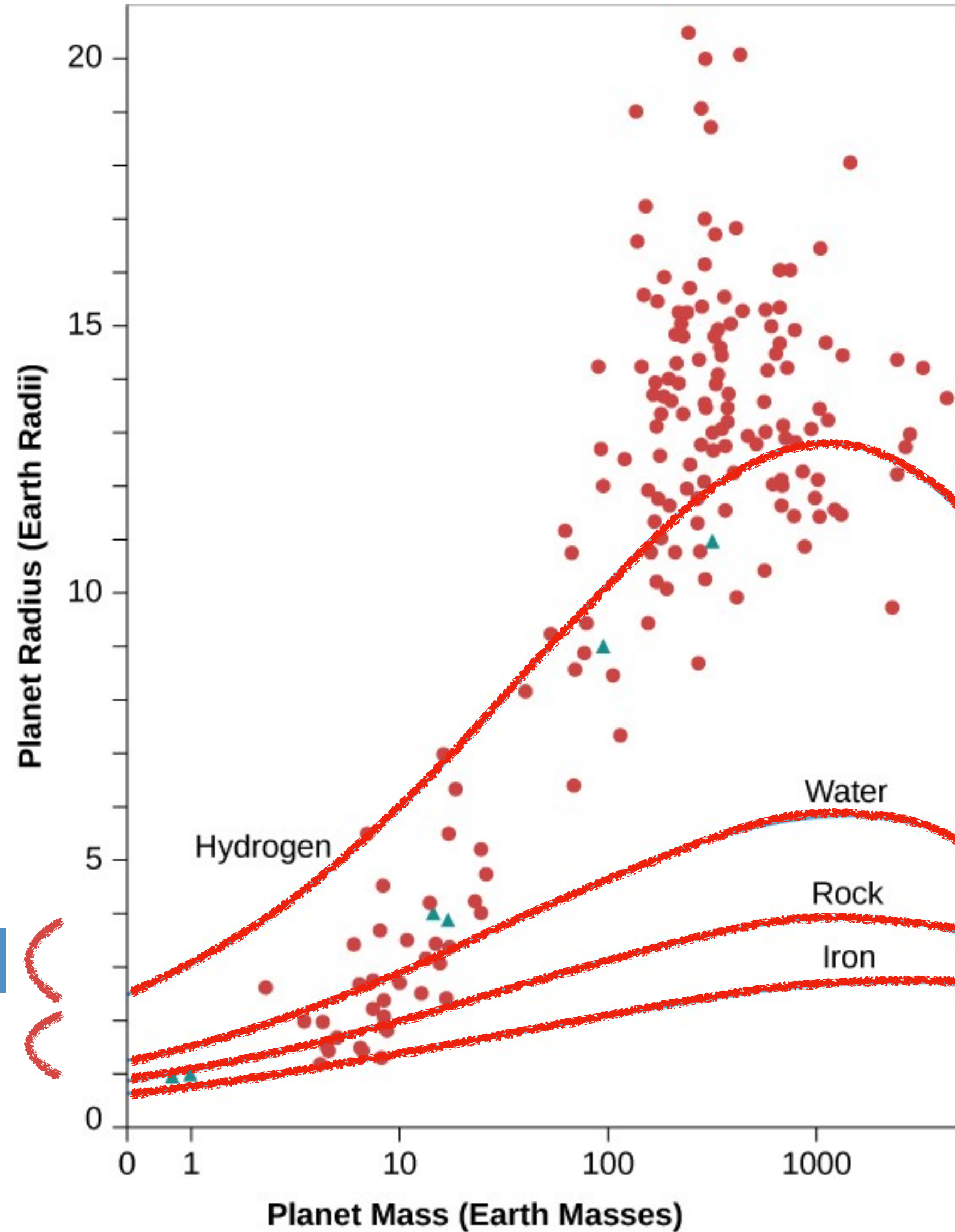


Mass-Radius



Sub-Neptunes

Super-Earths



Gas Giants



▲ Planets in our solar system

● Exoplanets with known masses and radii

Nanosecond X-ray diffraction of shock-compressed superionic water ice

Marius Millot^{1,3*}, Federica Coppari^{1,3*}, J. Ryan Rygg^{1,2}, Antonio Correa Barrios¹, Sebastien Hamel¹, Damian C. Swift¹ & Jon H. Eggert¹

High-pressure physics plays a vital role for understanding exoplanets

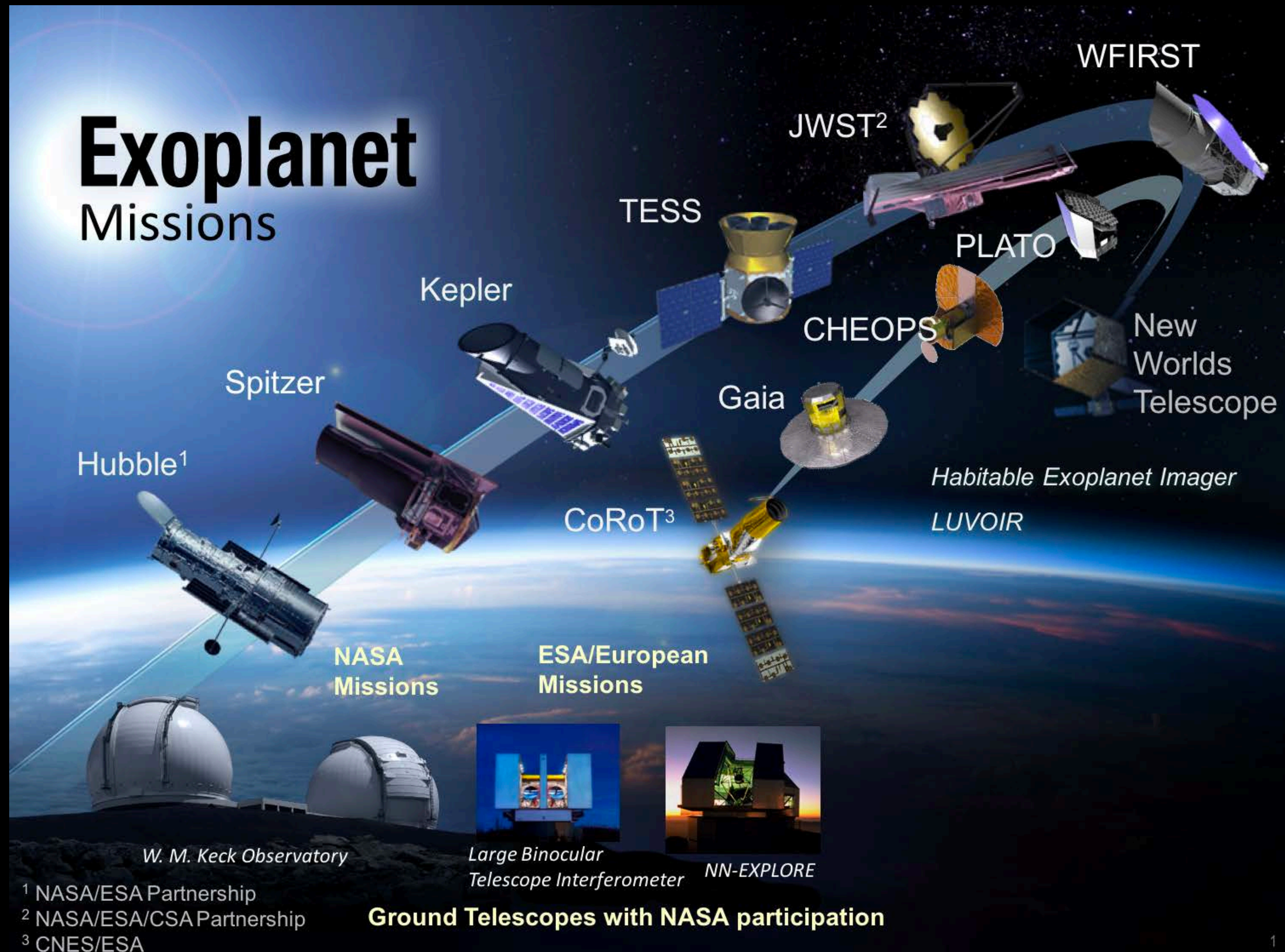
Kraus *et al.*, *Science* **375**, 202–205 (2022) 14 January 2022

PLANETARY SCIENCE

Measuring the melting curve of iron at super-Earth core conditions

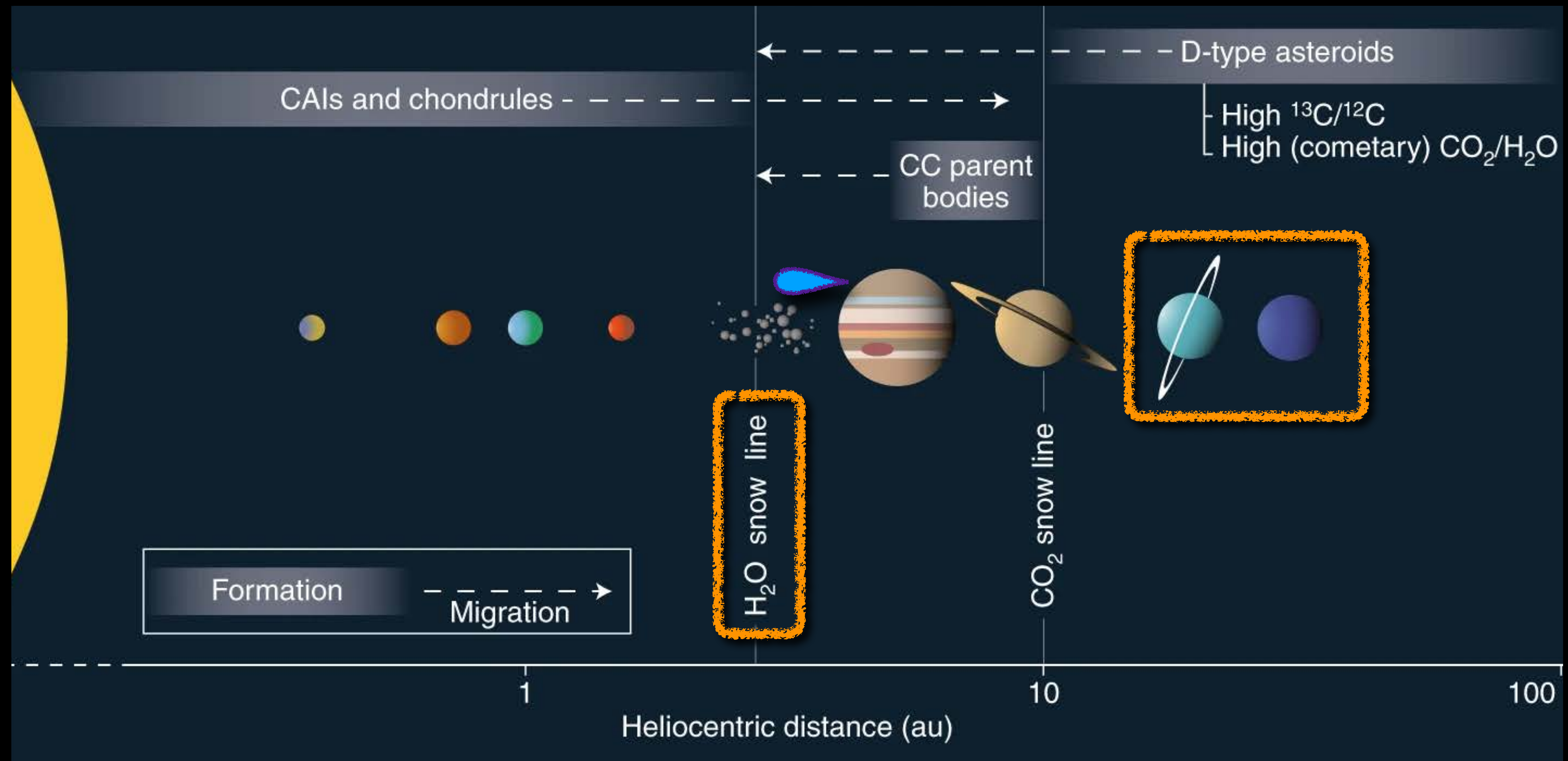
Richard G. Kraus^{1*}, Russell J. Hemley², Suzanne J. Ali¹, Jonathan L. Belof¹, Lorin X. Benedict¹, Joel Bernier¹, Dave Braun¹, R. E. Cohen³, Gilbert W. Collins⁴, Federica Coppari¹, Michael P. Desjarlais⁵, Dayne Fratanduono¹, Sebastien Hamel¹, Andy Krygier¹, Amy Lazicki¹, James Mcnaney¹, Marius Millot¹, Philip C. Myint¹, Matthew G. Newman⁶, James R. Rygg⁴, Dane M. Sterbentz¹, Sarah T. Stewart⁷, Lars Stixrude⁸, Damian C. Swift¹, Chris Wehrenberg¹, Jon H. Eggert¹

Atmosphere-Interior Interactions

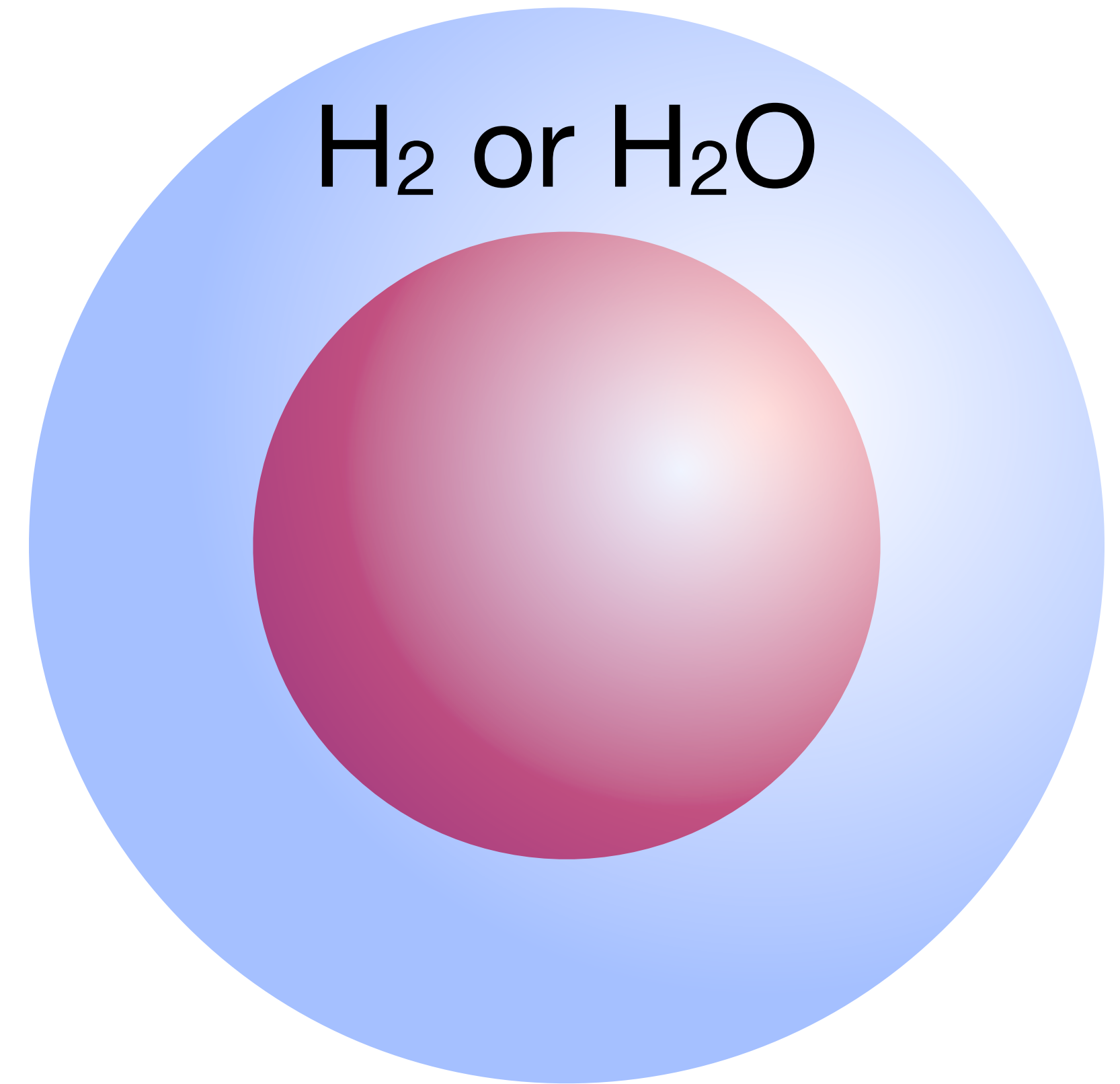
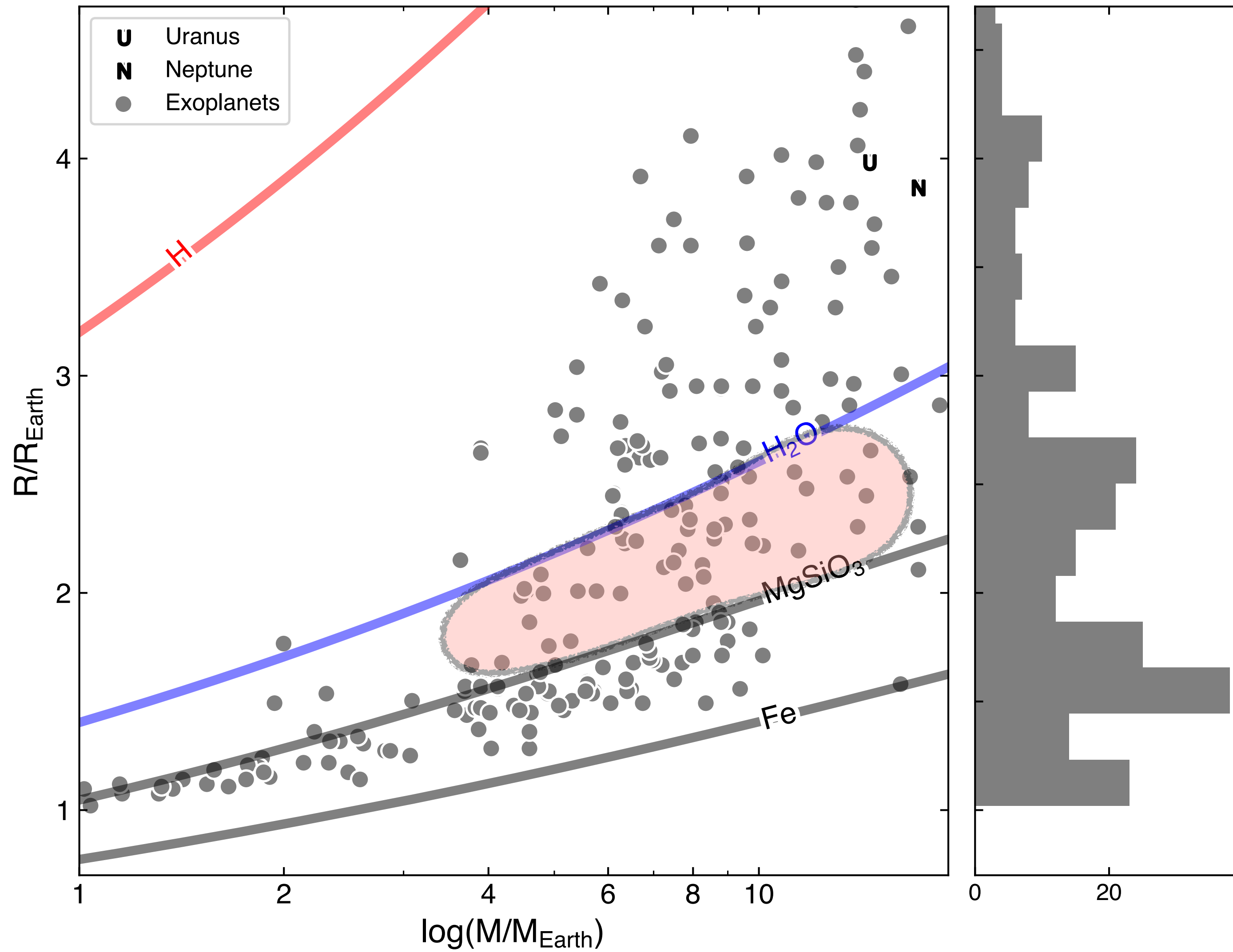




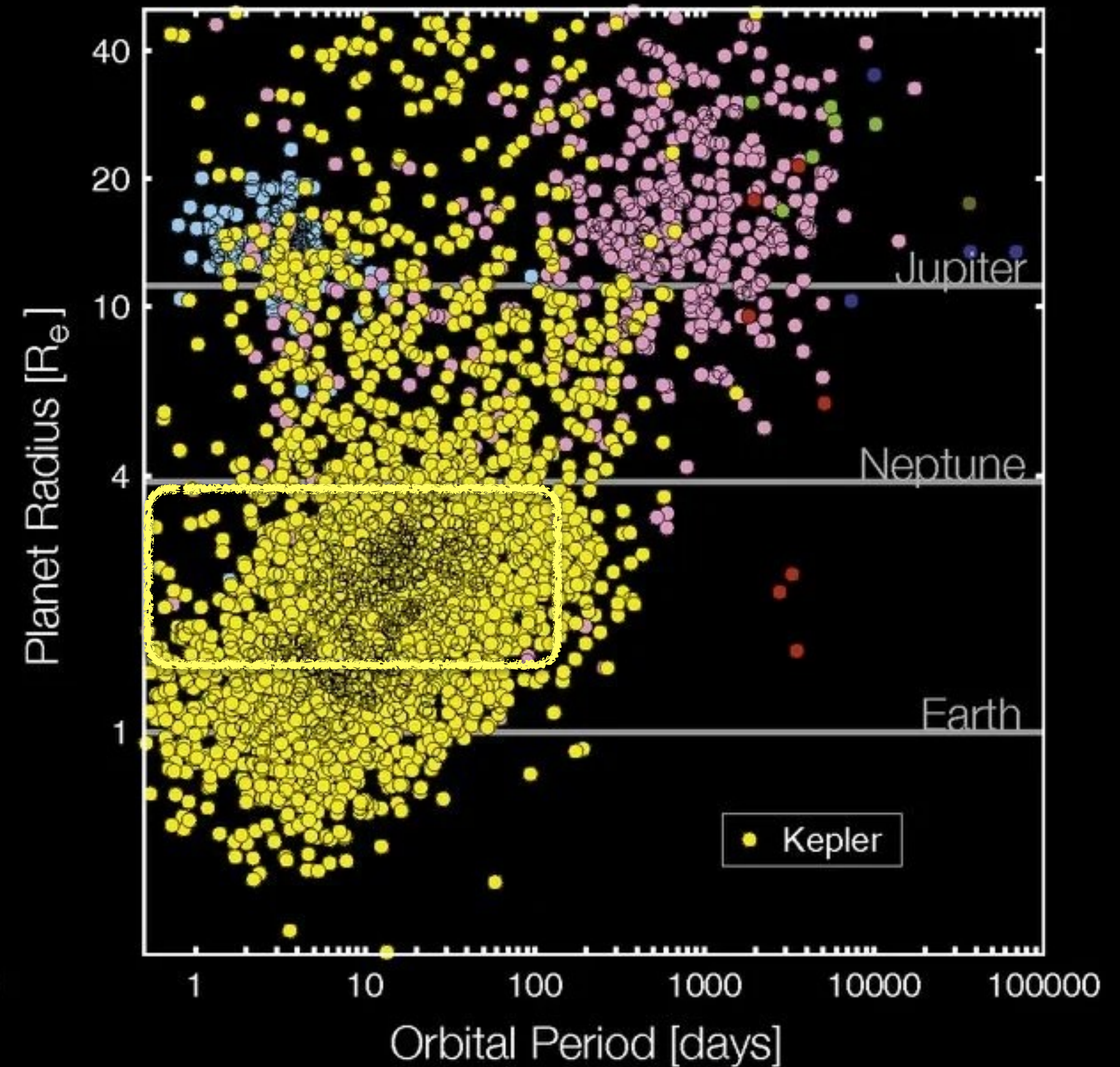
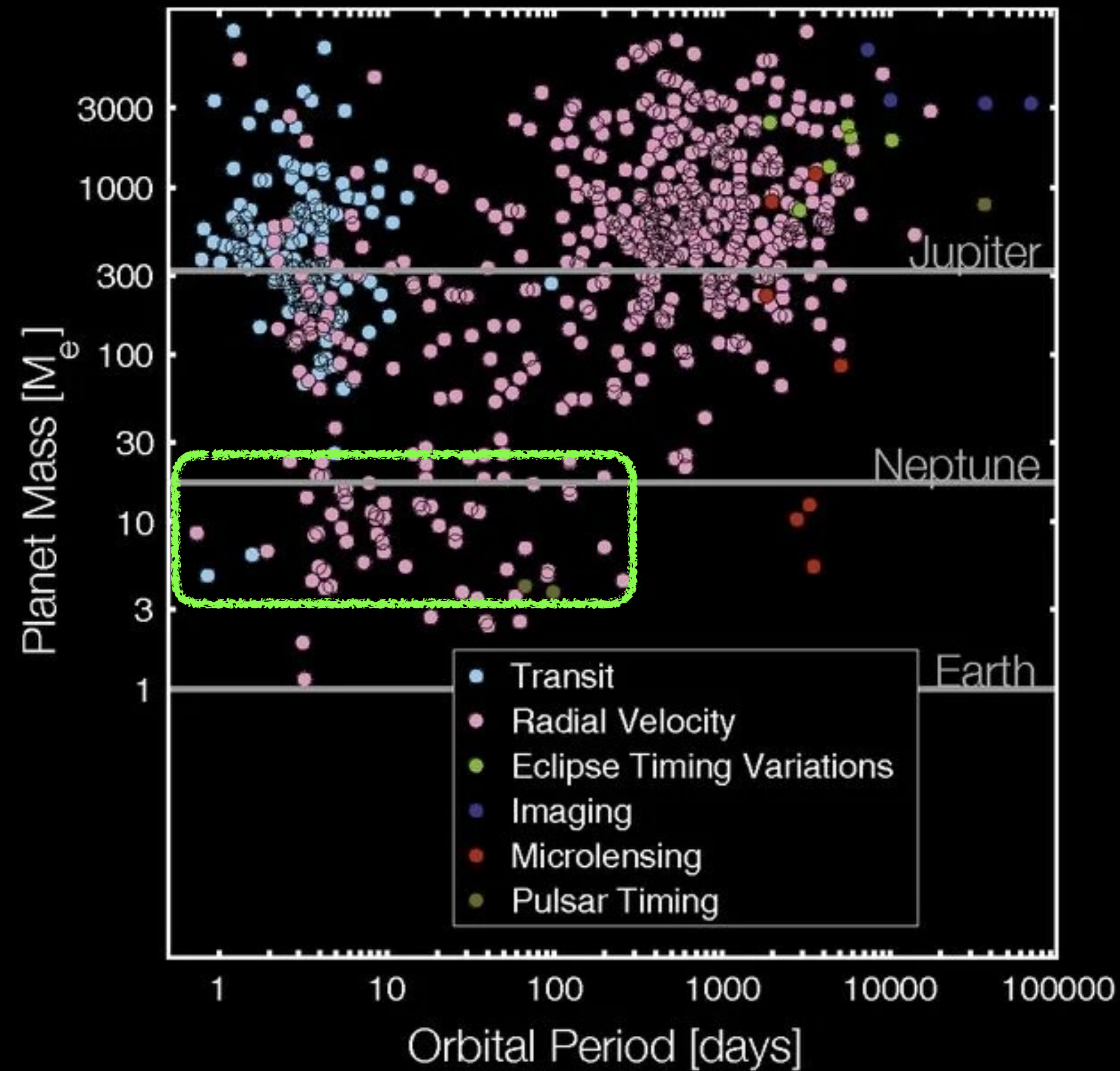
- Mixing-Demixing



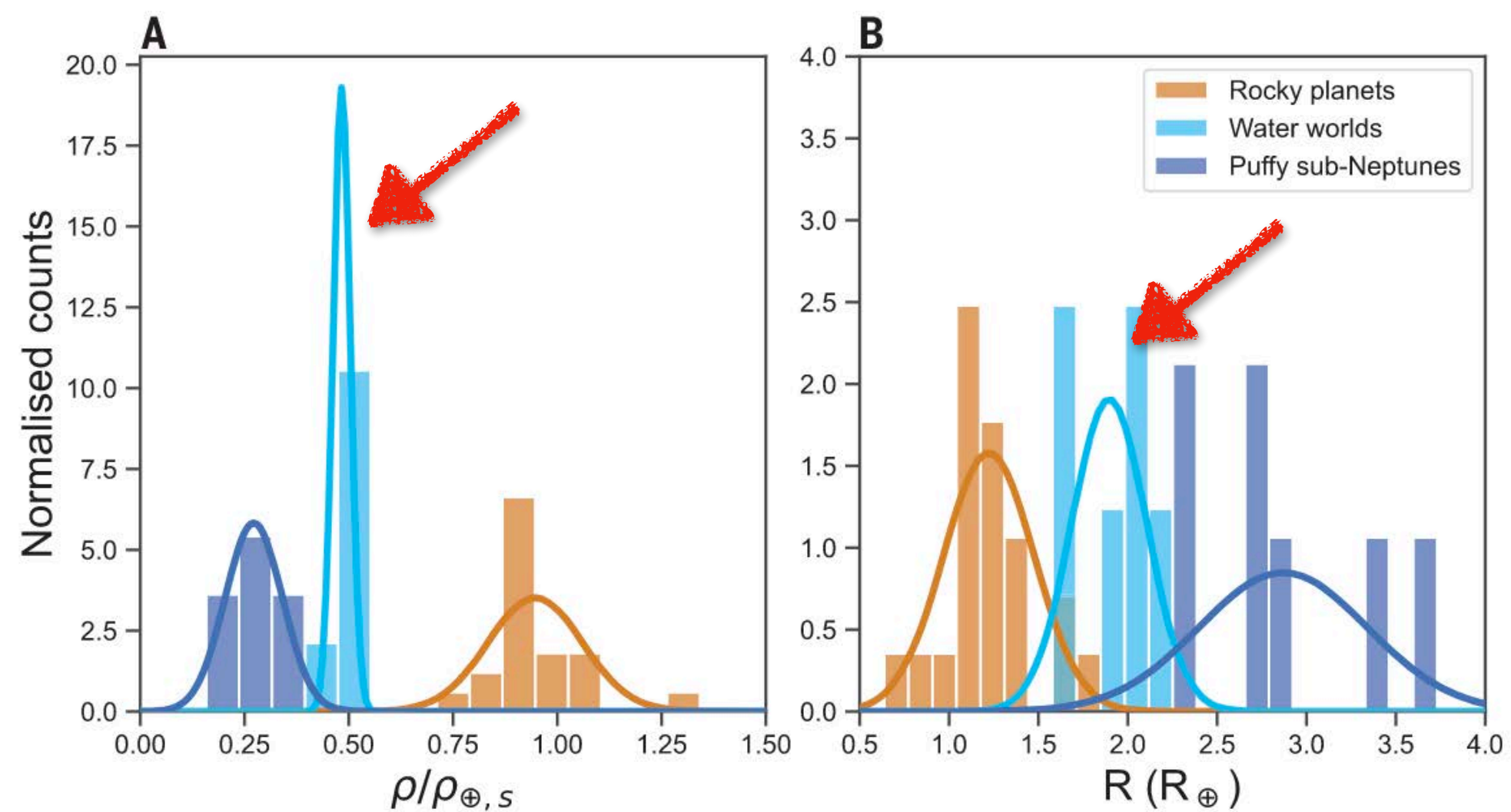
Sub-Neptunes



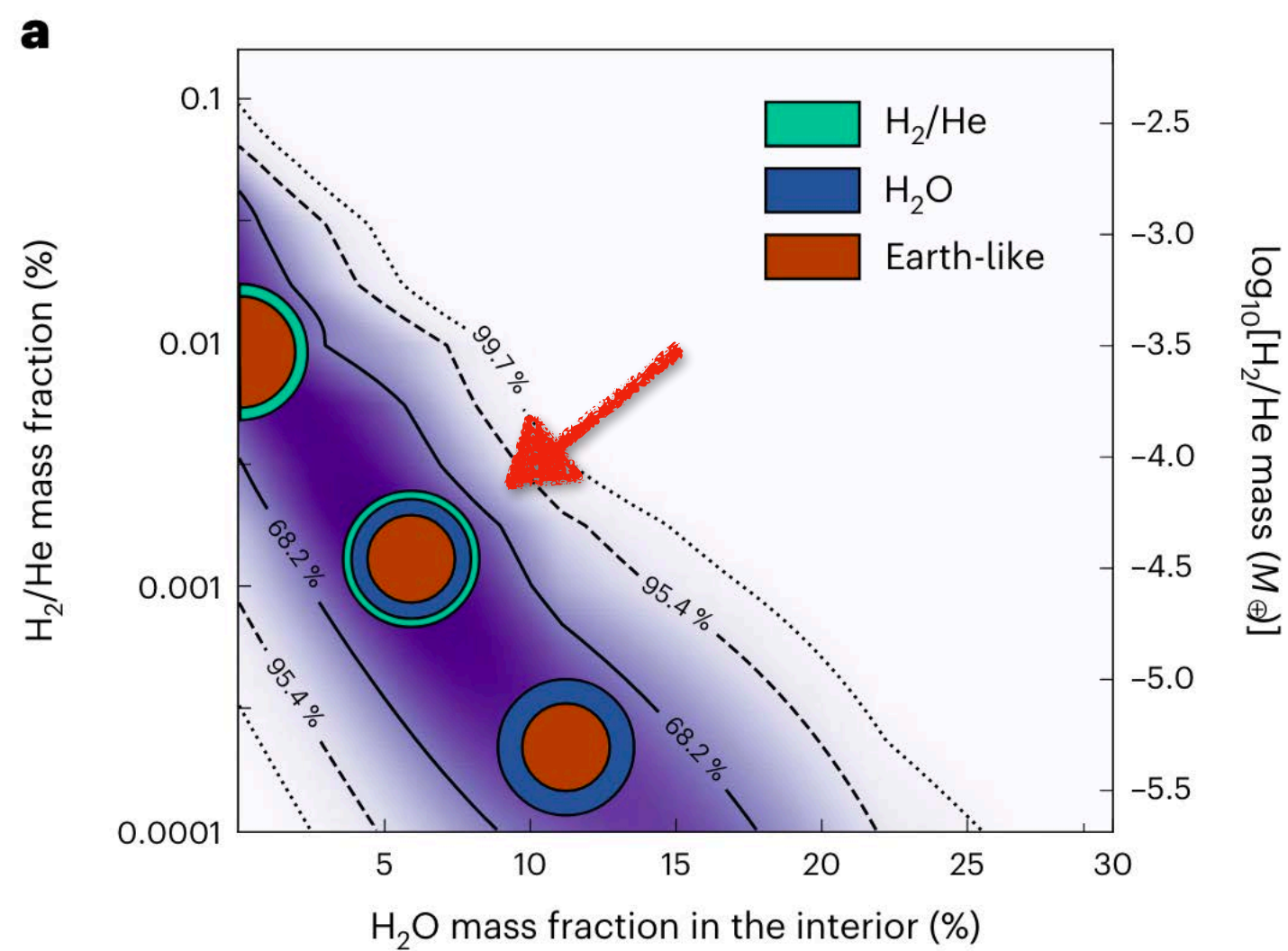
Close-in Sub-Neptunes



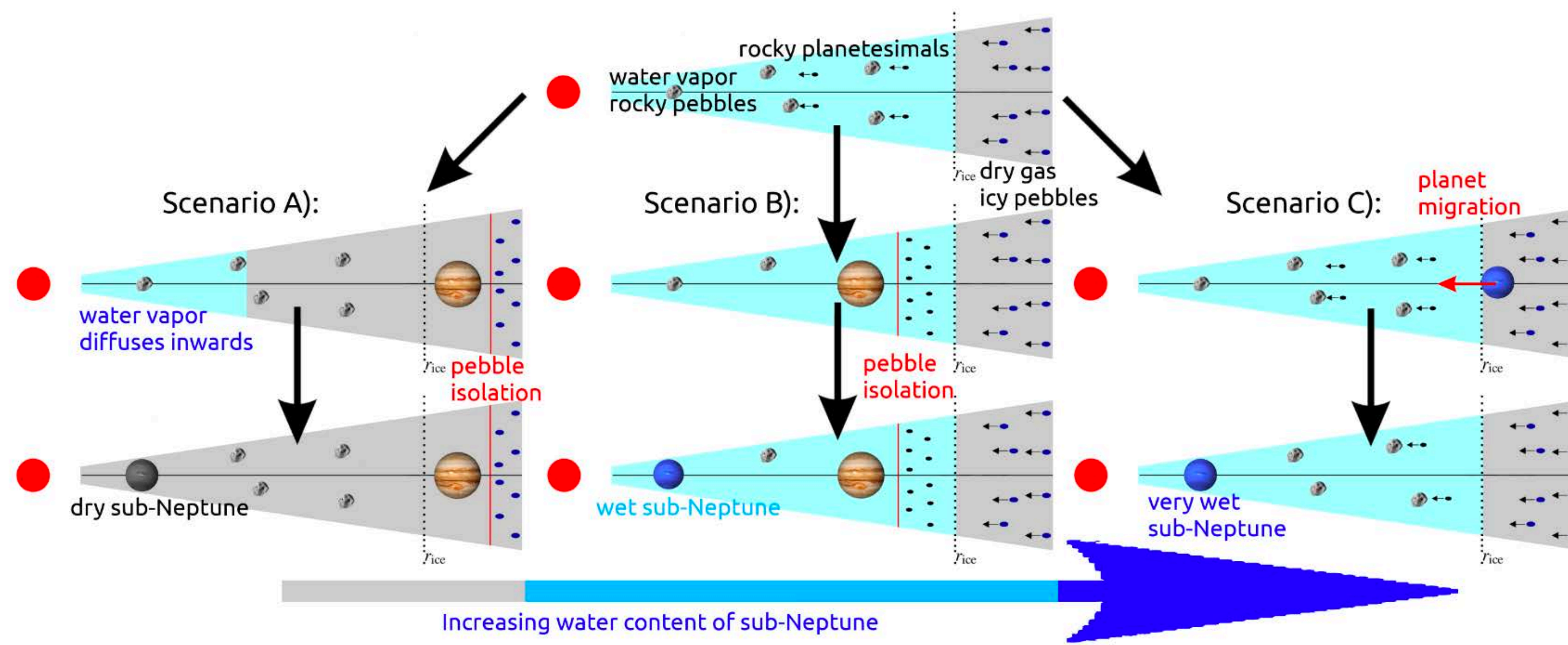
Water-rich Sub-Neptunes



Luque and Palle (2022) Science

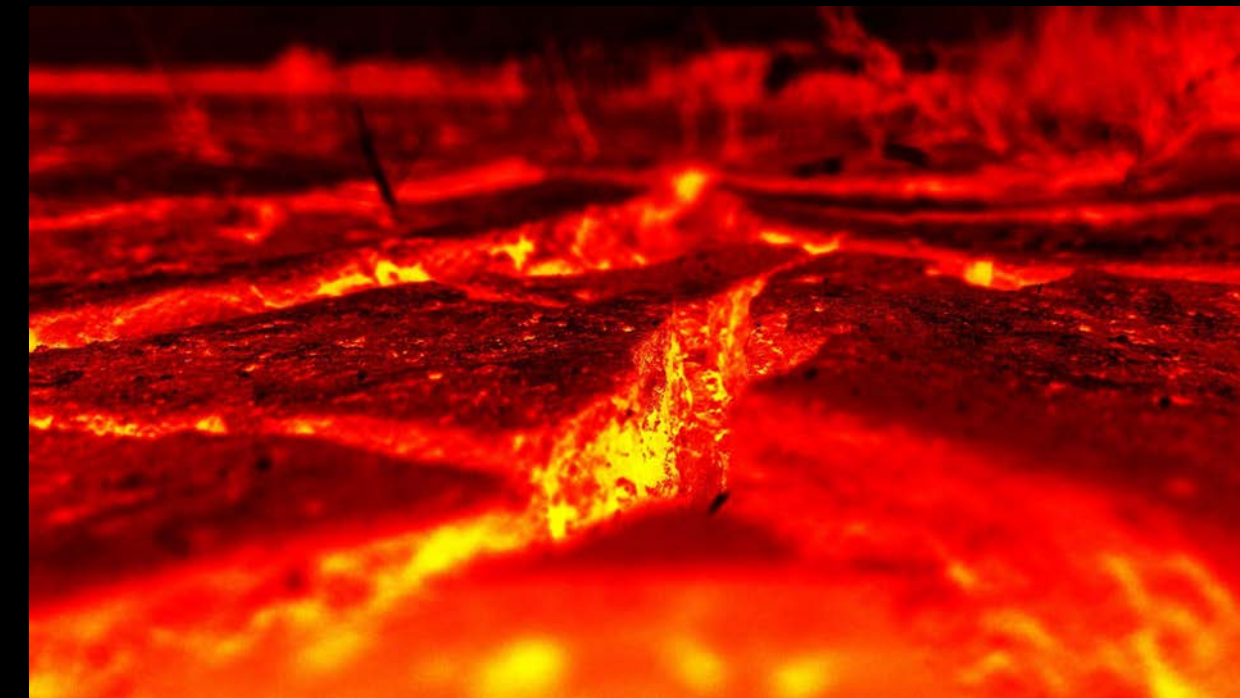
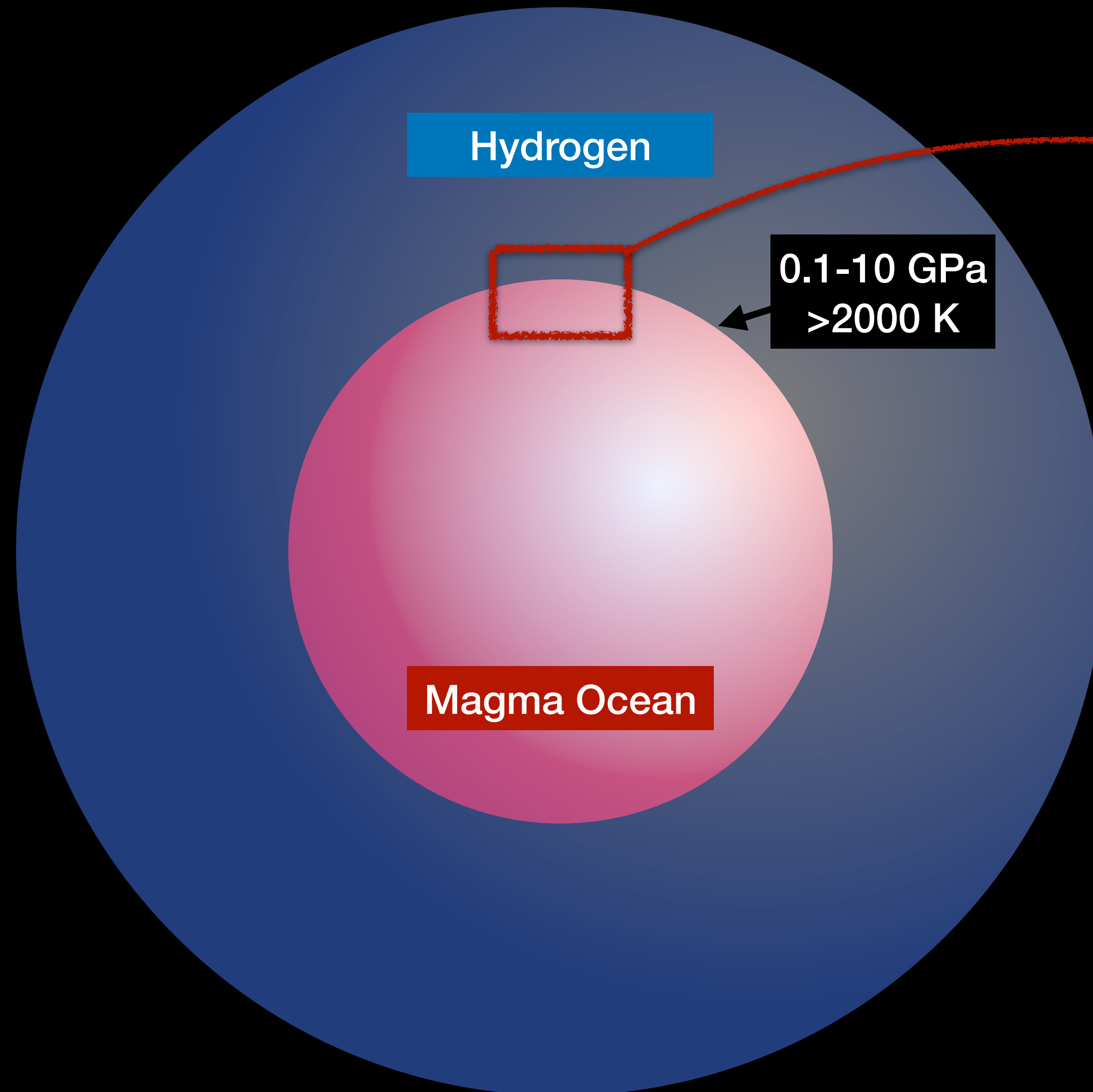


Kepler-138d, Piaulet et al. (2022) Nat. Astro.



Bitsch et al. (2021) Astron. & Astrophys.

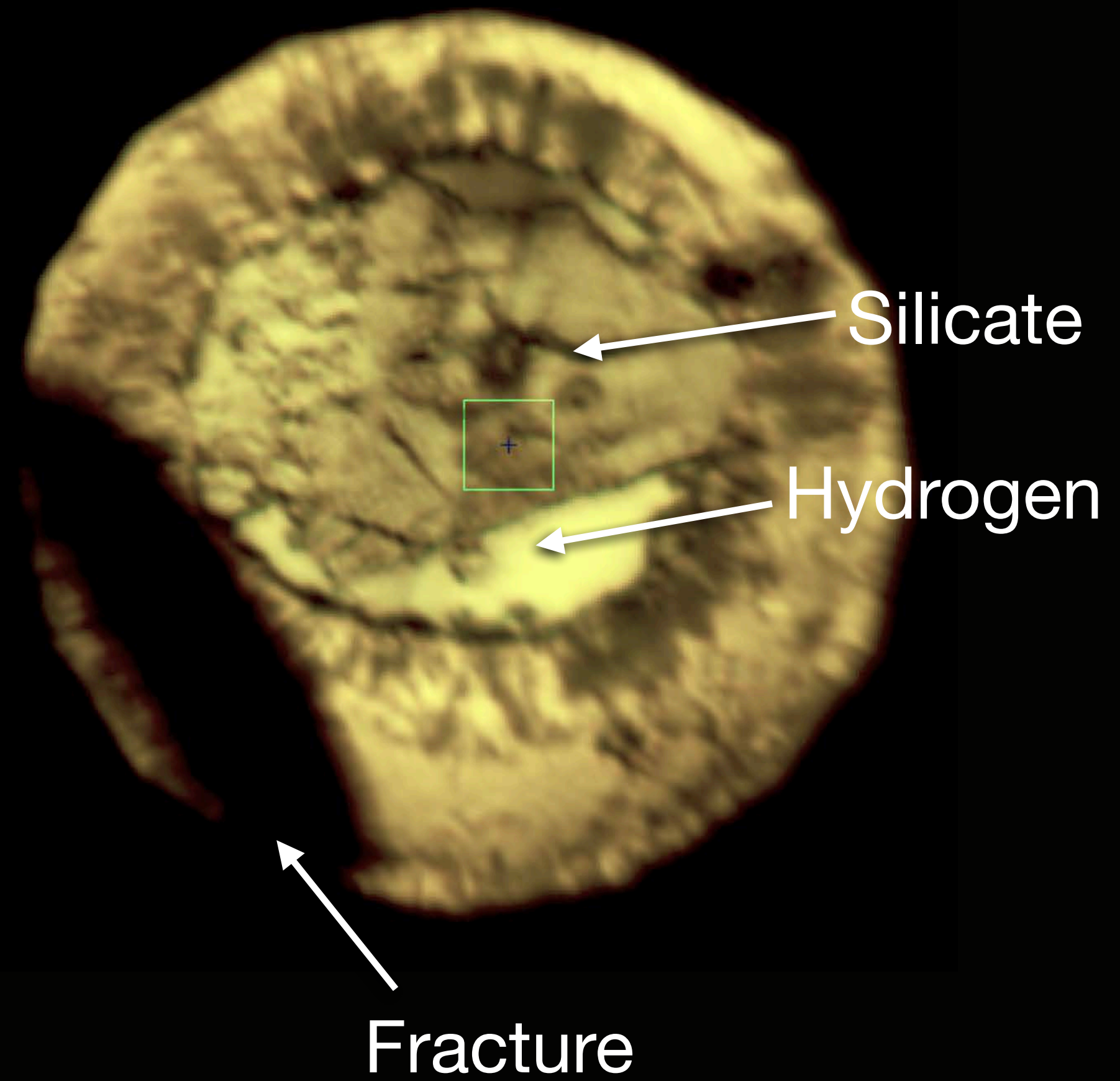
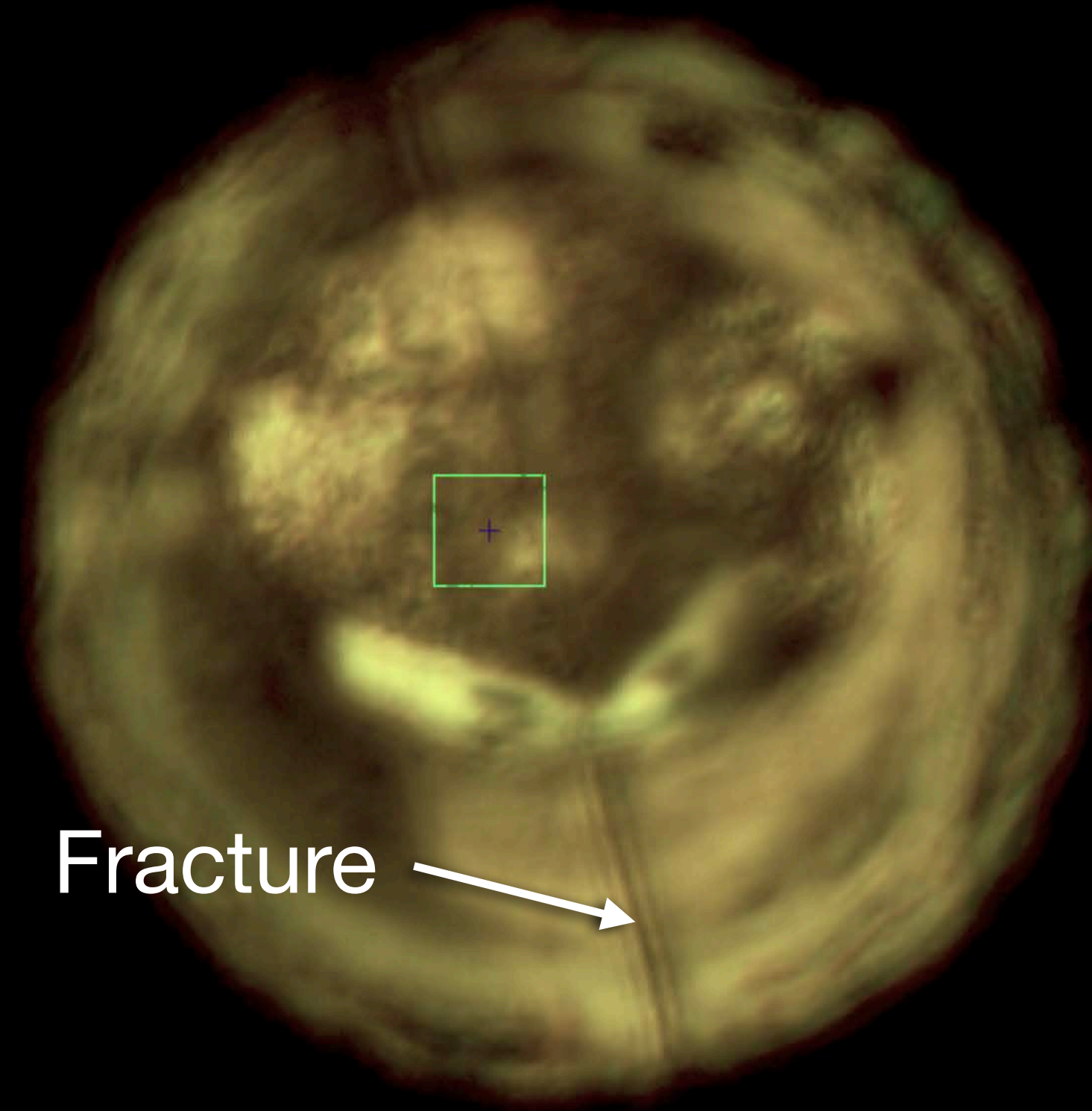
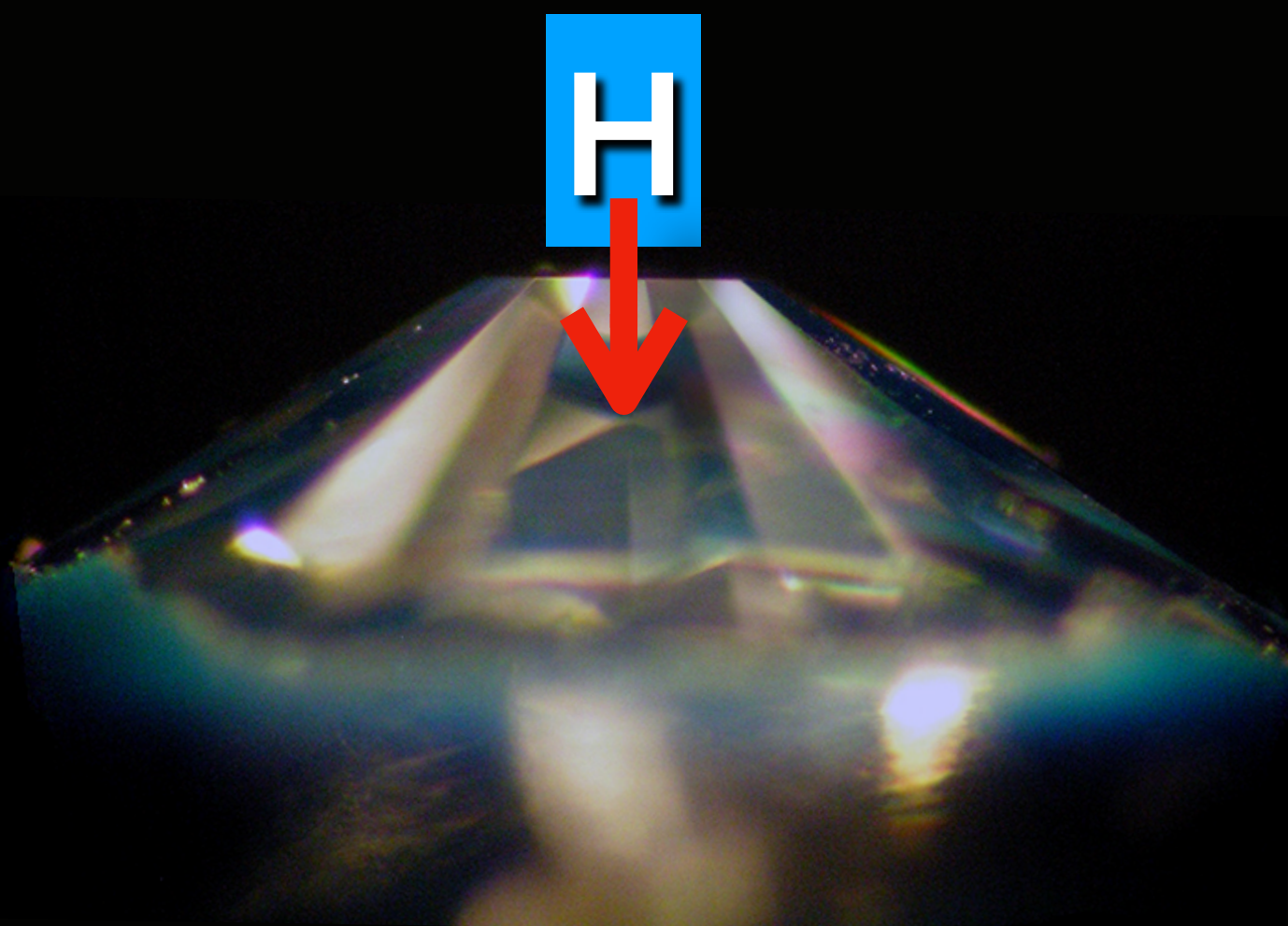
Interior Model for Sub-Neptunes



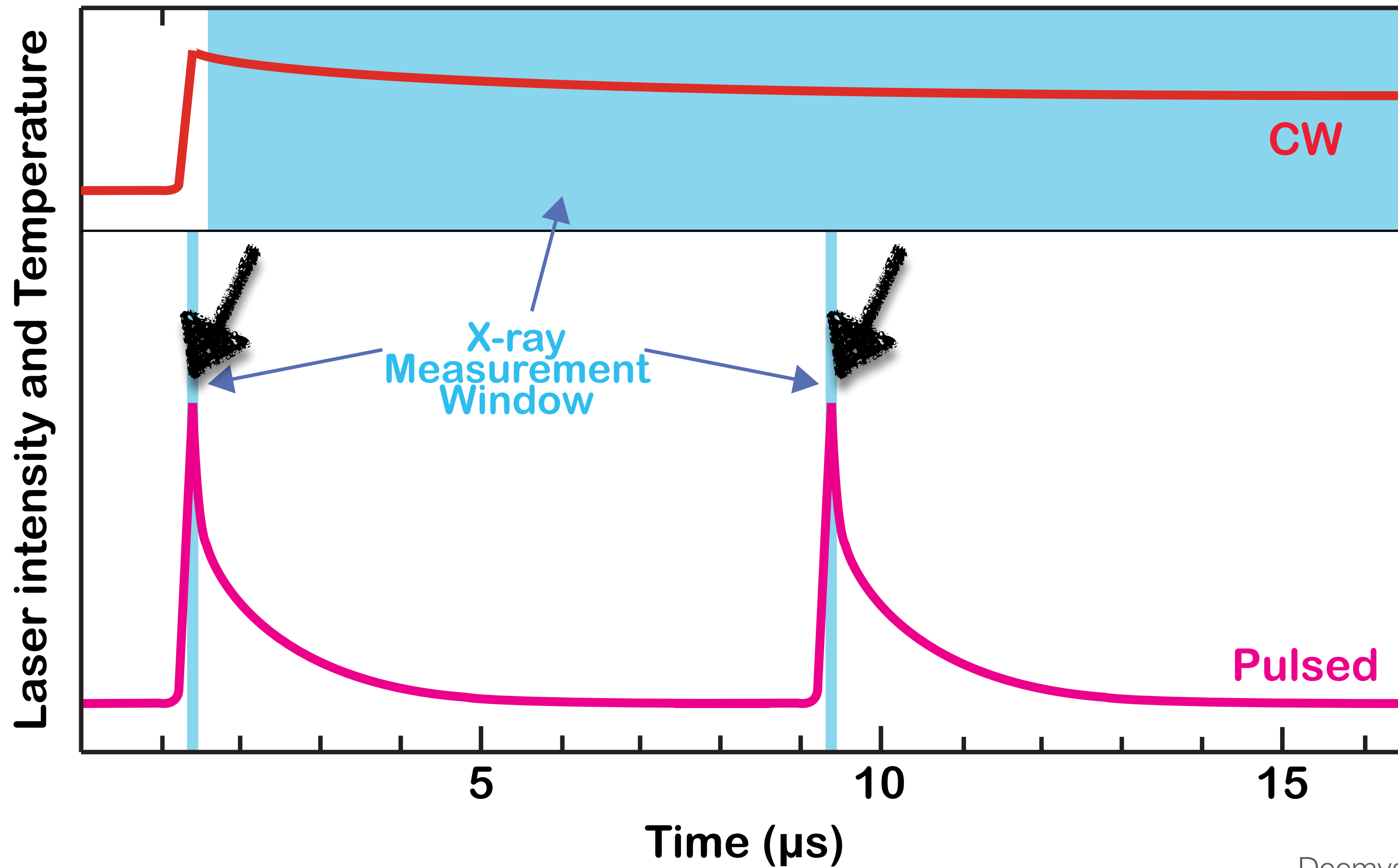
The magma ocean phase may last several gigayears

Vazan et al. (2018) ApJ

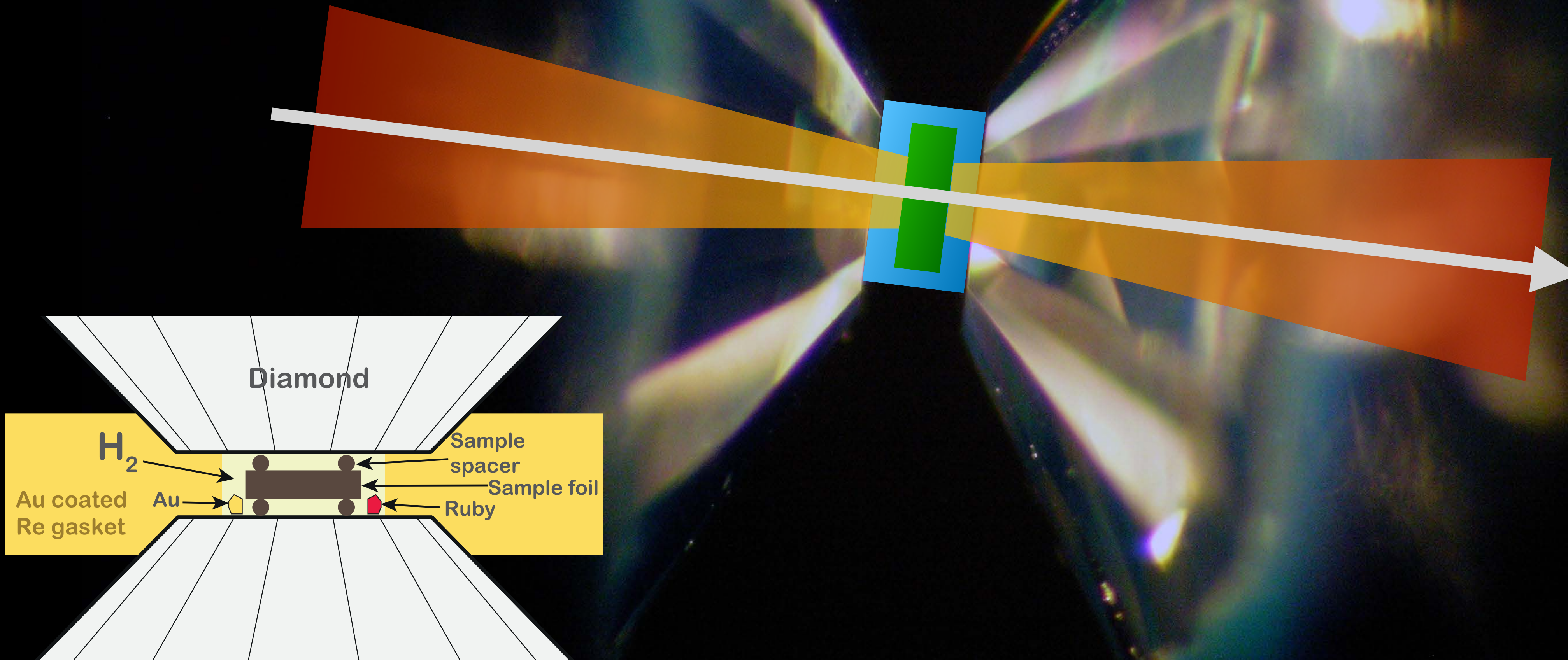
Diamond Embrittlement



Pulsed Laser Heating

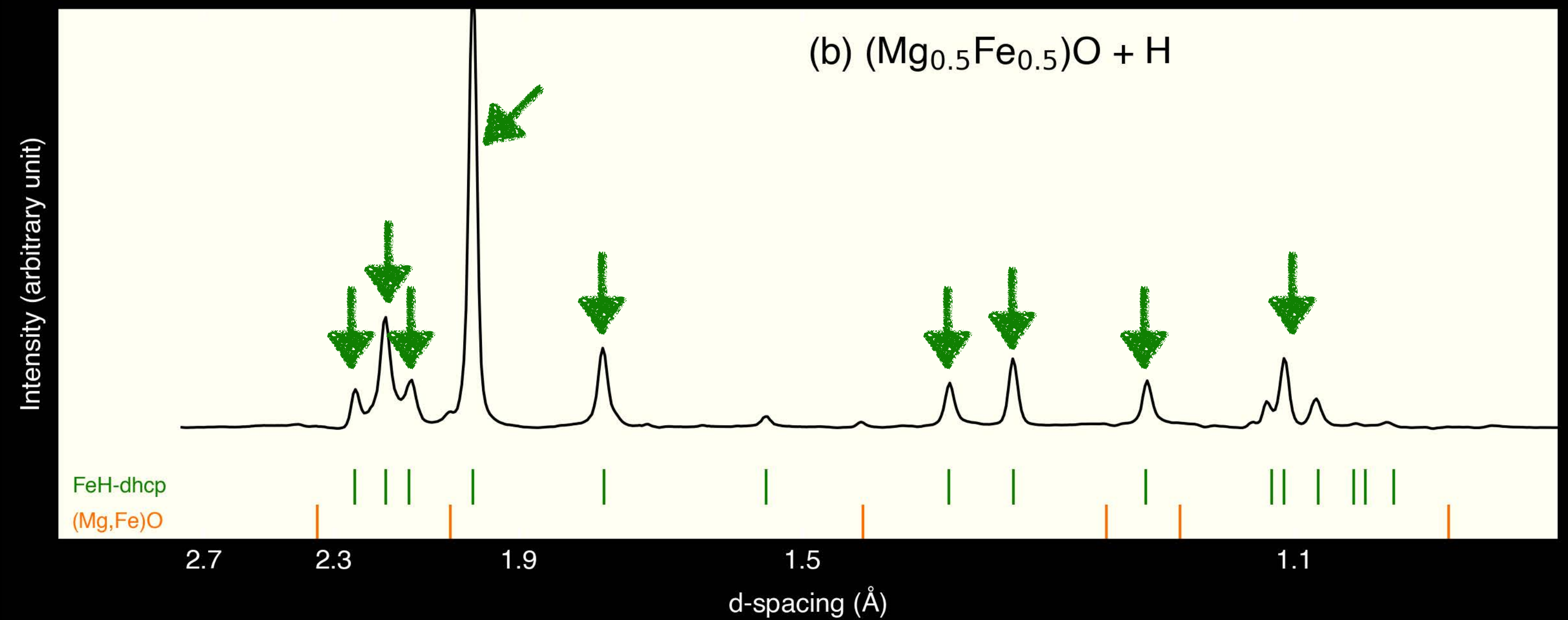
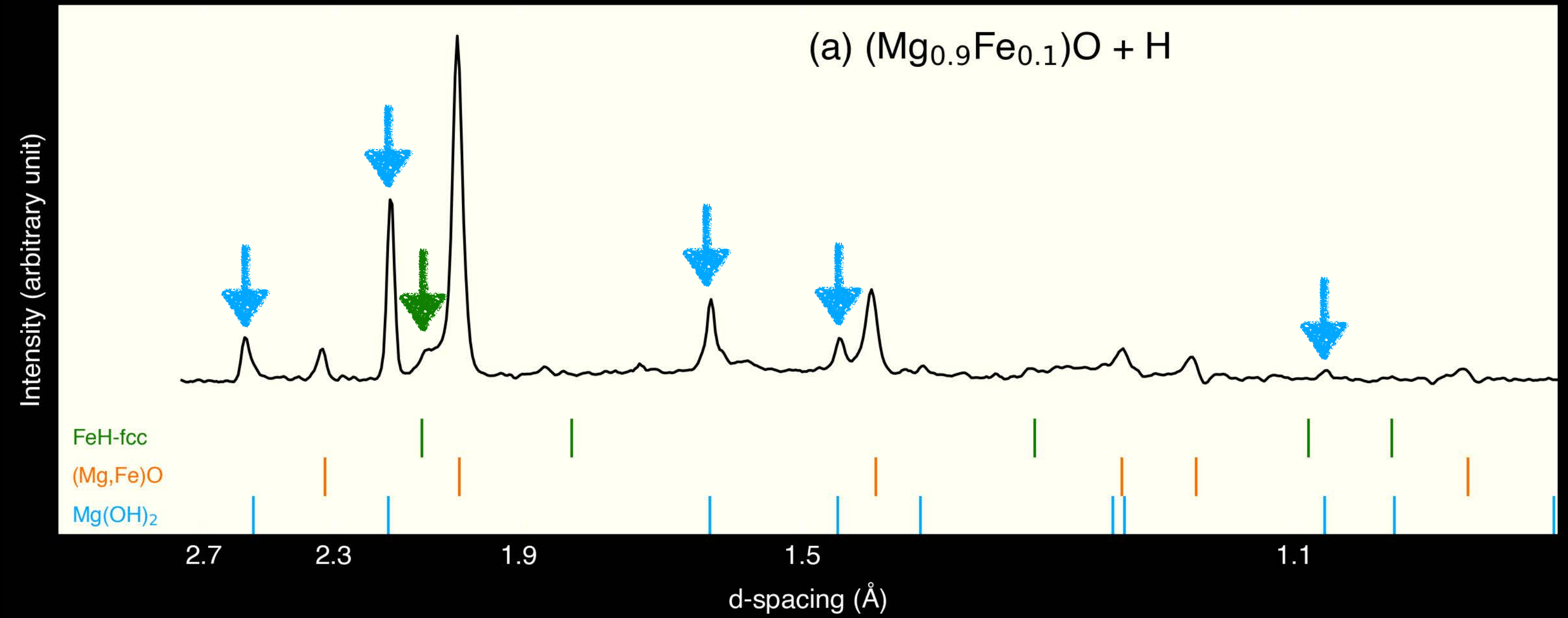
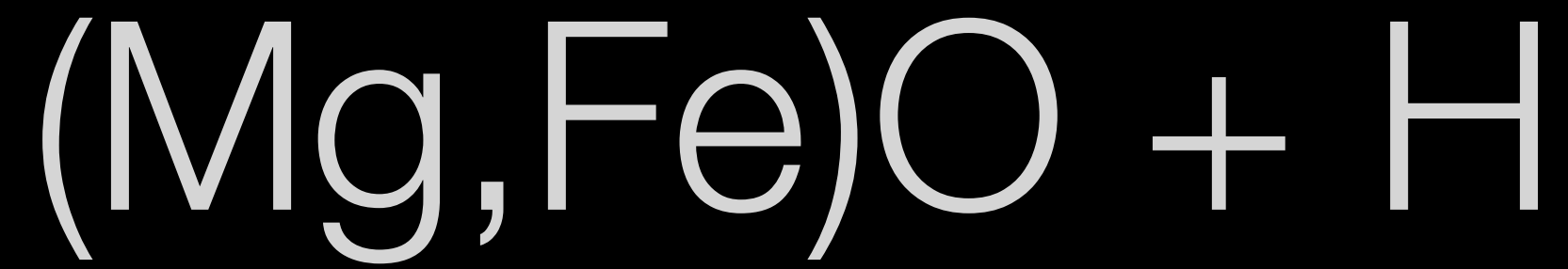


Diamond Anvil Cell



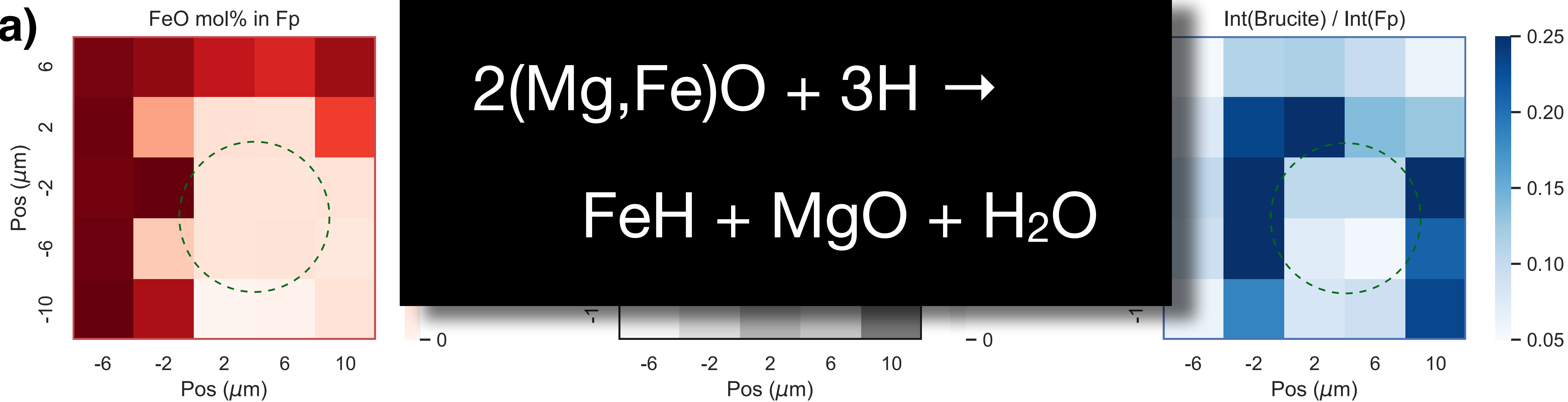


22-30 GPa, 1500-2000 K



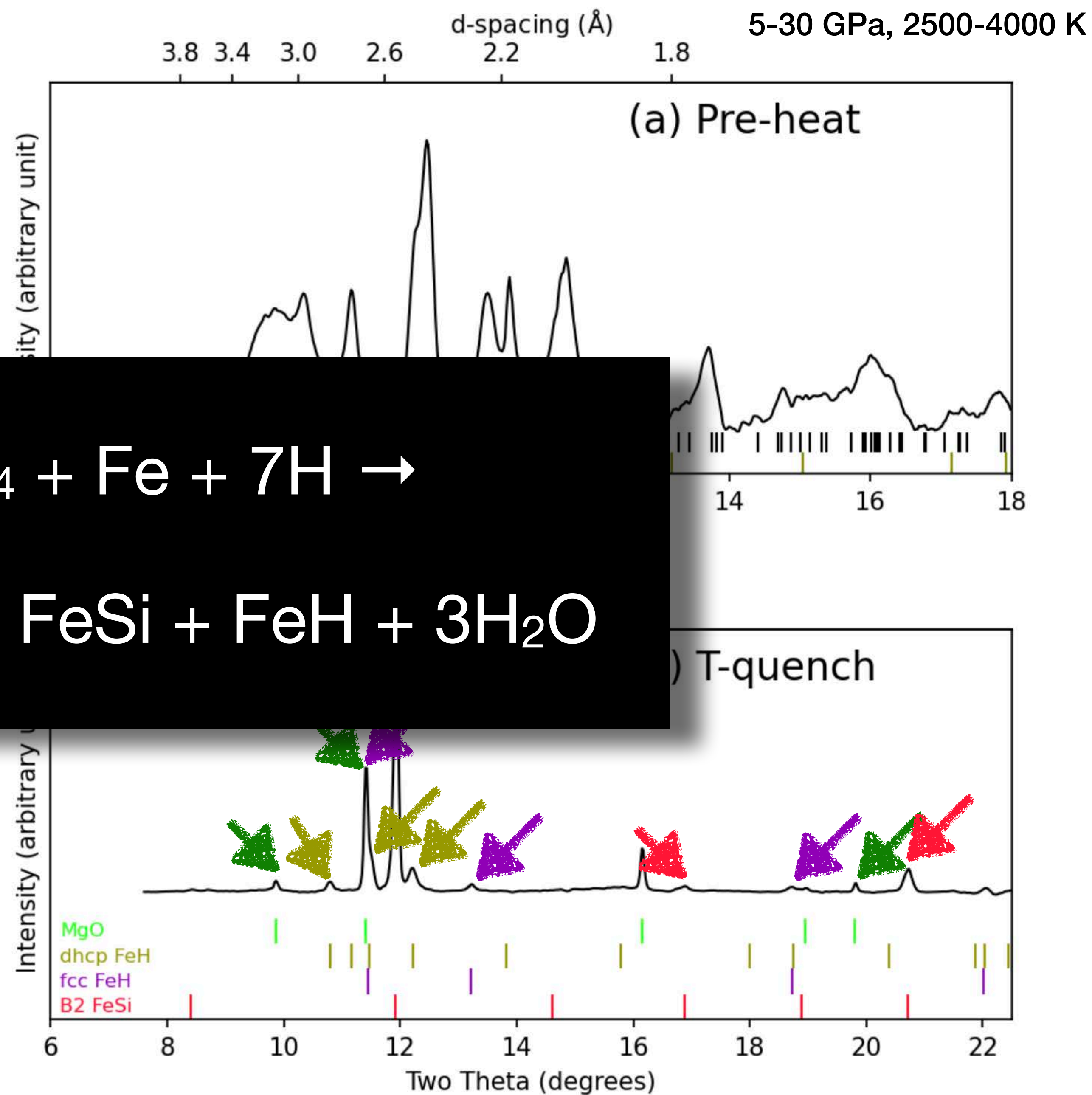
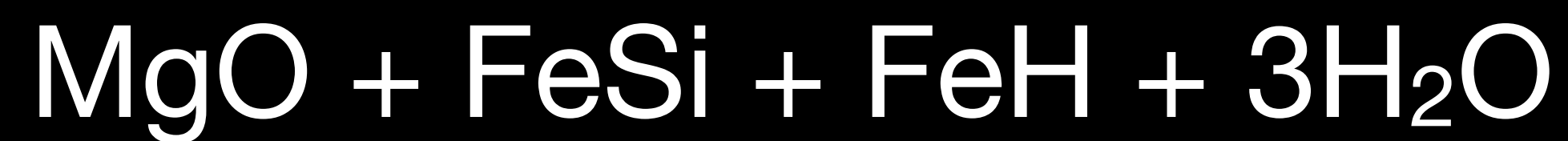
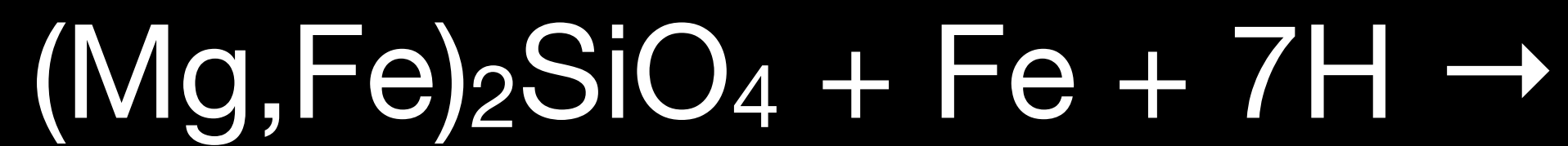


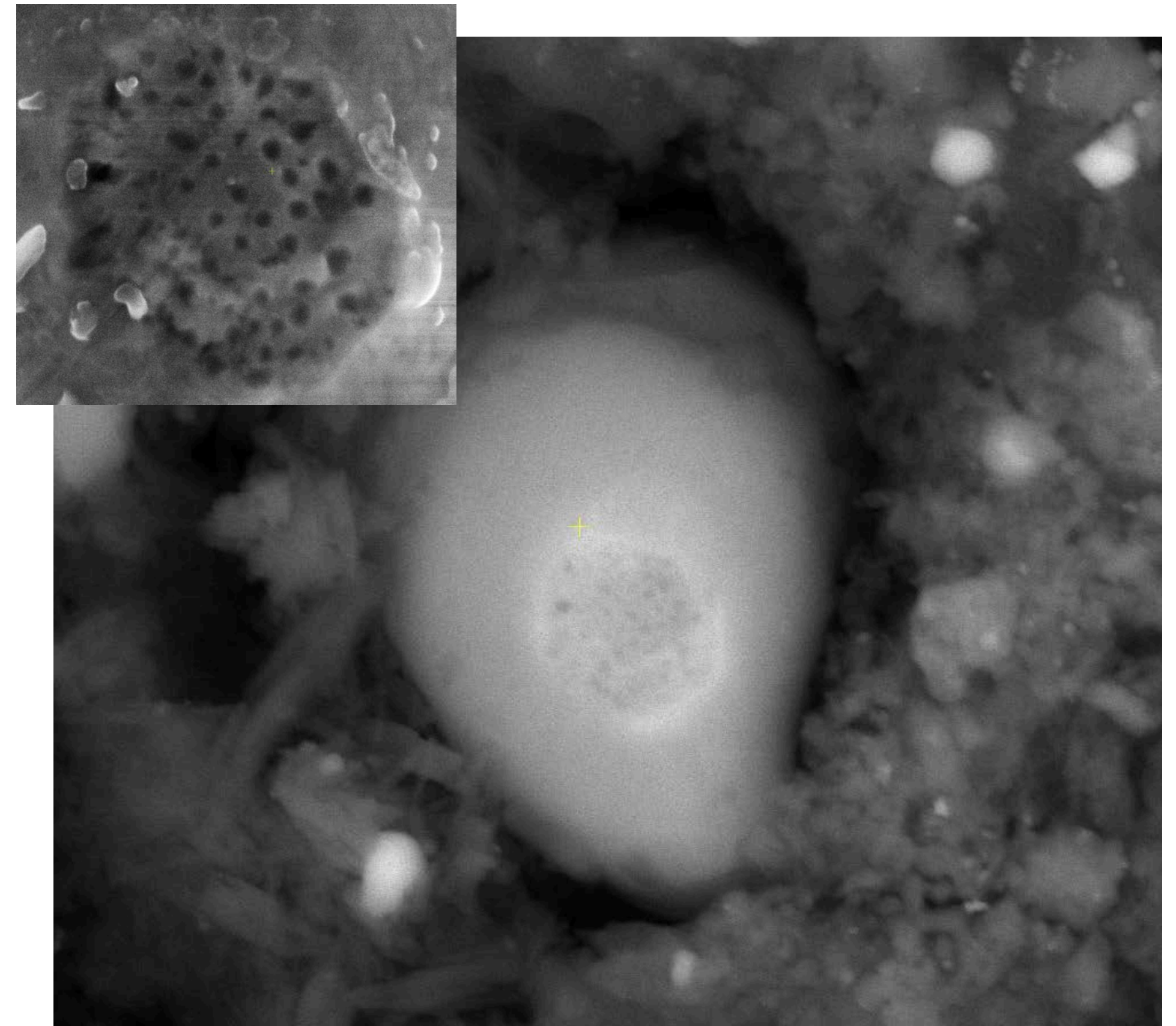
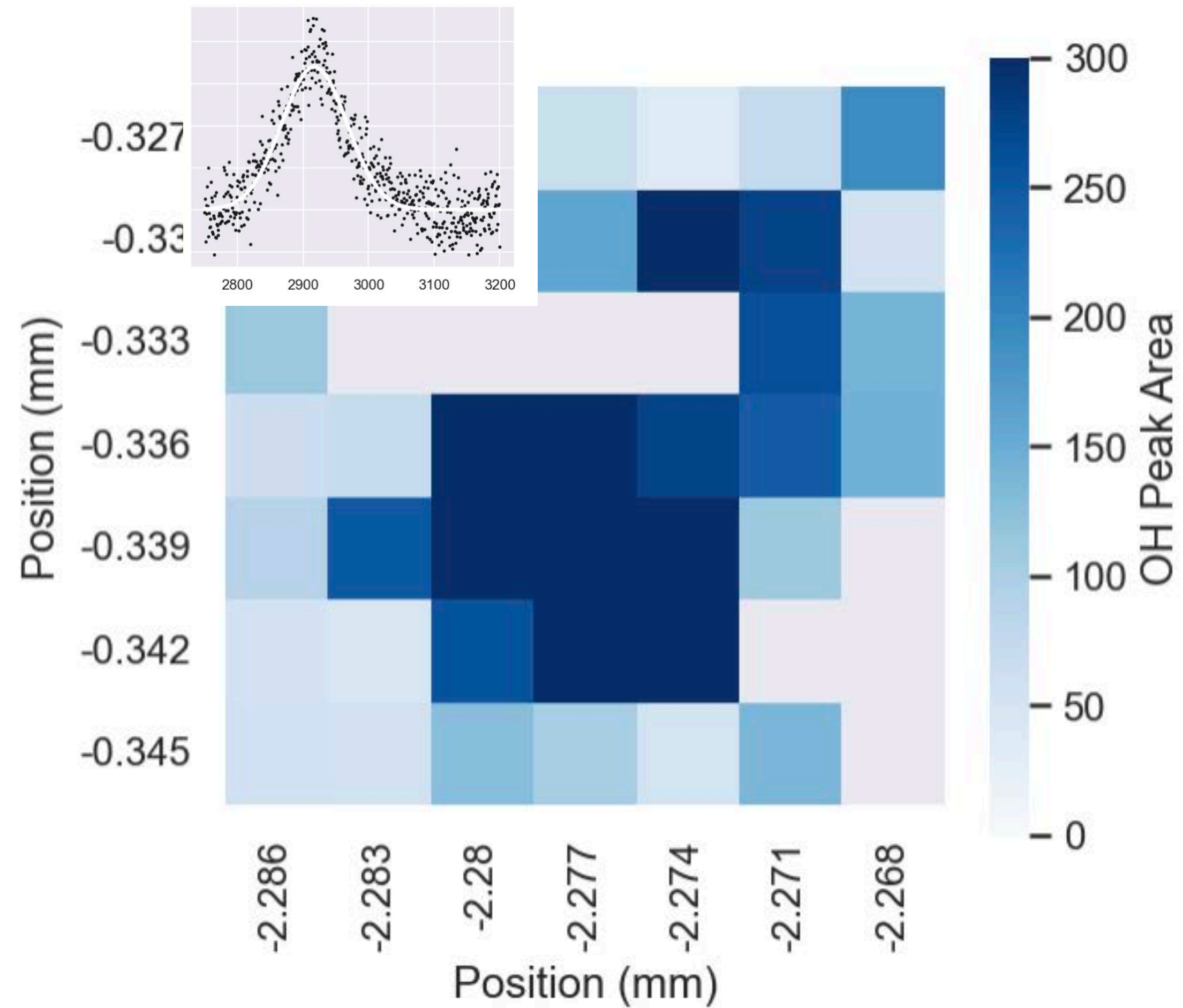
(a)





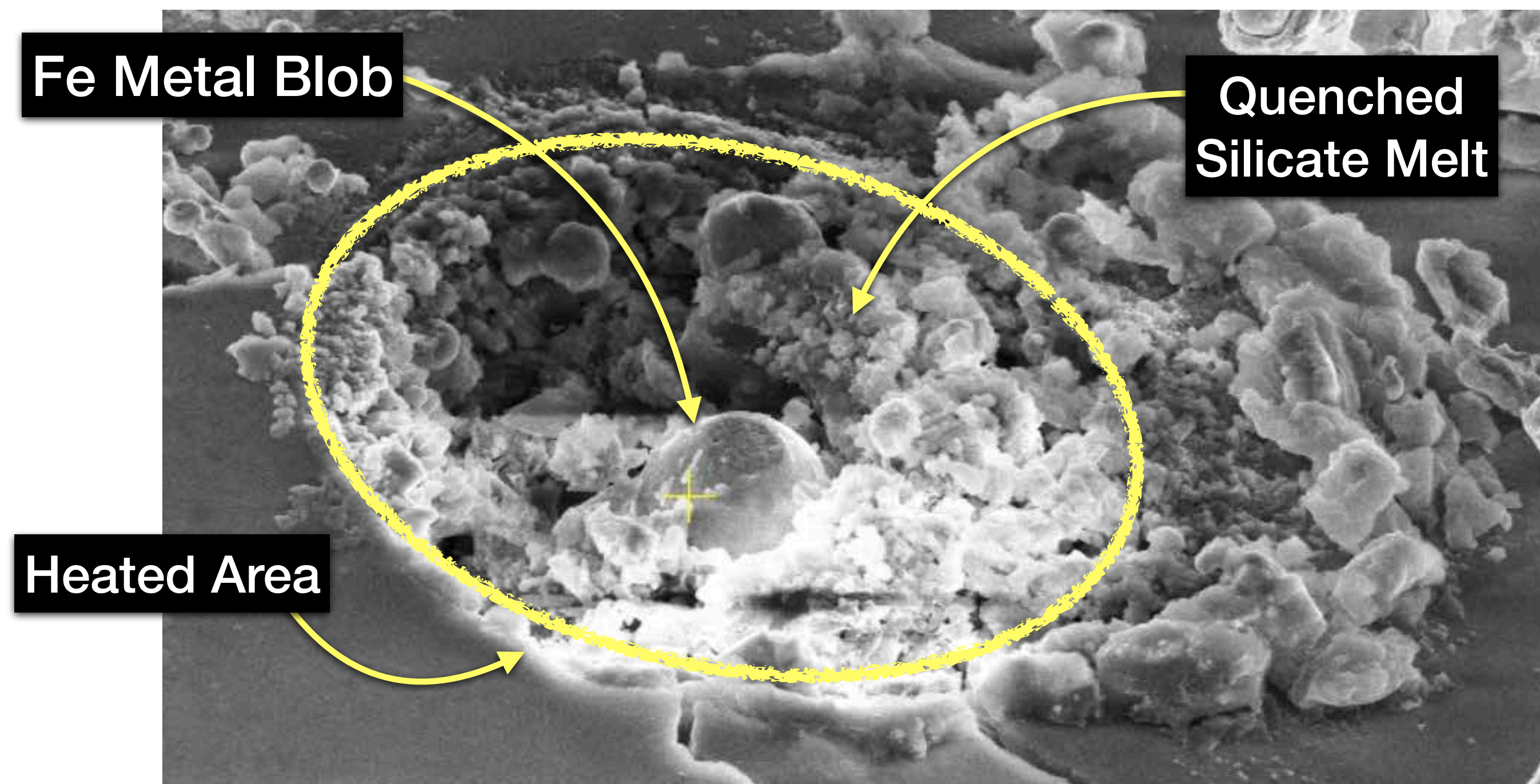
(Mg,Fe)₂SiO₄





Fayalite in a H medium at 12 GPa and 3000 K

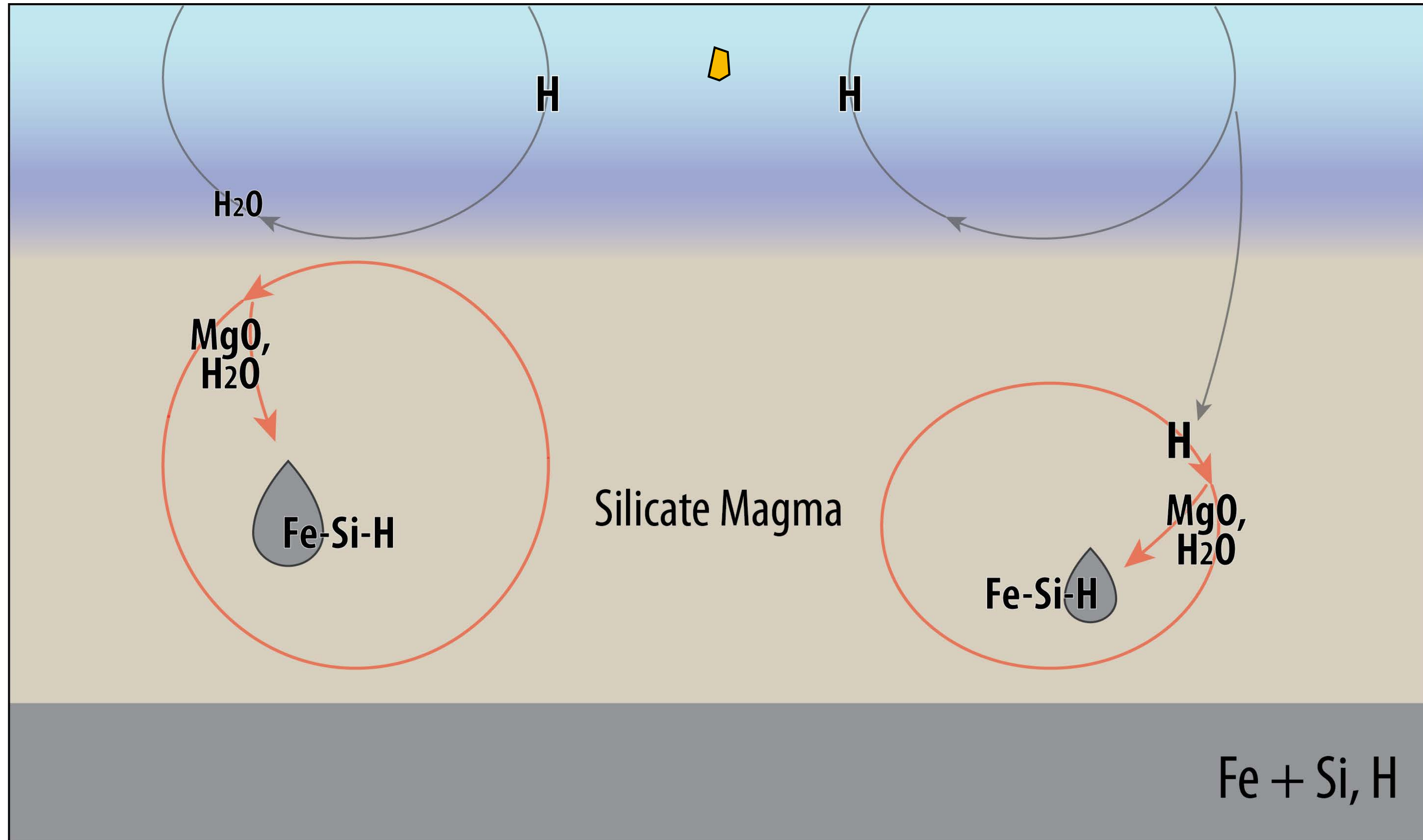
Electron Microscopy



Fayalite in a H medium at 12 GPa and 3000 K.

- **$\text{FeO} + \text{H}_2 \rightarrow \text{Fe} + \text{H}_2\text{O}$**
- **$\text{SiO}_2 + 2\text{H}_2 \rightarrow \text{Si} + 2\text{H}_2\text{O}$**
- **$\text{Fe} + 0.5\text{H}_2 \rightarrow \text{FeH}$**

Conversion of H to H₂O Planet



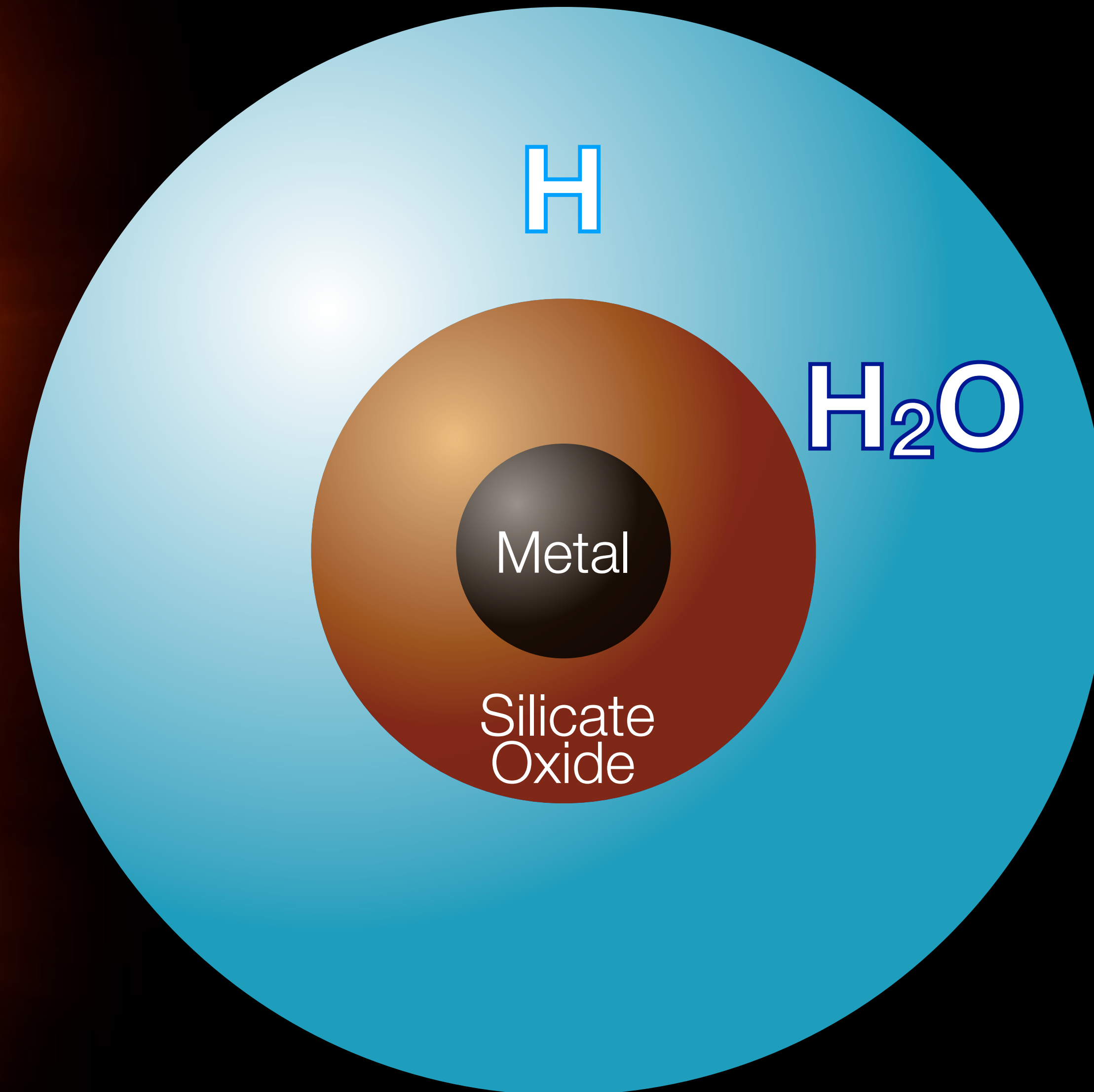
Impact of Si Reduction

Production of a large amount of H₂O

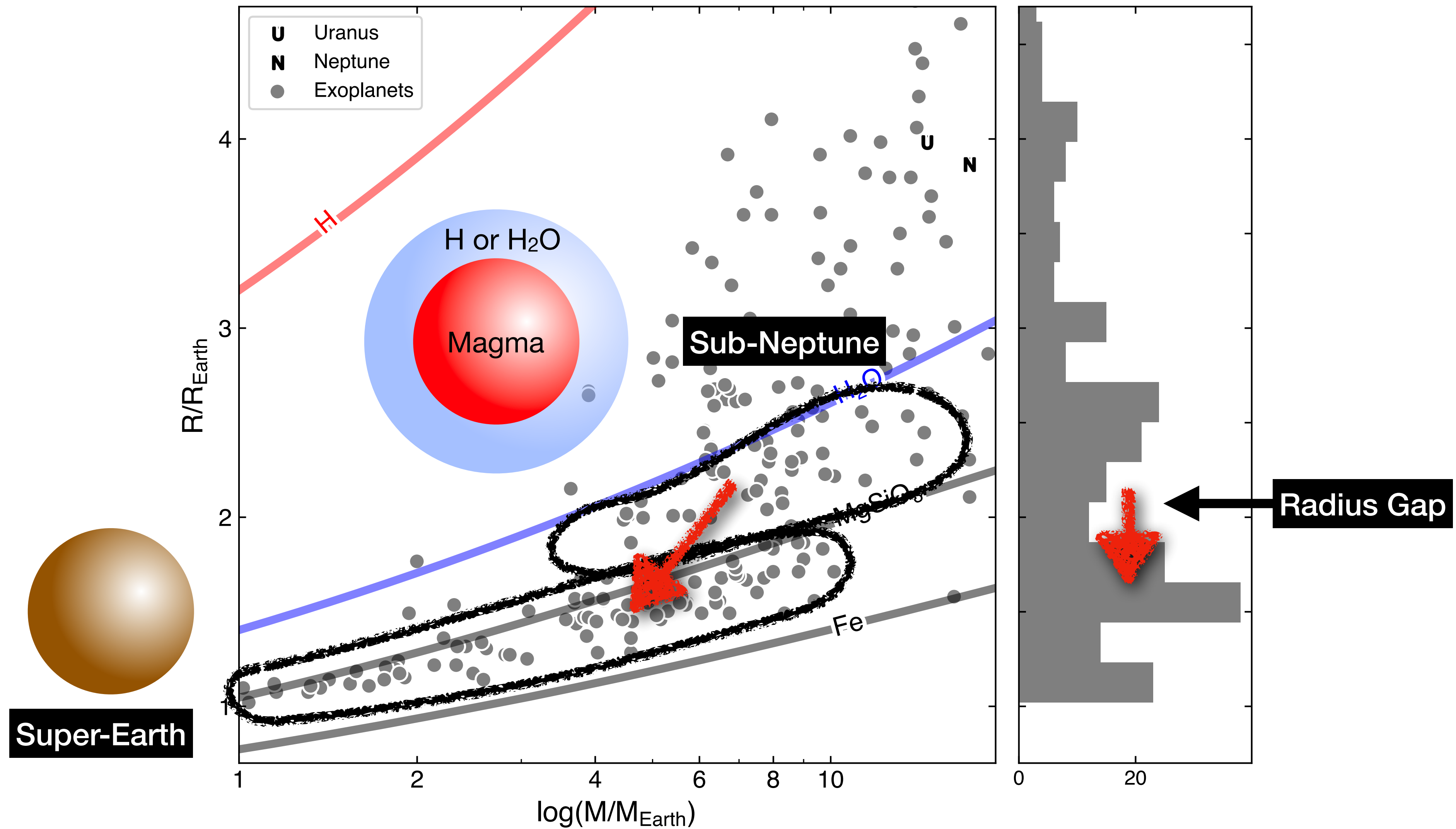
Growth of the metallic part and shrink of the rocky part of the core

Magma converts from silicate to oxide

Planet Conversion



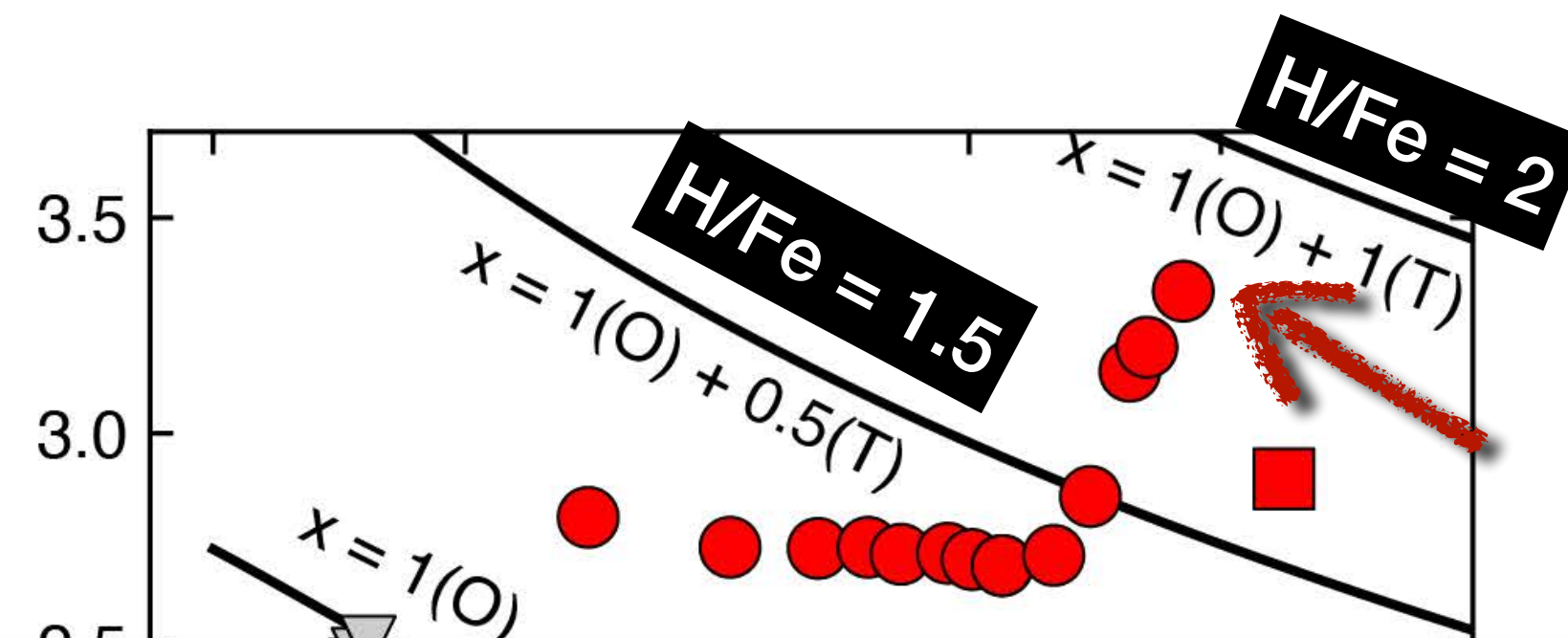
Sub-Neptunes and Super-Earths



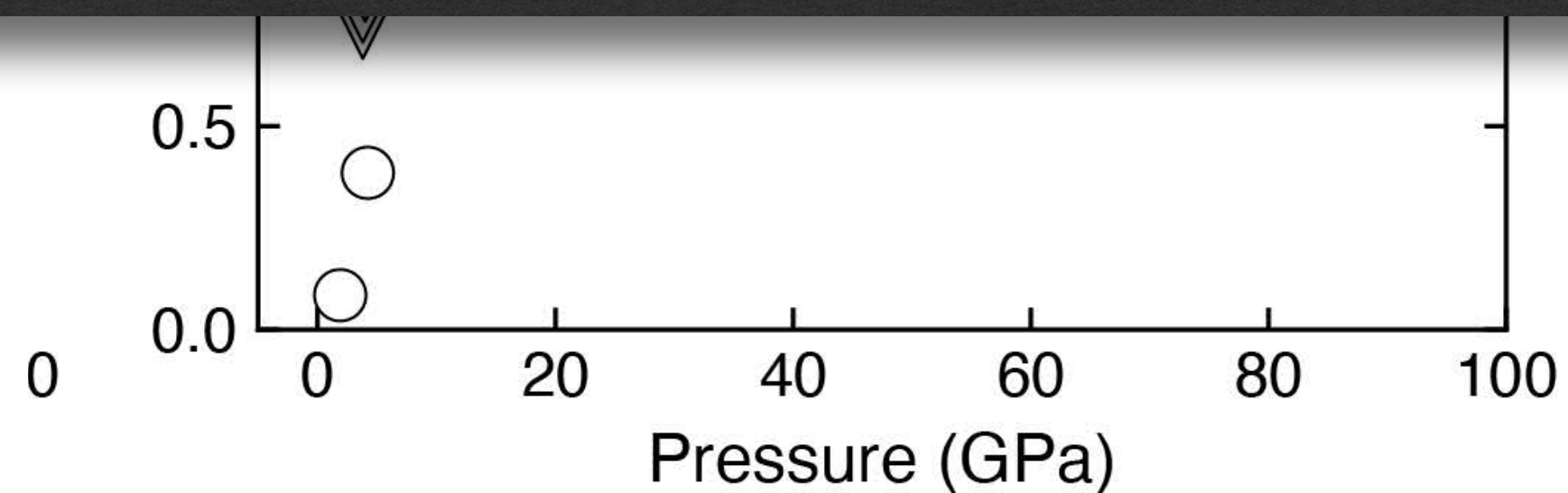
Mixing-Demixing



Super-Stoichiometric FeH_x Liquid



Very large amount of H can be ingassed in molten Fe metal.

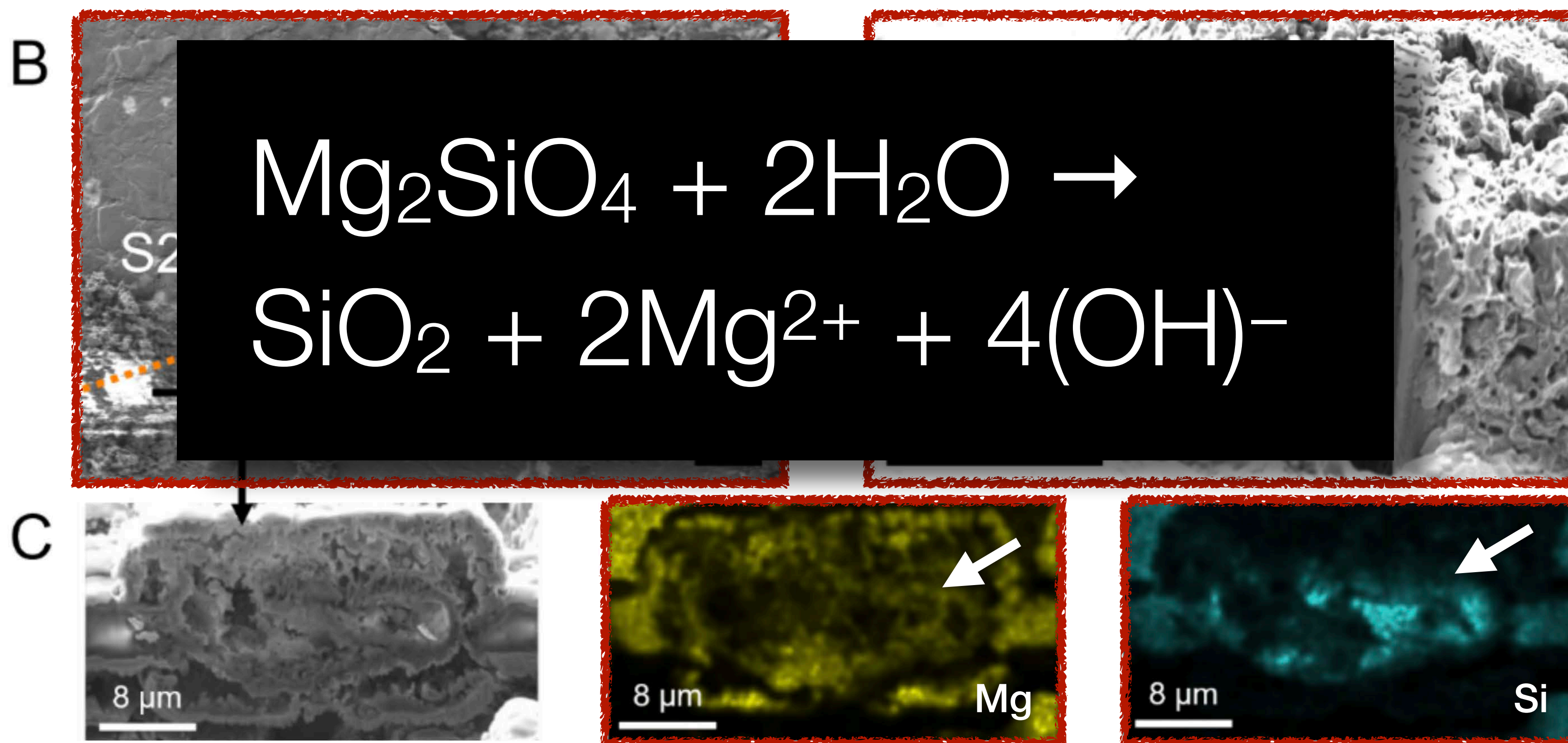




H₂O-Mg₂SiO₄

Atomic-scale mixing between MgO and H₂O in the deep interiors of water-rich planets

Taehyun Kim¹, Stella Chariton², Vitali Prakapenka², Anna Pakhomova³, Hanns-Peter Liermann³, Zhenxian Liu⁴, Sergio Speziale⁵, Sang-Heon Shim⁶ and Yongjae Lee¹

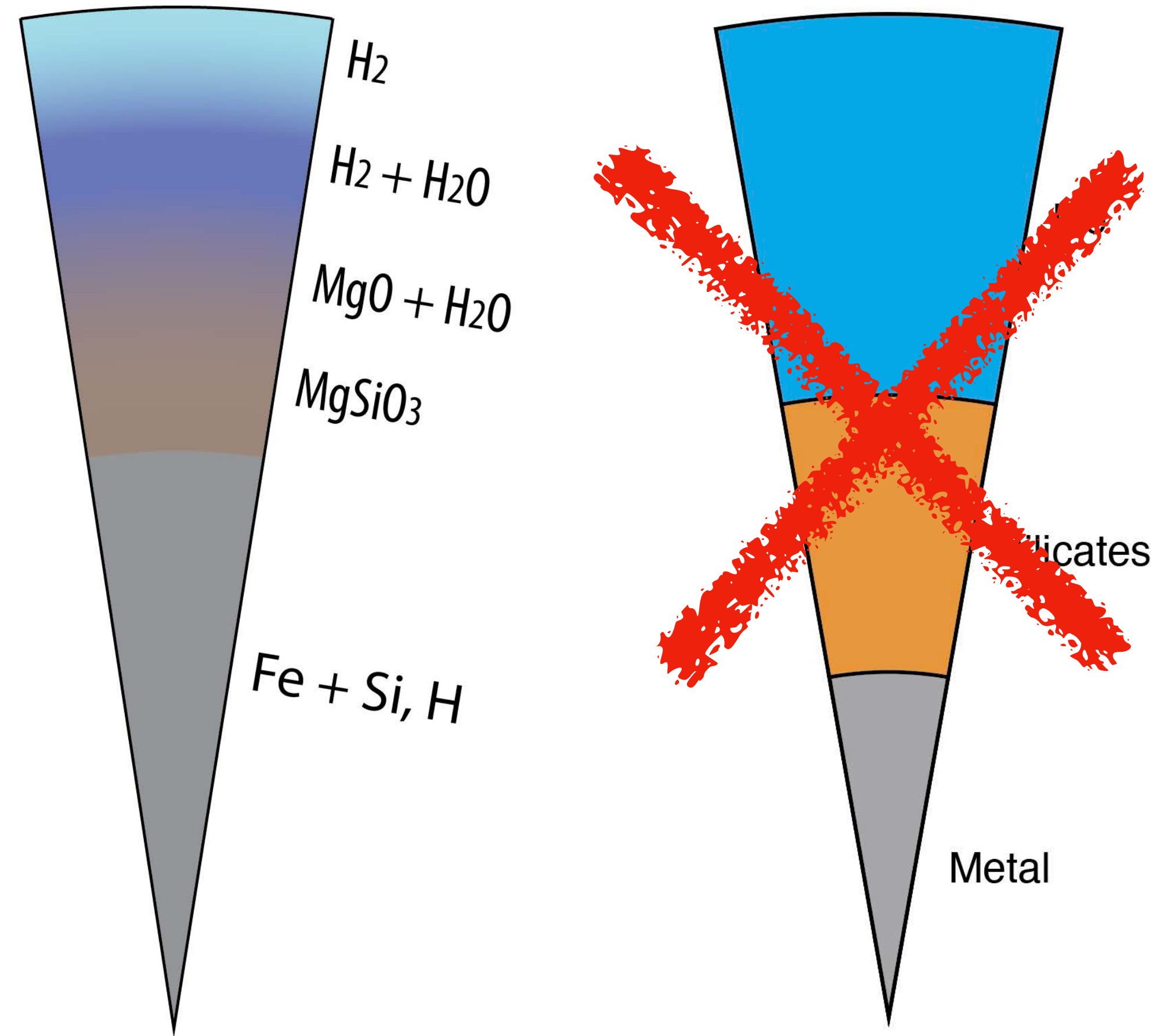


Mixing-Demixing

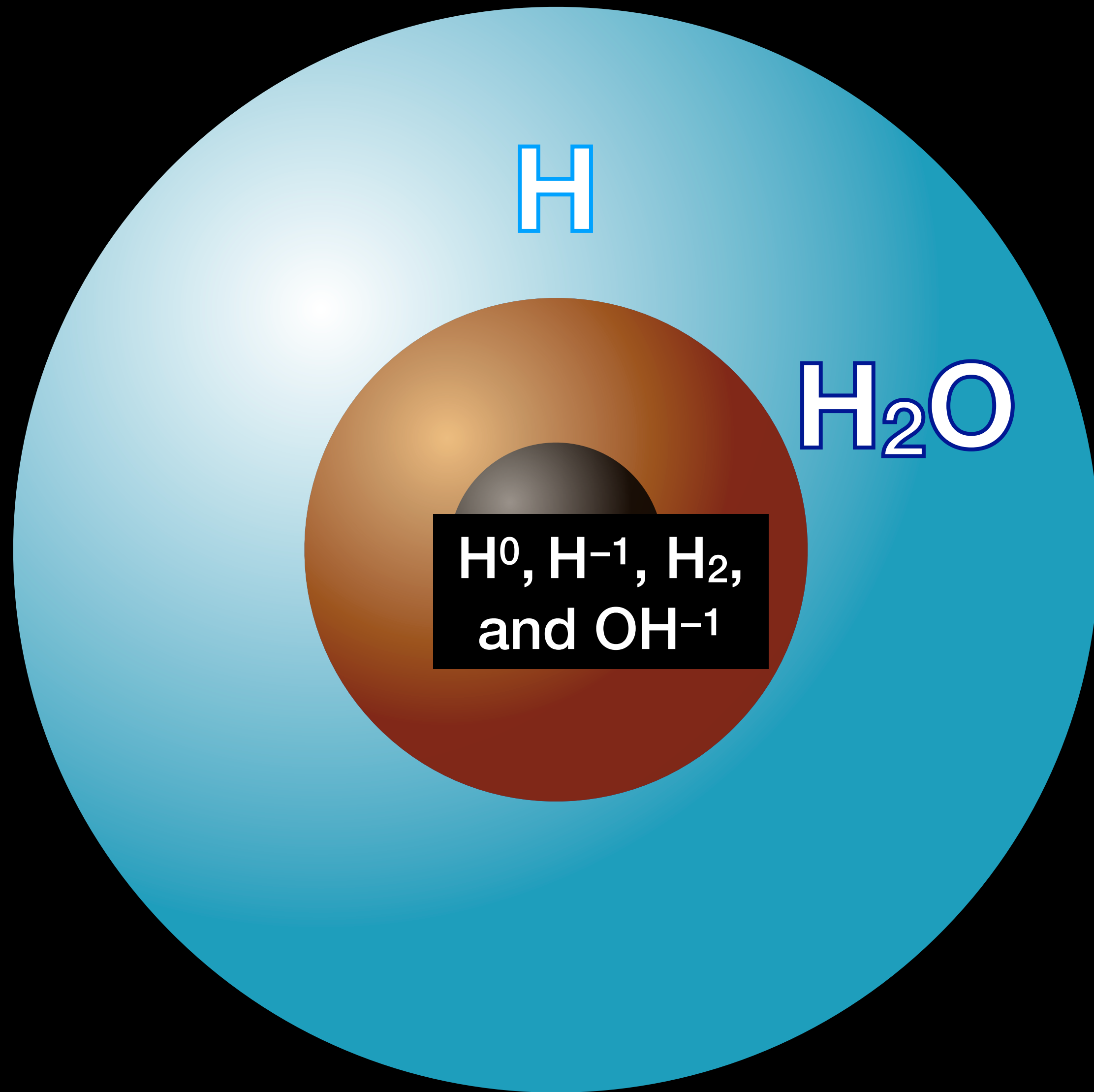
Mixing between Planetary Materials

H₂-H₂O: Lei et al. (2021) JPCL,
Soubiran and Militzer (2015)
ApJ

H₂O-MgO: Kim et al. (2021)
Nat. Astron.



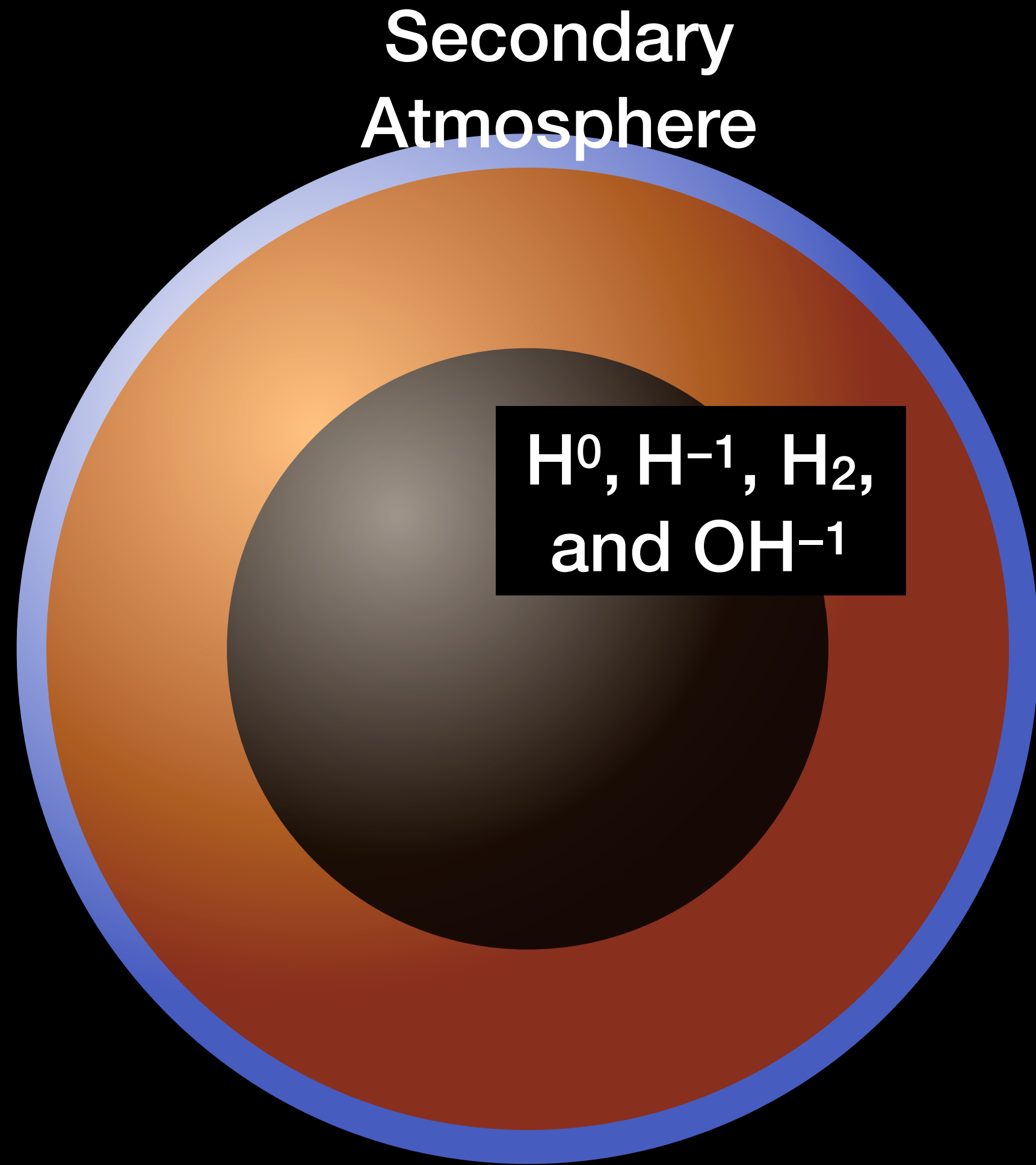
Sub-Neptune



Ingassing


- Ingassing through mixing
- Dynamic processes

Outgassing



- Low H/H₂O solubility in solid silicates and metal.
- Impact of interior processes for the chemical composition of secondary atmosphere

Super-Earths

The image features two realistic depictions of the Earth, one on the left and one on the right, showing blue oceans, white clouds, and brownish-green continents. They are positioned behind a central, semi-transparent purple rectangular box. The background is a solid black color.

Super-Earths may not be simple scale up versions of Earth.

Summary

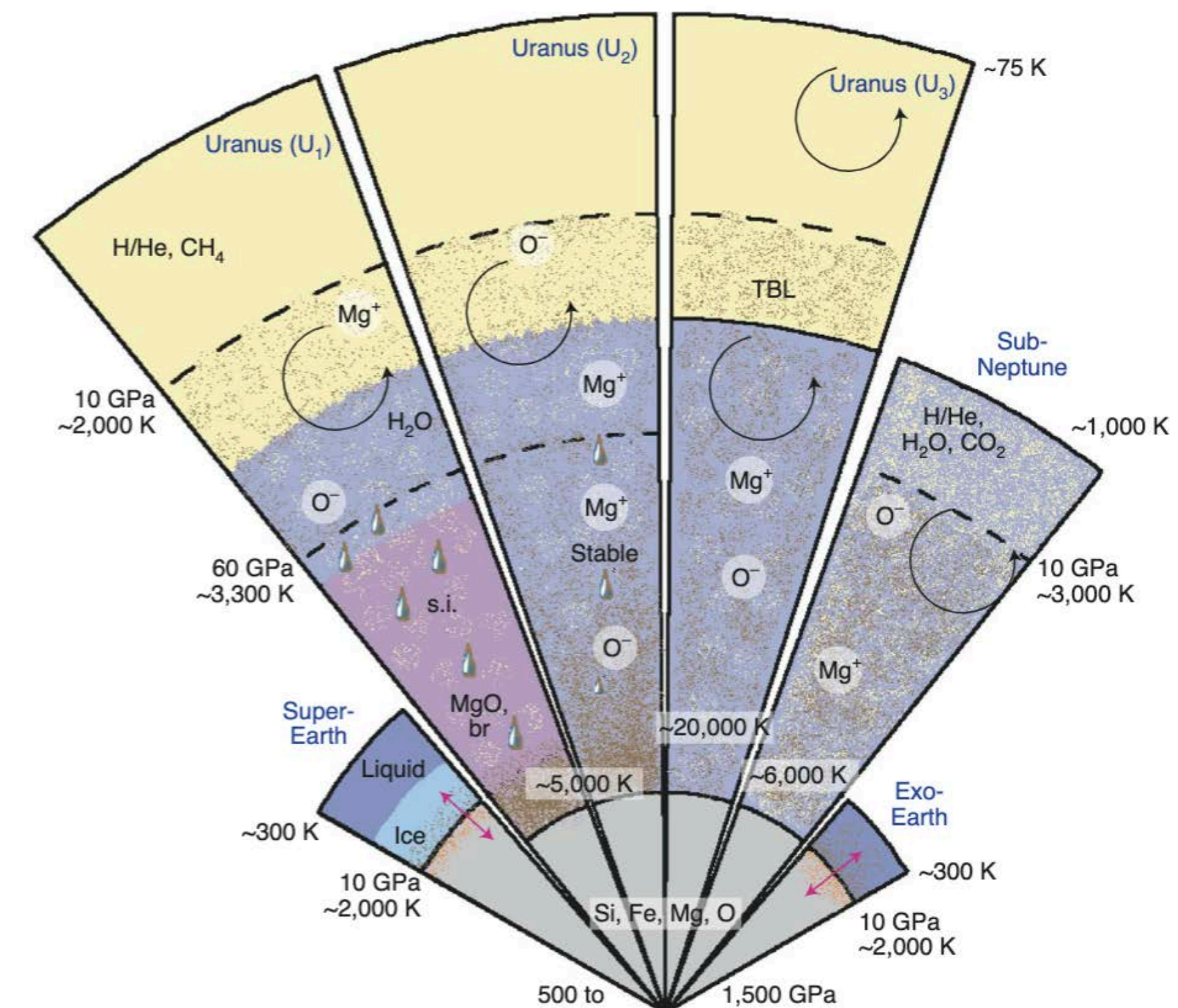
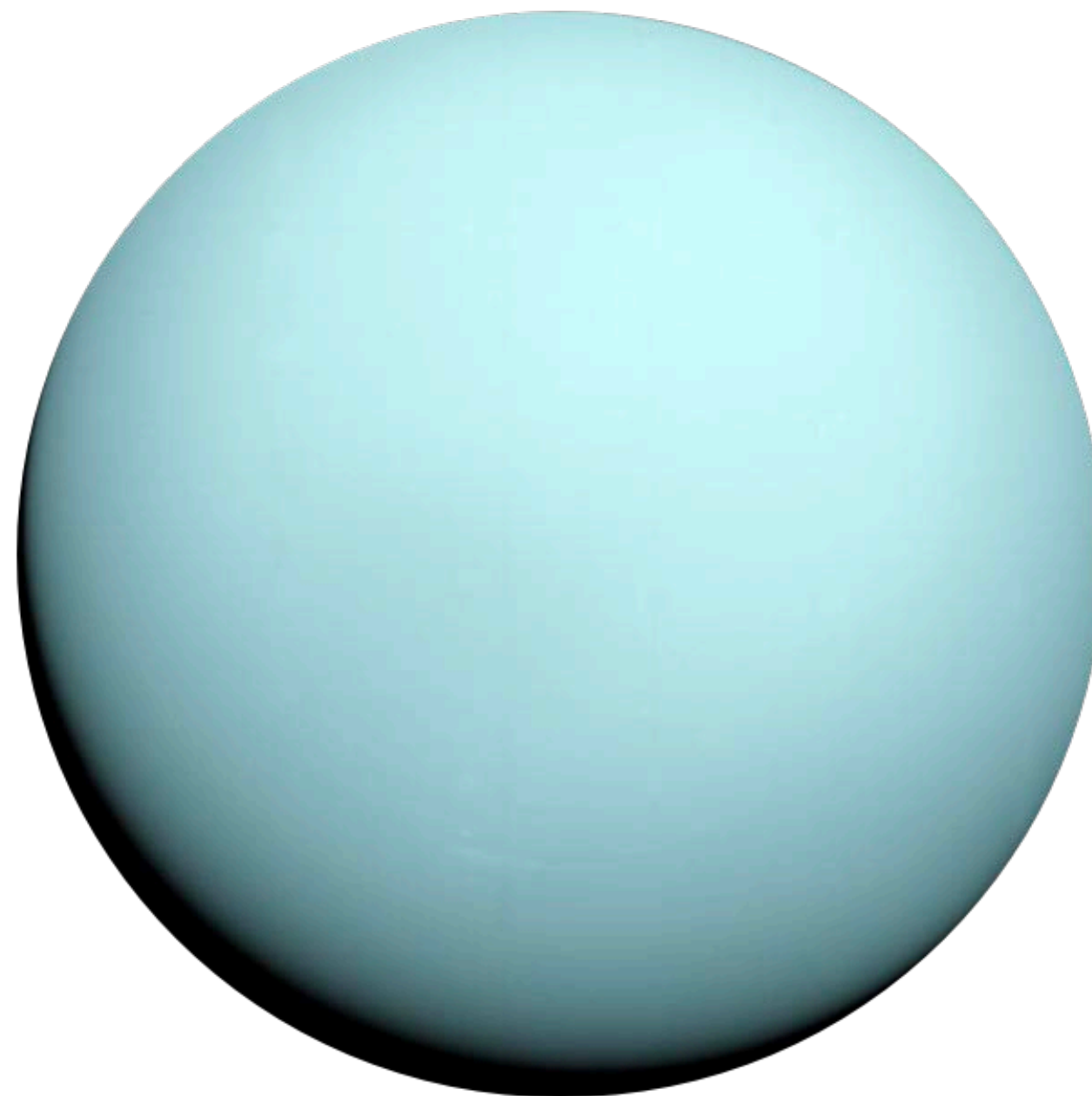
- Hydrogen-silicate reaction can produce water, converting a dry hydrogen rich planet to a wet water rich planet.
- Mixing-demixing could play an important role in the formation of secondary atmosphere of super-Earths converted from sub-Neptunes
- High-pressure chemistry will play a key role in advancing our understanding on exoplanets' atmospheres.

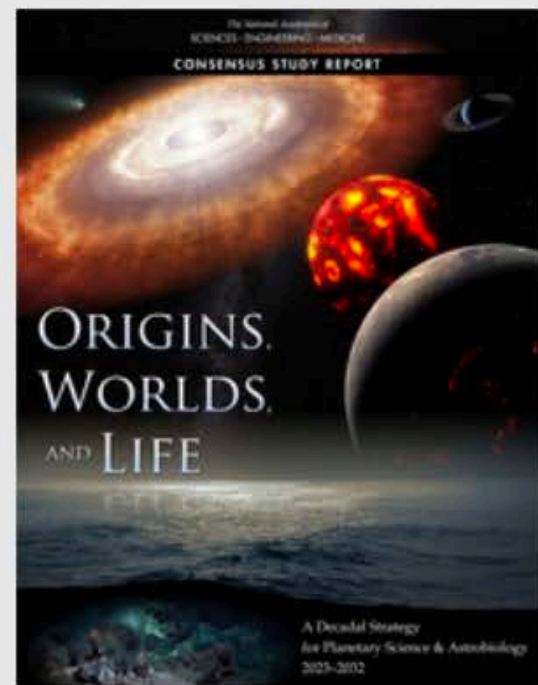
PLANETARY INTERIORS

Stardust in the deep interior of low-mass planets

The behaviour of minerals under high pressure affects the bulk properties of exoplanets and planets with rocky components, possibly influencing their observable radii. Obtaining a wide range of experimental data is key to understanding observations and informing planetary interior models.

Nadine Nettelmann



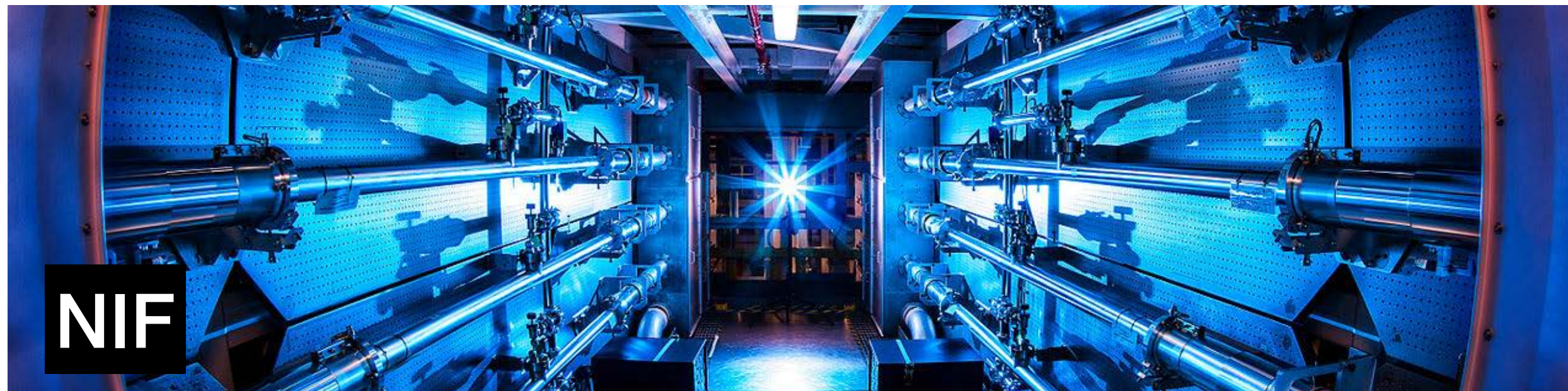


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DETAILS

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Facility for Open Research in
a Compressed Environment



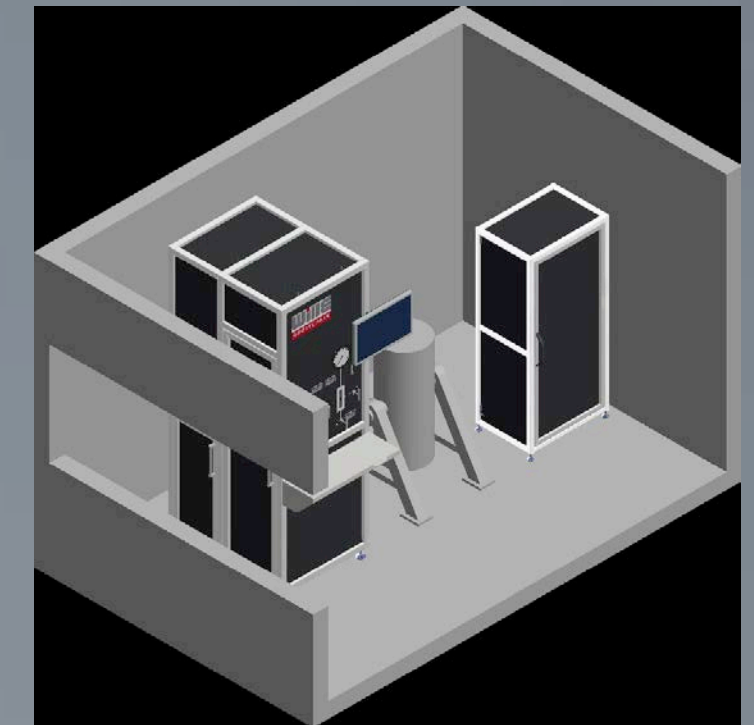
6000-t multi-anvil
press



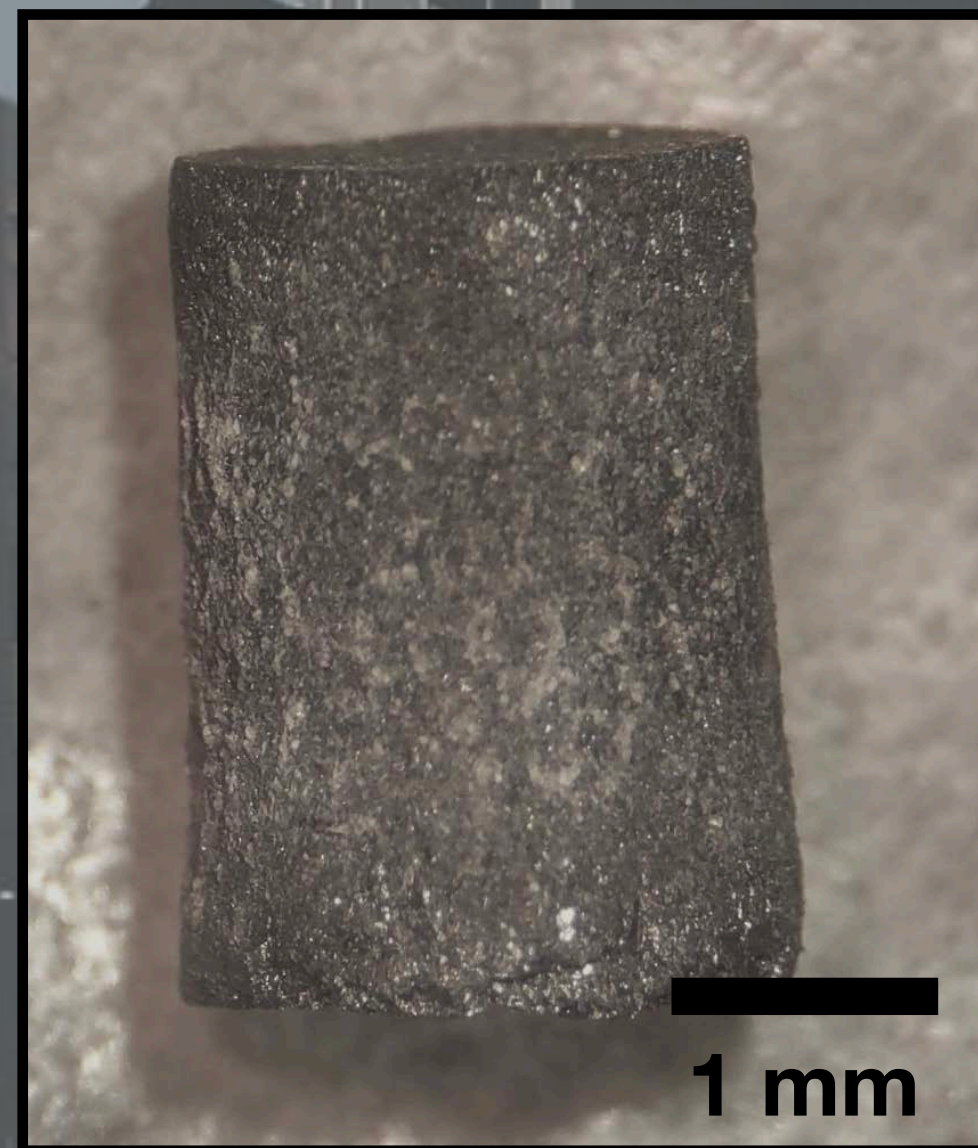
Ultrahigh pressure
cubic press



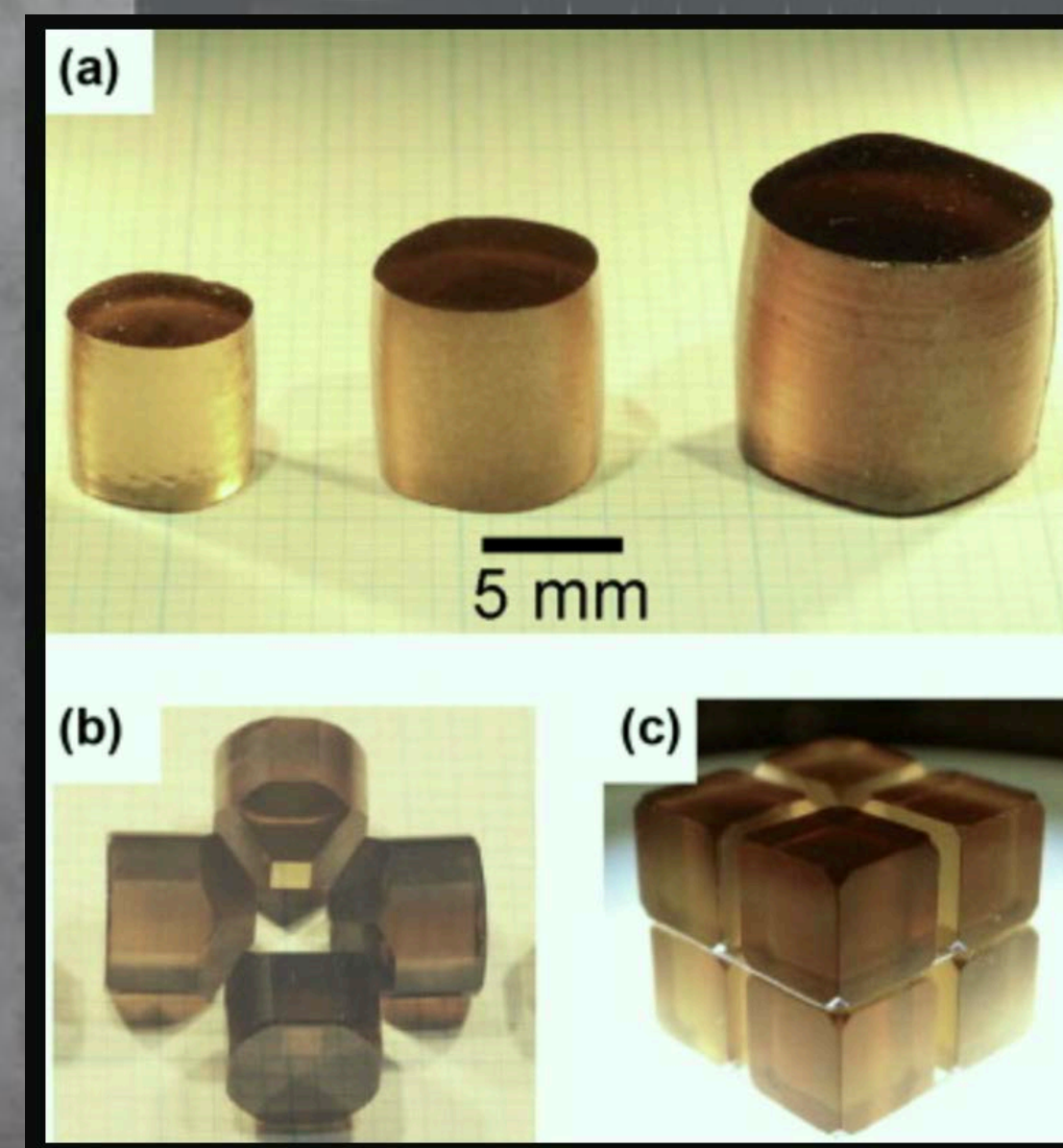
Torsional
apparatus



Gas pressure
vessel



Dense FeSiO_3
Rod



Nano-
polycrystalline
diamond
(Irifune et al., 2014)

