

# Nuclear Physics for and with the National Ignition Facility

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<https://nucleardata.berkeley.edu>

# The Bay Area Nuclear Data (BAND) Group

**NUCLEAR DATA GROUP**

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Our mission is to address the nuclear data needs of the applied and basic nuclear science and engineering community while training the next generation of nuclear scientists and engineers in the process

# Me in a nutshell

- Rutgers University Ph.D. (1994)
- LLNL Post-doc → Staff (1994-2016)
- Deputy Group Leader for Nuclear Diagnostics at NIF (2008-2013).
- Q-clearance holder (LLNL VSP)
- Joint Faculty Scientist (UC/LBNL) and Nuclear Data Group Leader
- Published 230+ papers in Nuclear Structure, Reactions, Plasma Physics
- Principal advisor to 12 graduate and 3 undergraduate students.
- Nuclear Science & Security Consortium Director for National Labs
- Nuclear Science Advisory Committee member (2021-2023)
  - Nuclear Data Subcommittee Chair (2022-2023)

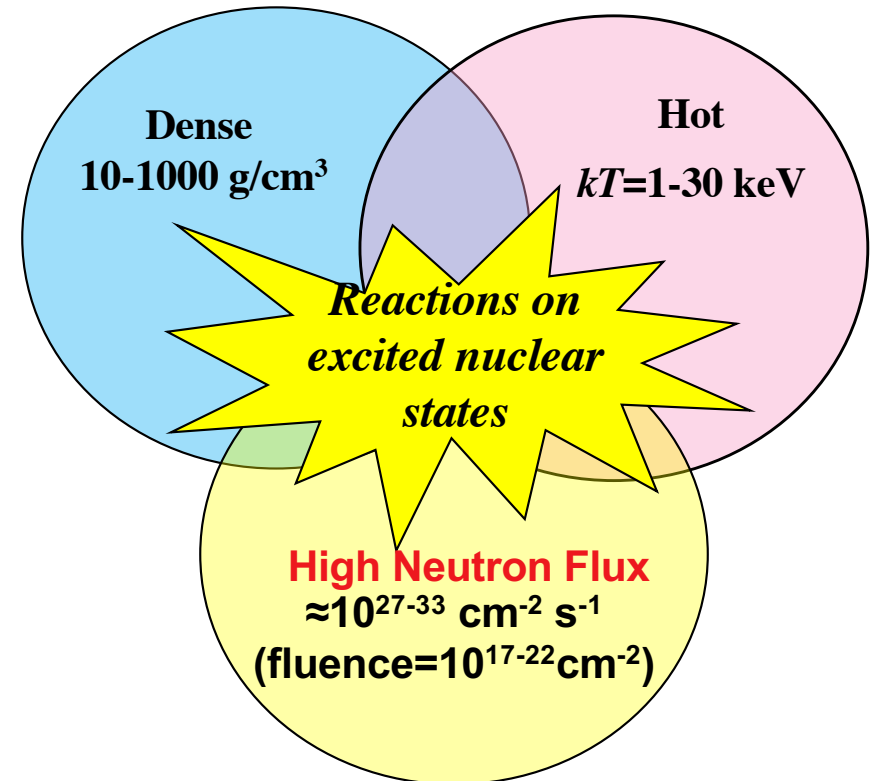
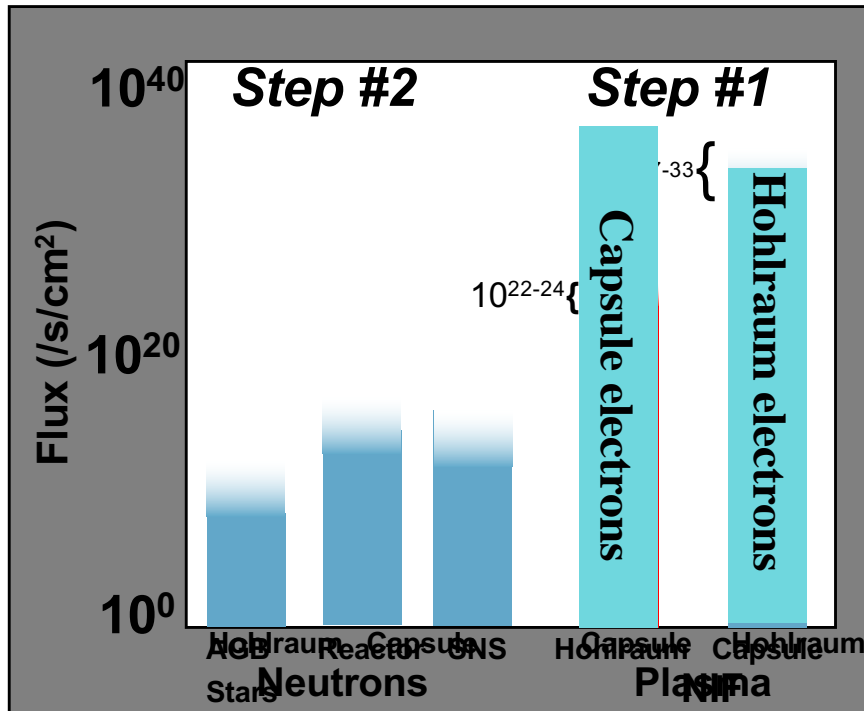


# Outline

- Introductions
- What NIF can do for Nuclear Physics
  - Nuclear reactions in HEDP environments
  - Historical example: The wheel experiments
  - Gedanken experiment #1: Using DD capsules to study neutron capture nucleosynthesis in an HEDP
  - Gedanken experiment #2: Using NIF+ARC to study fission in High Energy Density Plasmas
- What Berkeley can do for NIF & LLNL
  - The 88-Inch cyclotron high-intensity neutron source
  - The BELLA petawatt laser-driven HED facility
  - Example: New gamma diagnostic for  $E_\gamma < 3$  MeV

The high  $e$ ,  $\gamma$  and  $n$  flux in an ICF capsule allows us to explore the effects of a HEDP environment on nuclear reaction rates

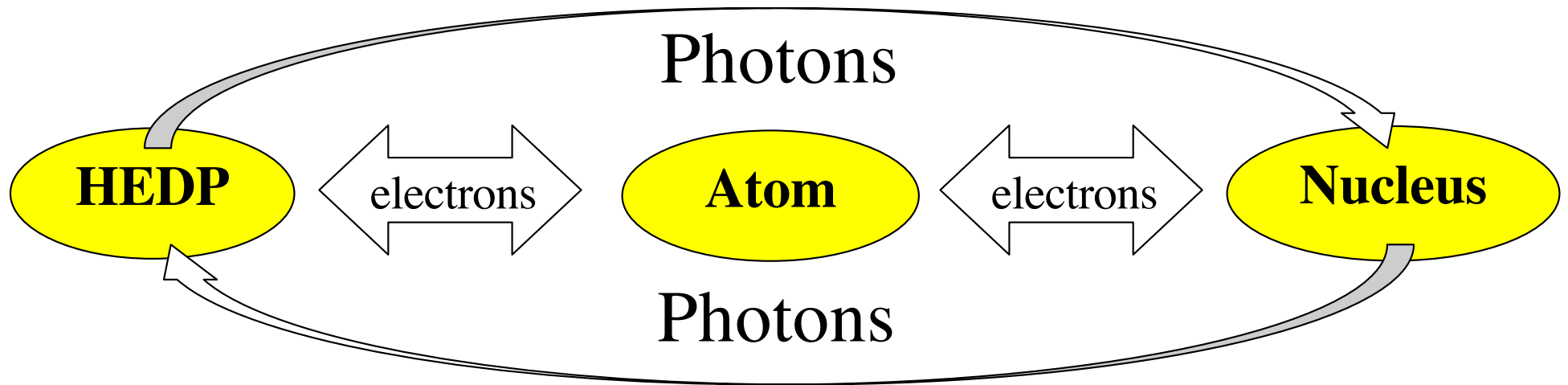
### NIF capsule/hohlraum



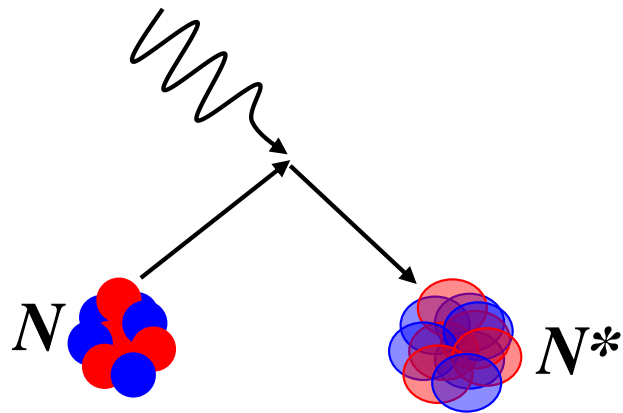
### Excited State Reaction Possibilities

*Nuclei are excited by and HEDP and then capture a neutron*

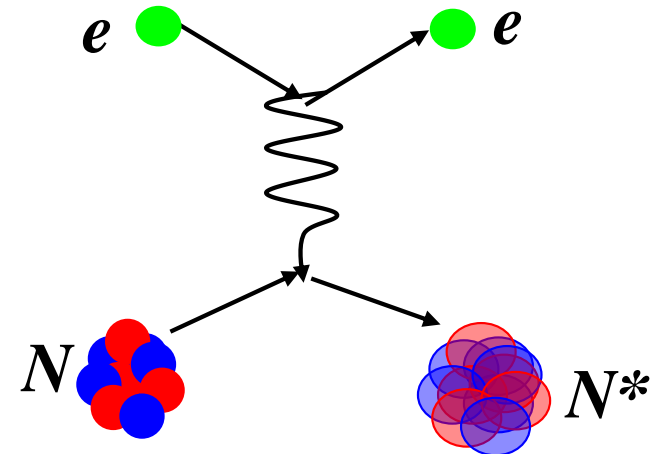
# Nuclear-plasma interactions are assumed to cause thermal population of low-lying nuclear states in HED plasmas



**Photo-absorption**  
Time Reverse:  $\gamma$ -ray decay



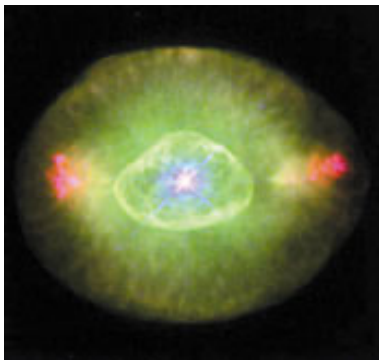
**Atomic-nuclear (electron) interactions**  
NEEC, NEET, IES\* (Time Reverse: IC, BIC)



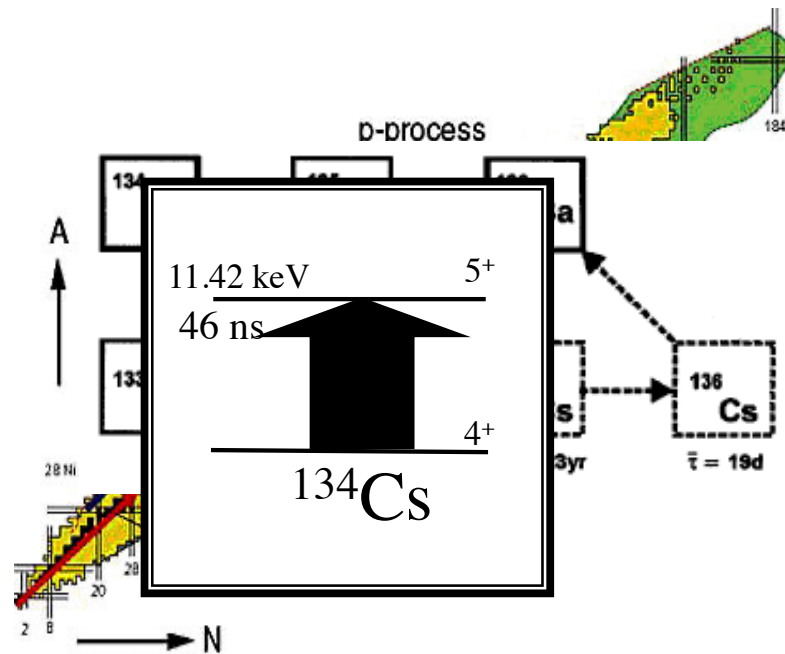
Electron-nuclear rates are likely to be the most important in stars

# Nearly half of the elements are made via neutron capture *in a stellar plasma*

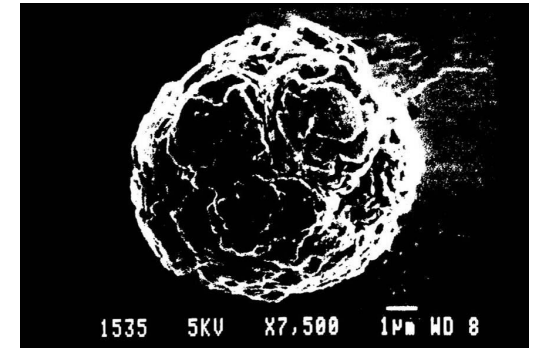
AGB stars →  
disperse elements



$$k_B T \approx 10-30 \text{ keV}$$



Pre-solar grains



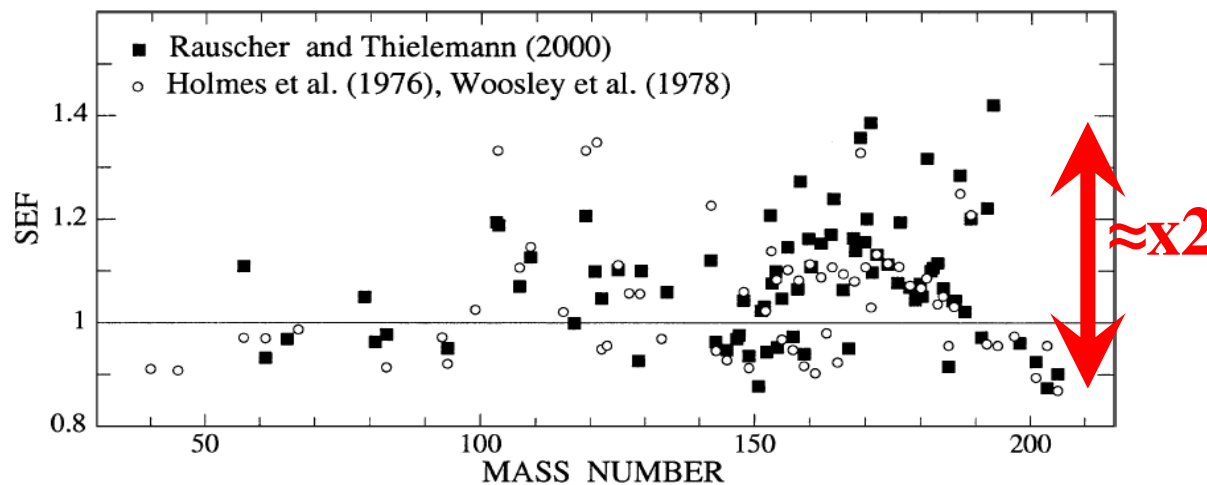
**(n,γ) cross section on s-process branch point nuclei allows a “forensic” study of the conditions in a stellar interior**

# The excitation of these low-lying levels leads to a change in the neutron capture cross sections

S-process (n,γ) cross sections due to excited states\*

$$\sigma_{stellar}^{(n,\gamma)} = \sum_i (2J_i + 1) e^{\frac{-E_i}{kT}} \sigma_i^{(n,\gamma)}$$

$$SEF = \frac{\sigma_{stellar}^{(n,\gamma)}}{\sigma_{terrestrial}^{(n,\gamma)}}$$

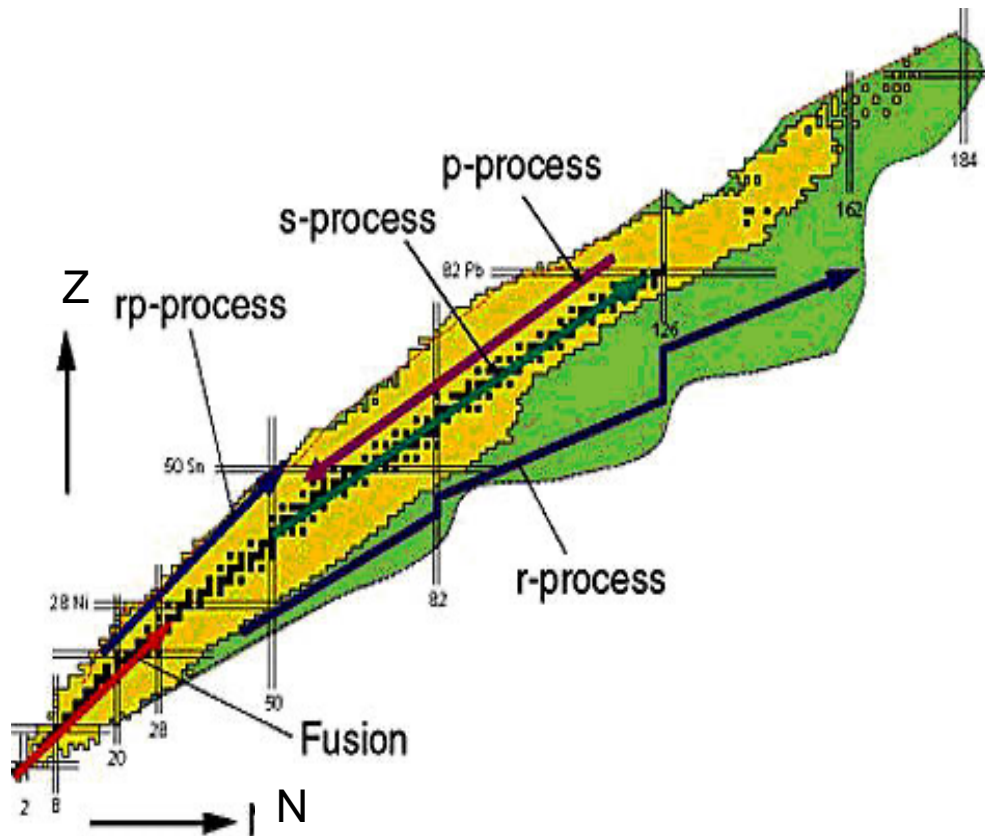


Branch Point	Gnd State J <sup>π</sup>	1 <sup>st</sup> Exc. State E <sub>x</sub> (keV)	1 <sup>st</sup> Exc. State J <sup>π</sup>
<sup>79</sup> Se	7/2 <sup>+</sup>	95.77	1/2 <sup>-</sup>
<sup>85</sup> Kr	9/2 <sup>+</sup>	304.871	1/2 <sup>-</sup>
<sup>147</sup> Pm	7/2 <sup>+</sup>	91.1	5/2 <sup>+</sup>
<sup>151</sup> Sm	5/2 <sup>-</sup>	4.821	3/2 <sup>-</sup>
<sup>163</sup> Ho	7/2 <sup>-</sup>	100.03	9/2 <sup>-</sup>
<sup>170</sup> Tm	1 <sup>-</sup>	38.7139	2 <sup>-</sup>
<b><sup>171</sup>Tm</b>	<b>1/2<sup>+</sup></b>	<b>5.0361</b>	<b>3/2<sup>+</sup></b>
<sup>179</sup> Ta	7/2 <sup>+</sup>	30.7	9/2 <sup>+</sup>
<sup>204</sup> Tl	2 <sup>-</sup>	414.1	4 <sup>-</sup>
<sup>205</sup> Pb	5/2 <sup>-</sup>	703.3	7/2 <sup>-</sup>
<sup>185</sup> W	3/2 <sup>-</sup>	23.547	1/2 <sup>-</sup>

A neutron-rich HED plasma like NIF is the *only* place where (n,γ) might be measured on a combination of ground & excited states



# Using NIF to measure $(n,\gamma)$ on s-process branch point nuclei was the topic of a 2010 workshop



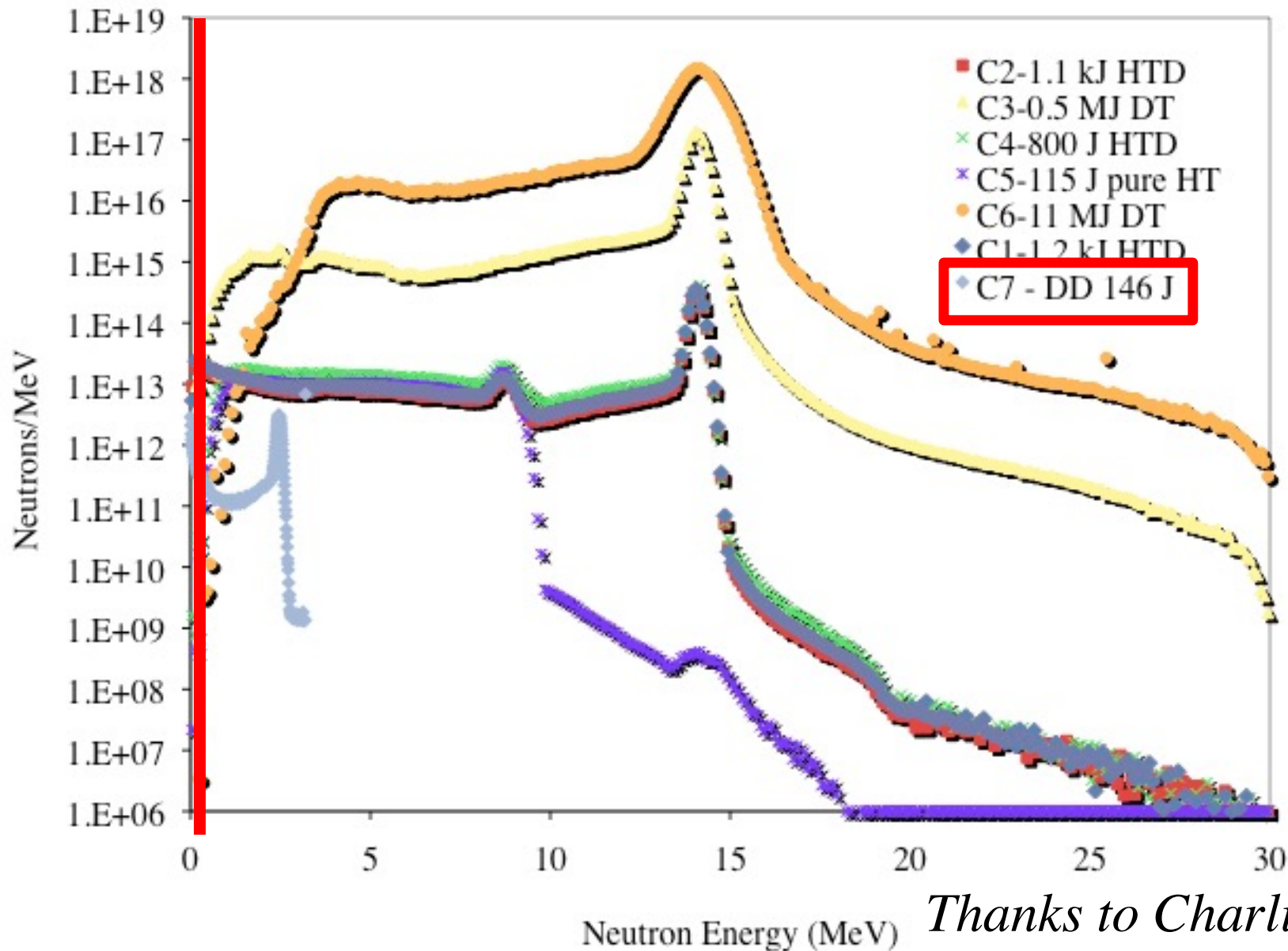
- Held at LBNL, March 2010 (!)
- 35 attendees from 8 institutions
  - LLNL, LBNL, LANL, Ohio Univ., Colorado-Mines, GSI-Darmstadt, Notre Dame, CEA-DAM
- Talks were presented on:
  - NIF diagnostics
  - s-process nucleosynthesis
  - Nuclear-plasma interactions

*This program would require DD fuel, target implantation and either solid radchem or Gamma spectroscopy*

# How would you measure an astrophysical (n, $\gamma$ ) cross section at NIF?

1. Create the correct environment (neutrons, T,  $\rho$ )
  - Fuel load and moderation environment
2. Get the material into the capsule
  - Ion-implantation
3. Measure target areal density
  - Energy resolved X-ray imaging
4. Measure the number of reactions and the neutron spectrum
  - Solid Debris Collection
  - Prompt  $\gamma$ -ray detection using Gas Cerenkov Detectors

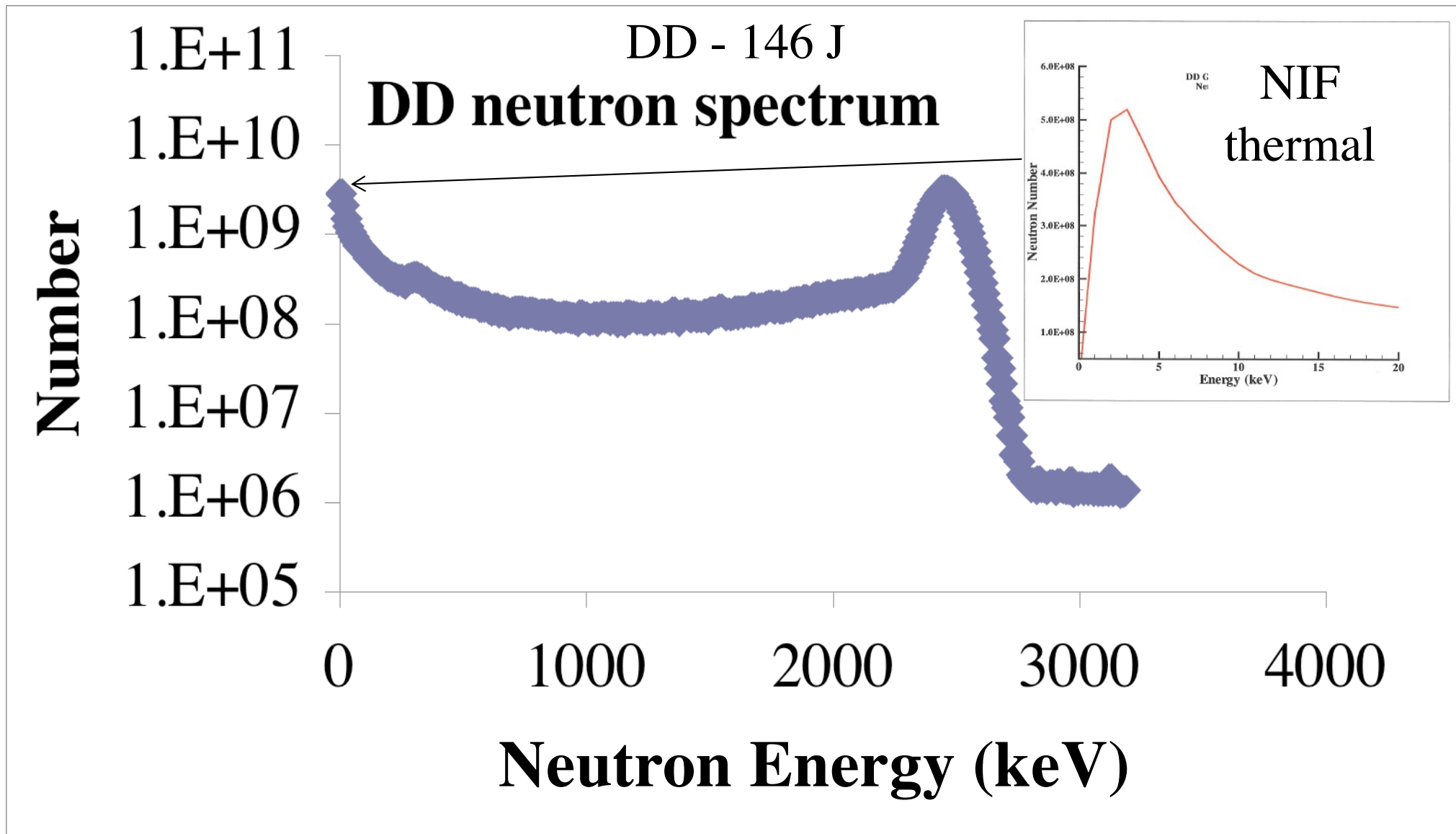
This type of measurement would require pure deuterium fuel to create the correct neutron spectrum



*Thanks to Charlie Cerjan!*

(Modeling courtesy of C. Cerjan)


This type of measurement would require pure deuterium fuel to create the correct neutron spectrum



(Modeling courtesy of C. Cerjan)

# The neutron capture reaction products could be measured either by prompt gamma signal or solid/gaseous radchem

$Ta_2O_5$



Tantalum oxide ( $Ta_2O_5$ ) aerogels

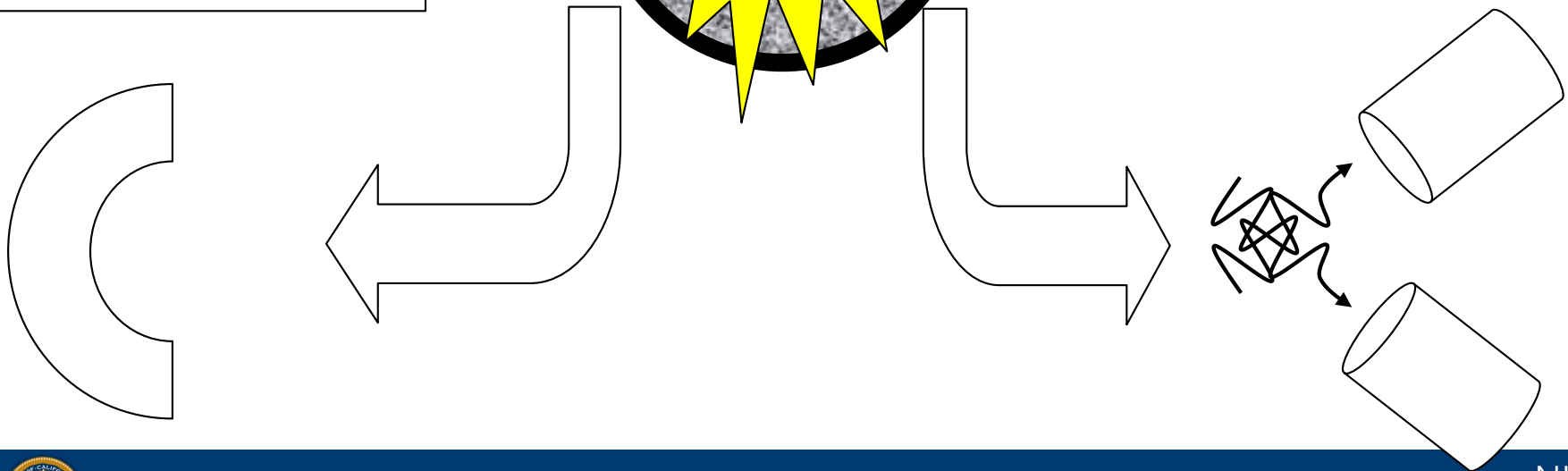
Application: HED and emission experiments

140 mg/cc

( $\leq 10^{16}$  atoms)  
*Hund et al.,*

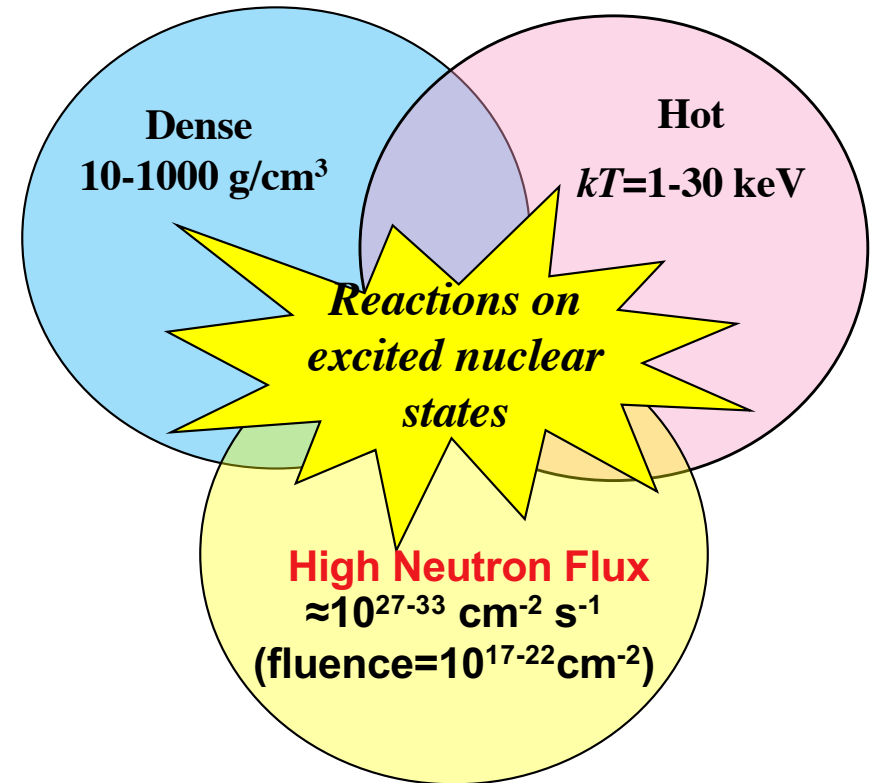
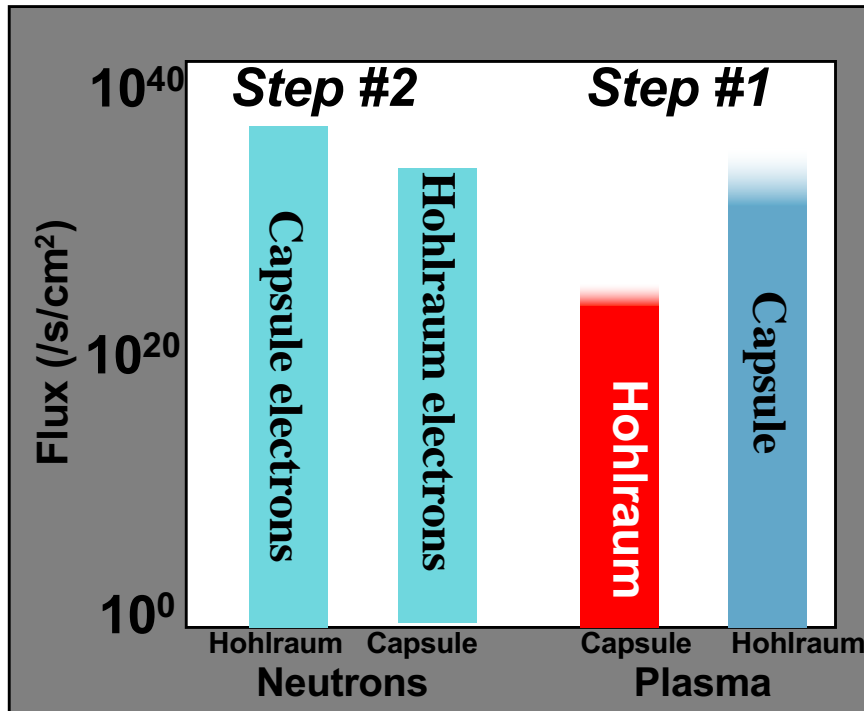
Collect “debris”  
and count-  
Tough if stable product  
(Shaughnessy,  
Grim, Greife...)

Try to measure the  
prompt  $\gamma$ -rays  
following capture



The high  $e$ ,  $\gamma$  and  $n$  flux in an ICF capsule allows us to explore the effects of a HEDP environment on nuclear reaction rates

### NIF capsule/hohlraum

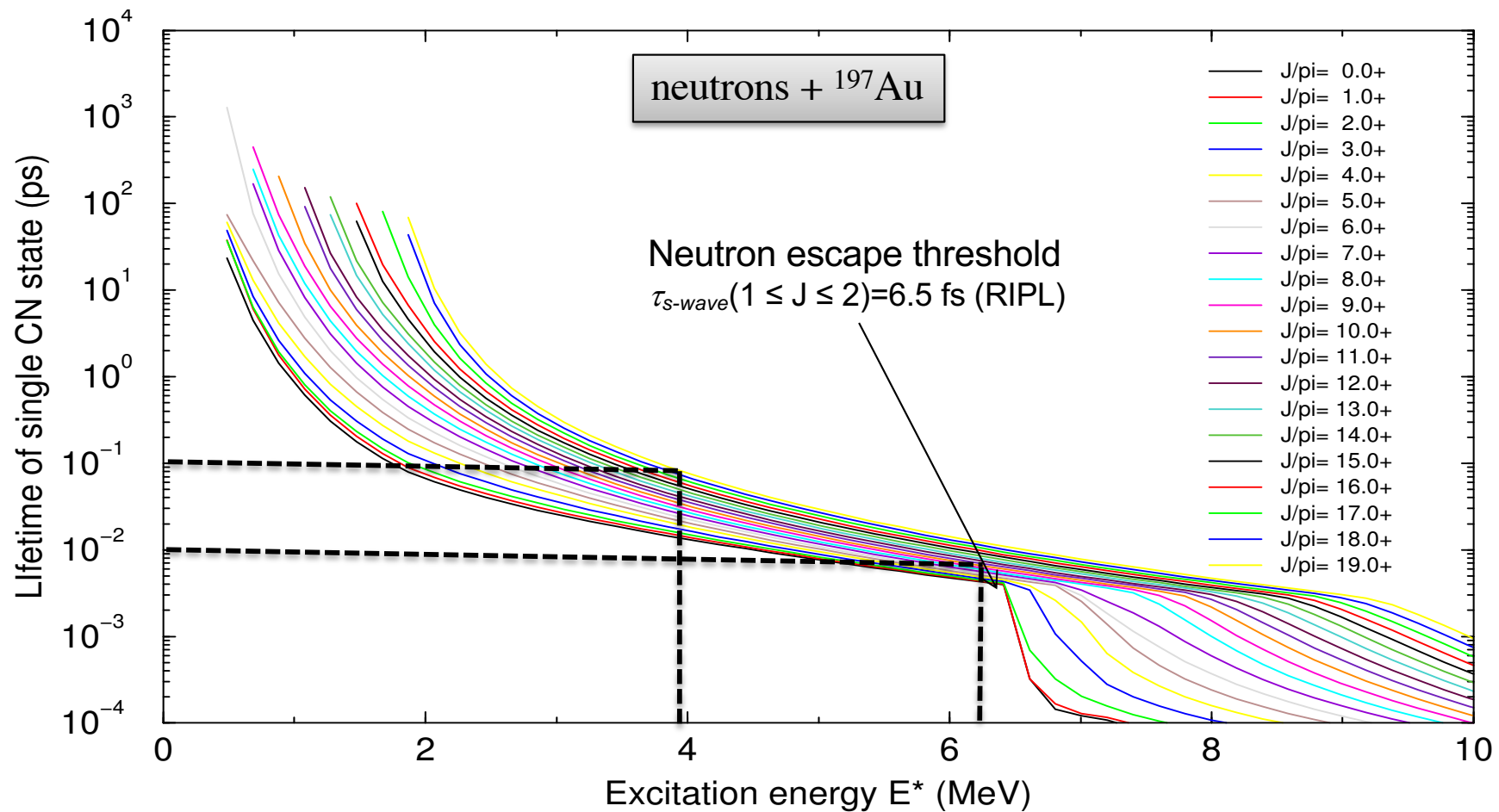


### Excited State Reaction Possibilities

*Nuclei are excited by an HEDP and then capture a neutron*

*Nuclei absorb a neutron and interact with an HEDP before particle emission*

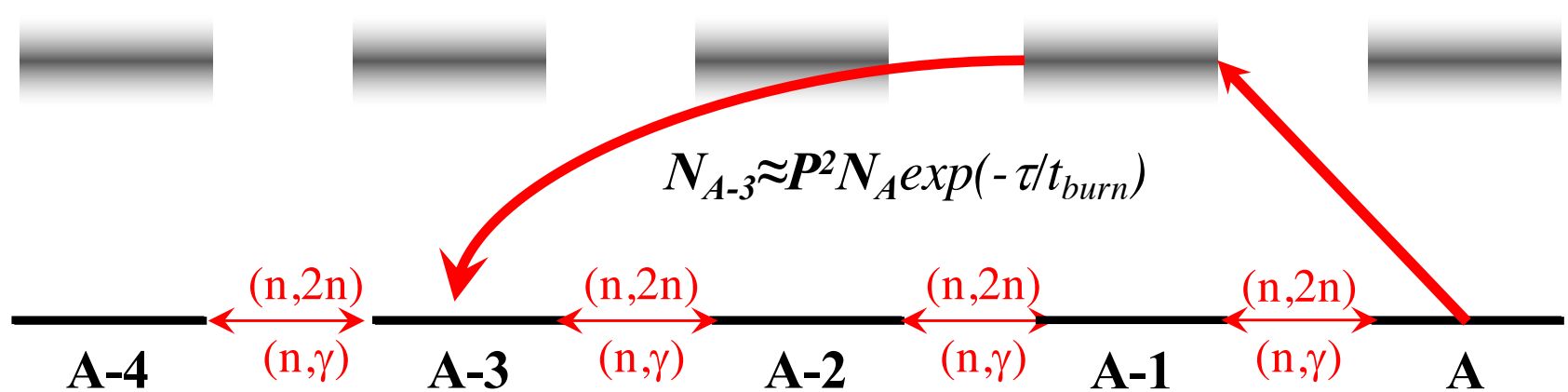
# The lifetimes of highly-excited nuclear states in heavy nuclei is comparable to the NIF burn time



What would happen if an excited nucleus were hit by a second neutron prior to decaying to the ground state?

# The large flux of 14 MeV neutrons in the NIF would drive multiple (n,2n) reactions, rather than (n,γ)

$E_x \approx 3-5$   
MeV



$$N_{A-4} = P^4 N_A \quad N_{A-3} = P^3 N_A \quad N_{A-2} = P^2 N_A \quad N_{A-1} = P N_A \quad N_A$$

- The probability that a nucleus  $A$  will be converted via (n,2n) to a nucleus  $A-1$  is given by:

$$P_{reaction} = \int_0^{\tau_{burn}} \int_0^{14} \phi_n(E_n) \times \sigma_n(E_n) \times e^{-t/\tau_{state}} dE_n dt$$

$$P_{reaction} \approx \frac{10^{17} \text{ neutrons/MeV}}{4\pi(30 \mu\text{m})^2 \times 10^{-11} \text{ s}} \times 10^{-24} \text{ cm}^2 \times 10^{-13} \text{ s} \cdot \text{MeV}$$

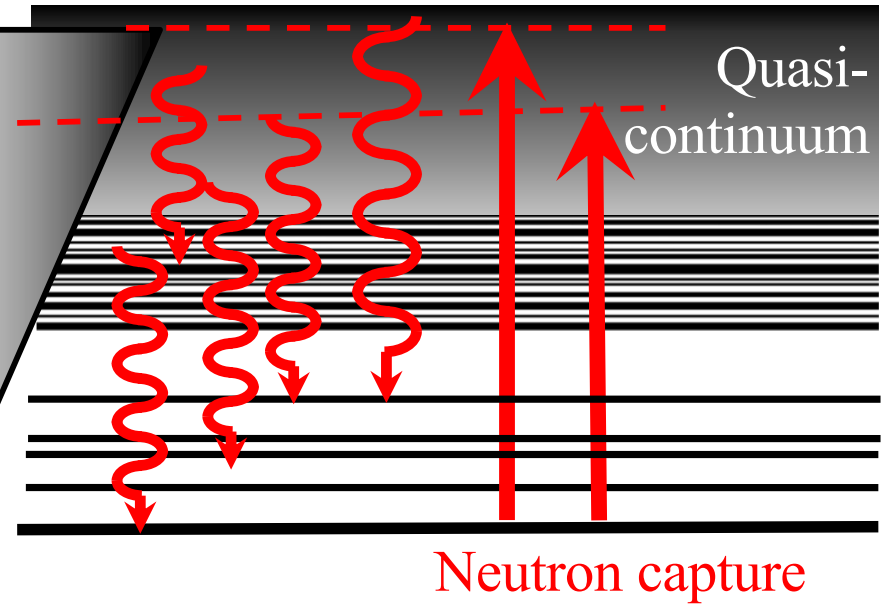
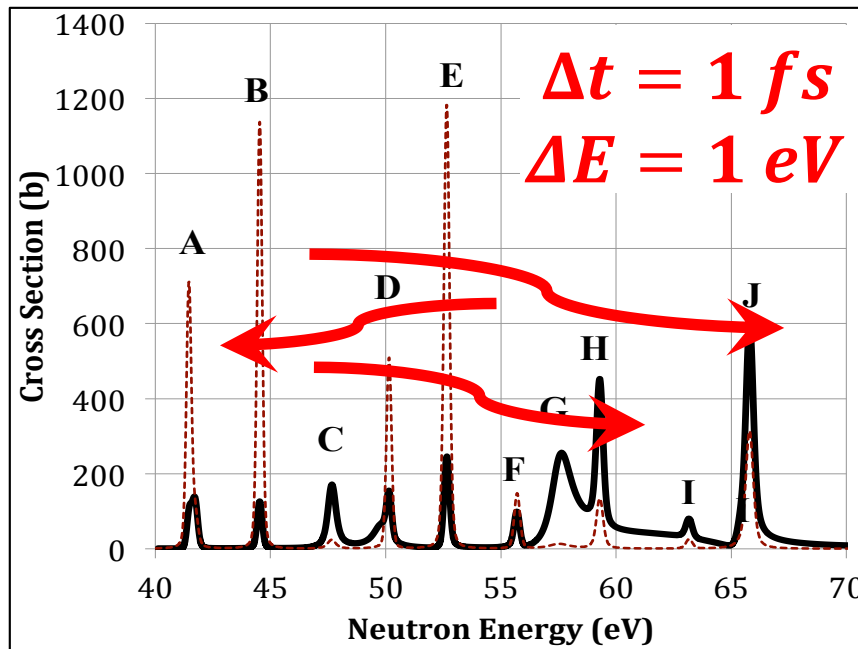
$\Delta t \approx 1\% t_{burn}$

$$P_{reaction} \approx \frac{10^{17} \text{ neutrons/MeV}}{10^{-4} \text{ cm}^2 \times 10^{-11} \text{ s}} \times 10^{-24} \text{ cm}^2 \times 10^{-13}$$

The  $\frac{N_{A-3}}{N_{A-1}}$  ratio would produce a solid radchem observable

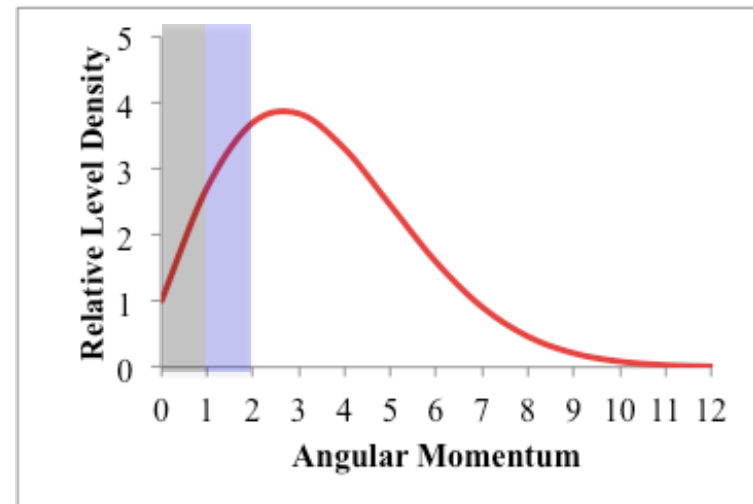


At multi-MeV energies the high nuclear level density makes the nucleus a much better absorber of energy from the HEDP

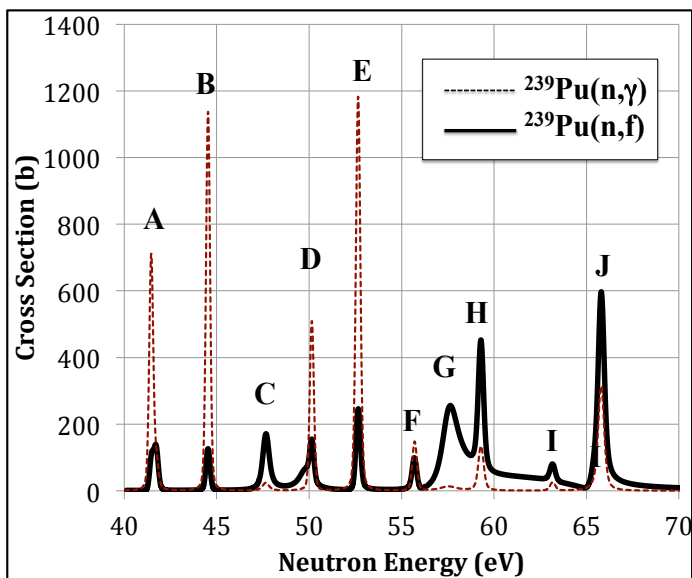


The large density of states would make these nuclear-plasma rate  $\approx 10^{14}$  Hz!

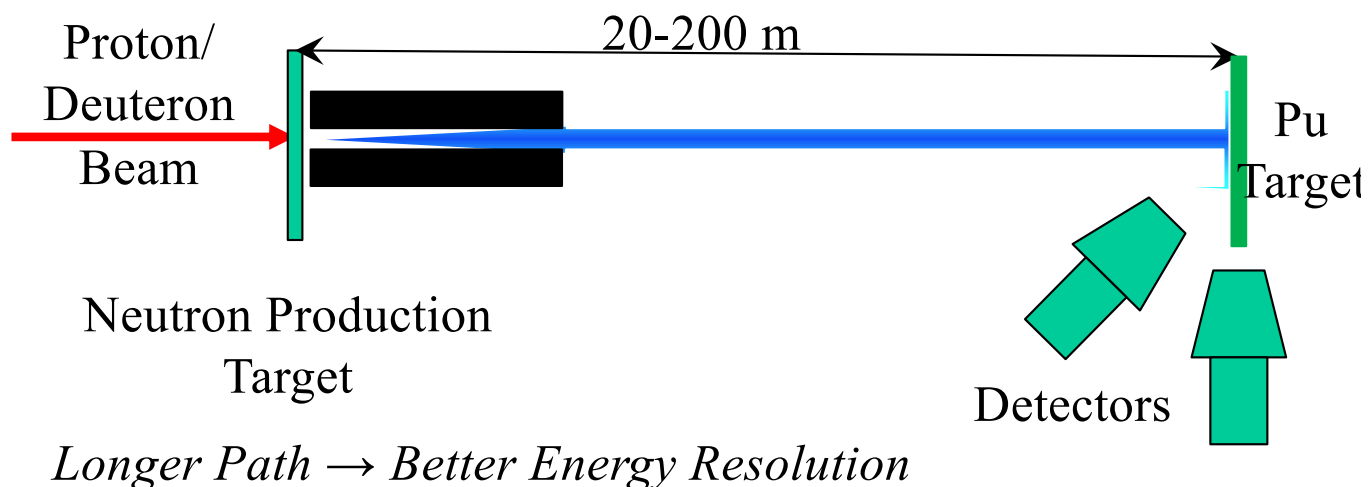
Each photon (real or virtual) will shift  $J^\pi$  toward the most likely value



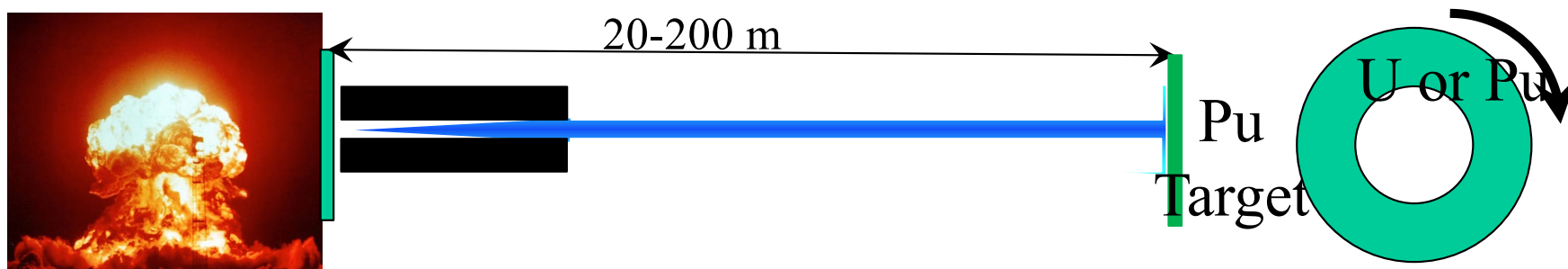
# Changing excited states could have a big influence on fission as well - The (in)famous “wheel” experiments



“Time-of-Flight” (TOF) technique – Multi-Day Experiment



Best TOF – Experiment – LOTS of neutrons in one pulse viewed from REALLY far away



# Yes. We did that (not me personally)

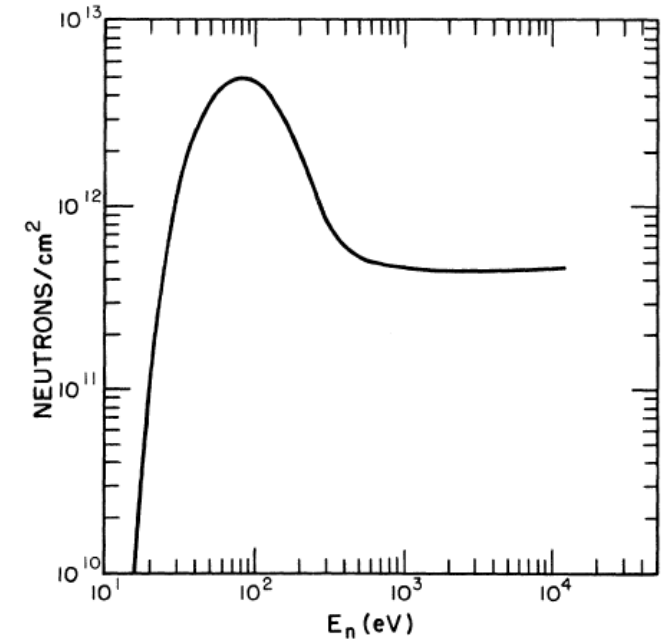
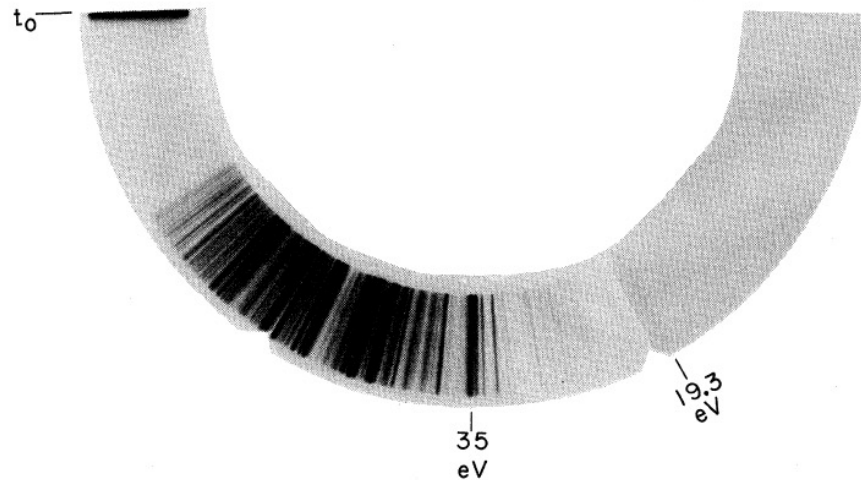
PHYSICAL REVIEW C

VOLUME 2, NUMBER 2

AUGUST 1970

## Symmetry of Neutron-Induced $^{235}\text{U}$ Fission at Individual Resonances. III\*

G. A. Cowan, B. P. Bayhurst, R. J. Prestwood, J. S. Gilmore, and G. W. Knobeloch  
Los Alamos Scientific Laboratory, University of California, Los Alamos, New Mexico 87544  
(Received 13 February 1970)



The symmetry of neutron-induced fission in  $^{235}\text{U}$  has been measured at a large number of epithermal levels with a time-of-flight energy-resolved beam from an underground nuclear explosion. The energy resolution was improved over previous similar "wheel" experiments. The symmetry of fission, as measured radiochemically by the ratio of  $^{115}\text{Cd}$  to  $^{99}\text{Mo}$ , shows a bimodal distribution. If the results are averaged separately for the two apparent groups, the first group (I) has an average  $^{115}\text{Cd}/^{99}\text{Mo}$  ratio which is 0.593 times the thermal value; the second-group (II) average is 1.11 times the thermal value. The average fission width in I is one-half the average fission width in II. The numbers of levels assigned to I and II are 24 and 14, respectively. The energies of levels assigned to I are 19.30, 21.07, 22.94, 23.63, 26.44, 27.81, 30.91, 32.10, 33.55, 34.85, 40.51, 41.91, 43.41, 44.04, 45.79, 46.93, 48.06, 48.82, 51.37, 56.58, 57.83, 58.15, 60.25, and 61.10 eV; assignments to II are at 24.23, 25.65, 28.35, 34.39, 35.21, 35.75, 36.64, 38.42, 39.44, 44.75, 49.51, 52.27, 53.56, and 58.68 eV. From analogy with conclusions based on a similar set of observations in the fission of  $^{239}\text{Pu}$  and on arguments derived from fission-channel theory, it is hypothesized that  $J=4$  in I and 3 in II.

*If the results are averaged separately for the two apparent groups, the first group (I) has an average  $^{115}\text{Cd}/^{99}\text{Mo}$  ratio which is 0.593 times the thermal value; the second-group (II) average is 1.11 times the thermal value.*

NIF+ARC could be used to study the effects of an HEDP on fission yields in a controlled manner

Use ARC to put a small uranium sample into an HEDP state just as the 14 MeV neutron pulse arrives



Timing is everything 😊

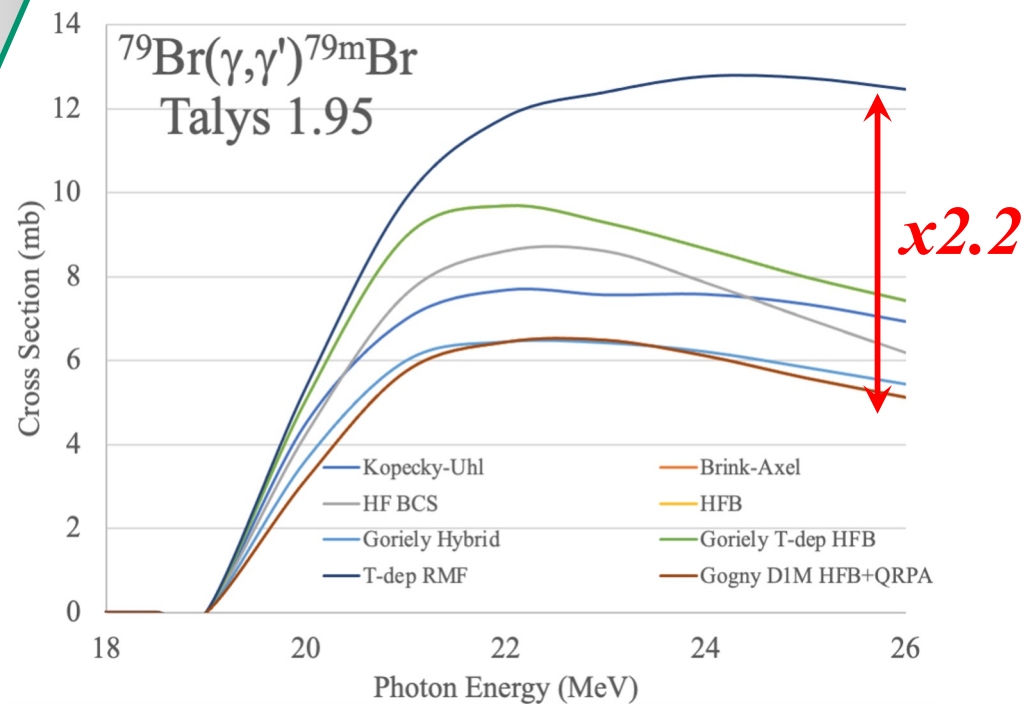
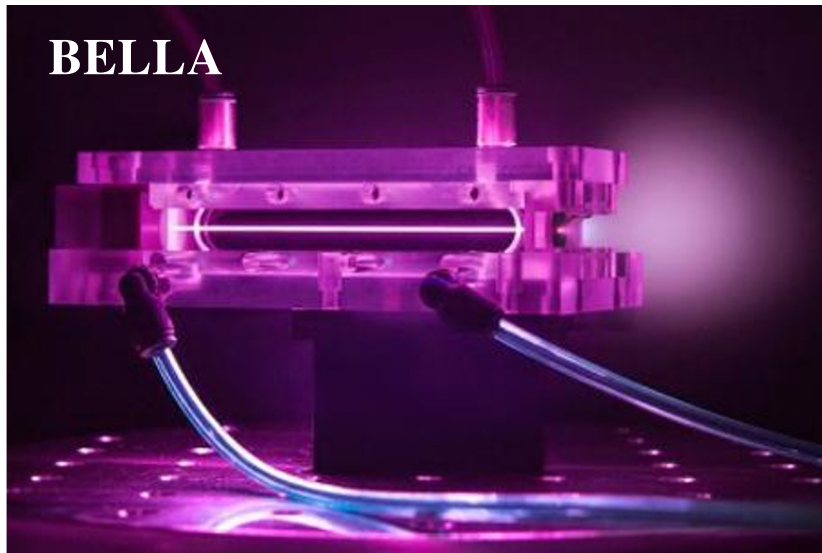
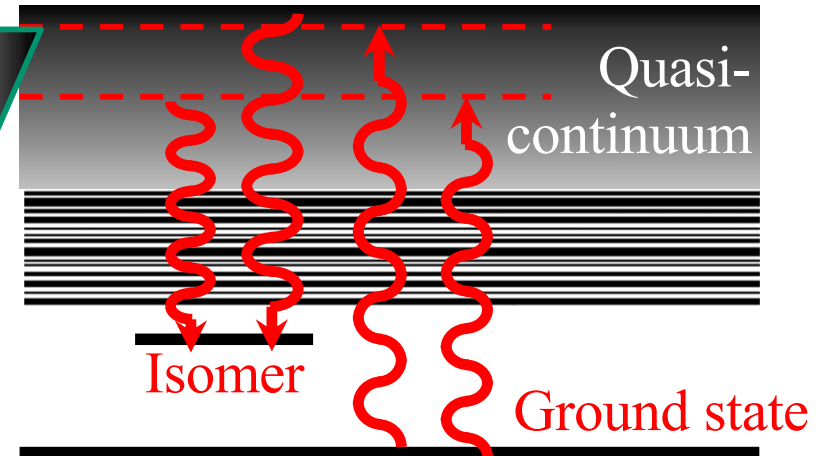
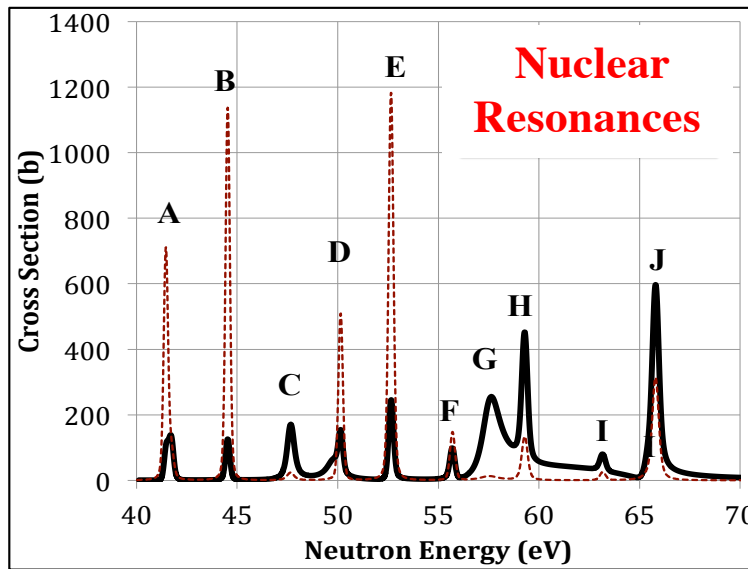
NIF+ARC could be used to study the effects of an HEDP on fission yields in a controlled manner

Repeat, but delay the ARC pulse to arrive well-after the 14 MeV neutrons



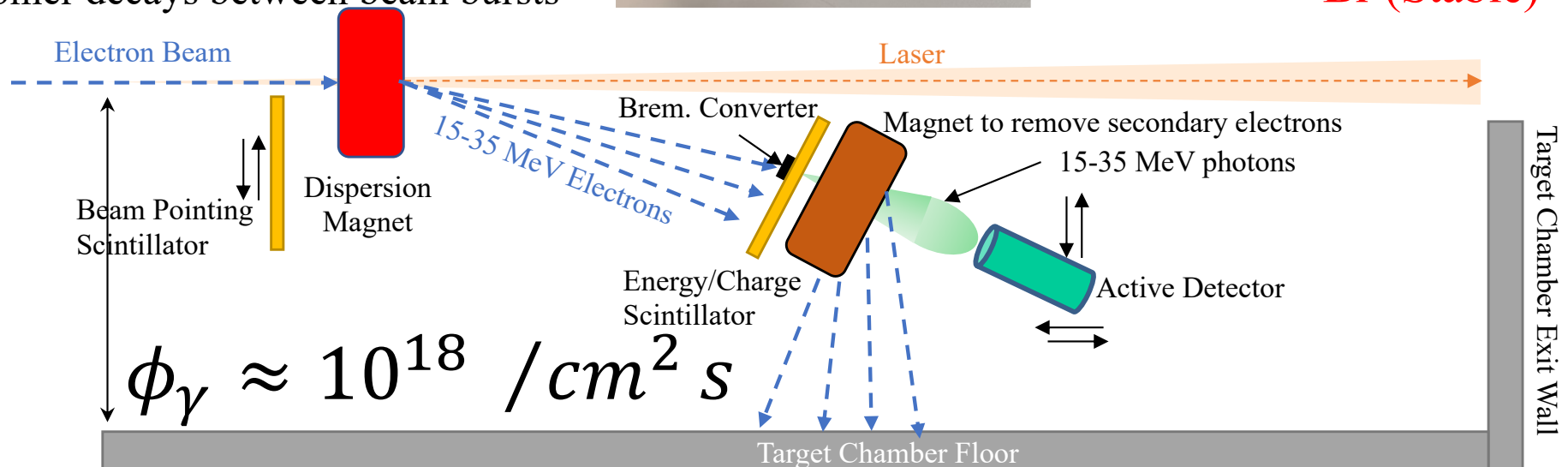
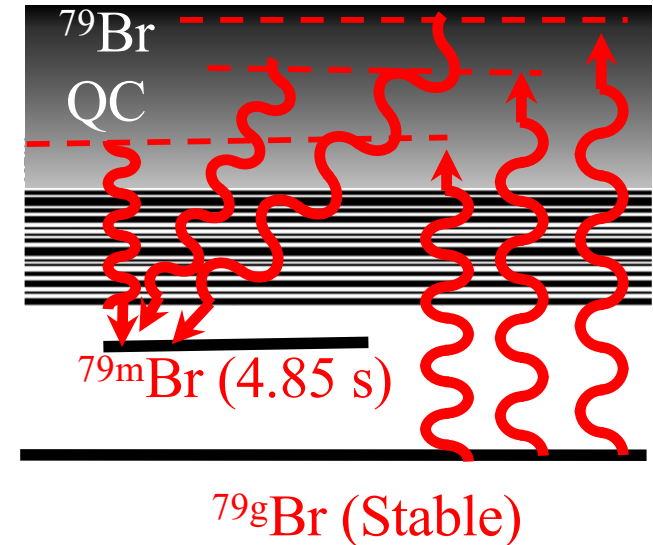
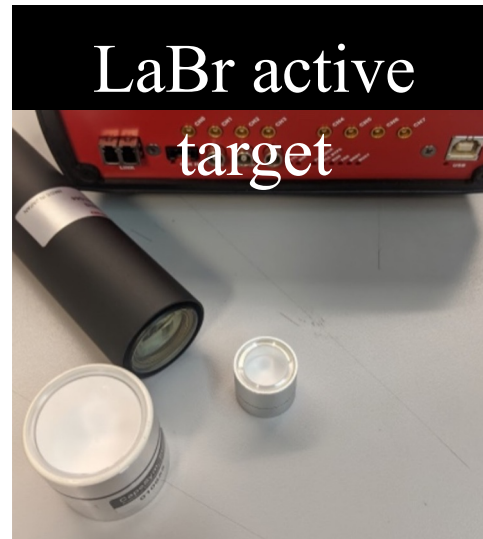
Controlled experiments are always best

# Moving to Berkeley: The BELLA facility has provided a capability to study Nuclear Plasma Interactions in HEDPs



# Part I: $^{79}\text{Br}(\gamma, x\gamma')^{79m}\text{Br}$ via Bremsstrahlung

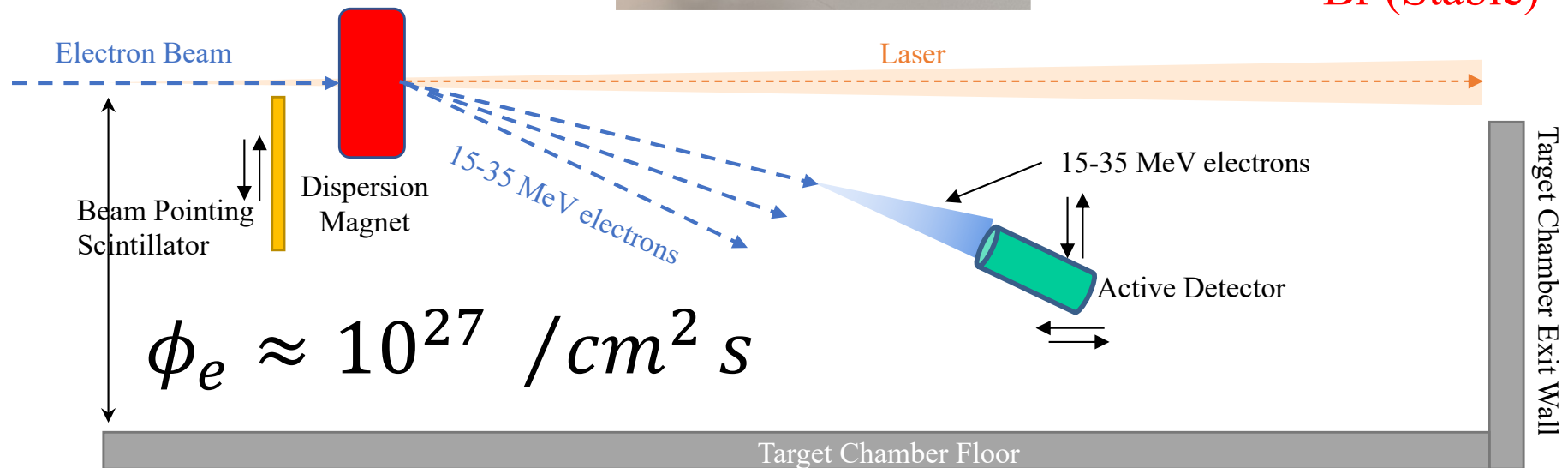
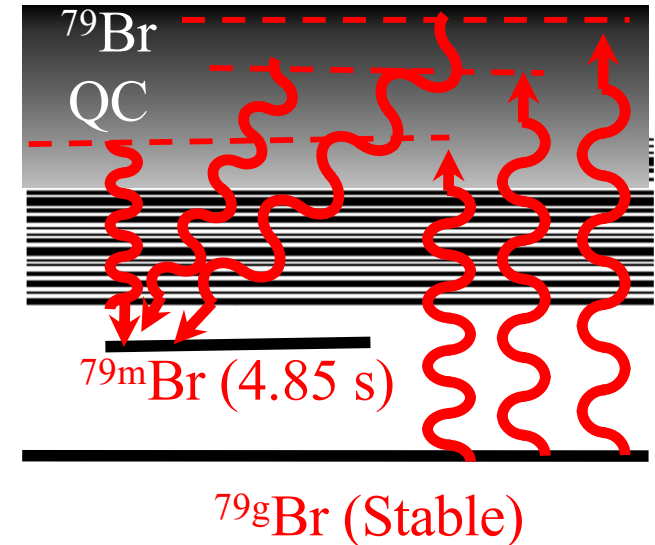
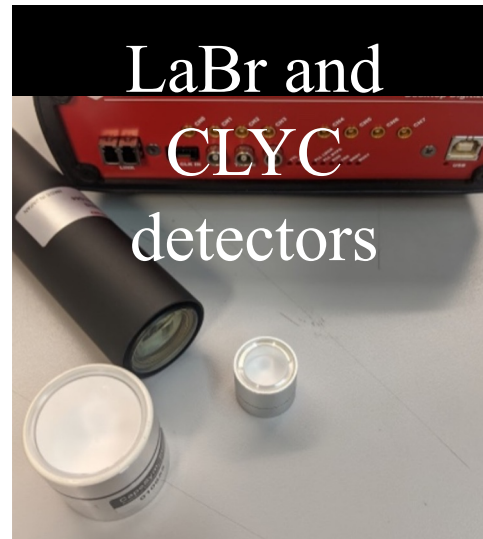
1. Install LaBr detector for use as a  $^{79}\text{Br}$  active targets
2. Install adjustable *Bremsstrahlung* converter (energy selection via geometry, few pC/MeV expected)
3. Use photons to populate  $^{79m}\text{Br}$
4. Repeat at up to 1 Hz & count isomer decays between beam bursts



*This will determine photoexcitation to the QC*

# Part II: $^{79}\text{Br}(e,e')^{79m}\text{Br}$ using the direct electron beam

1. Remove the *bremsstrahlung* radiator and the bending magnet to allow the electrons hit the target
2. Perform the experiment again.
3. Measure the additional isomer population signal to determine the electron-driven NPI contributions.



*The higher flux could allow for reactions on QC states*



# Experimental Set-up @ BELLA

Dispersion Magnet

Gas target

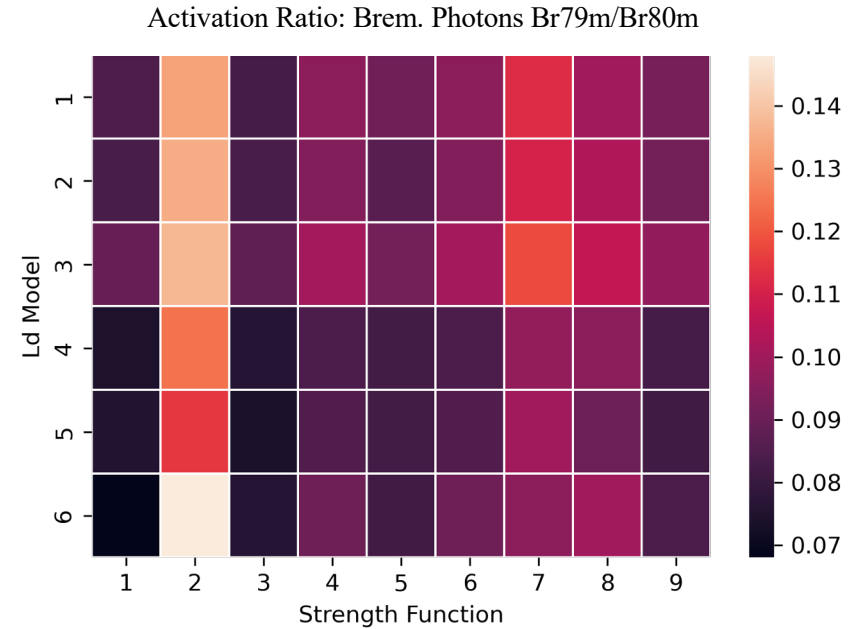
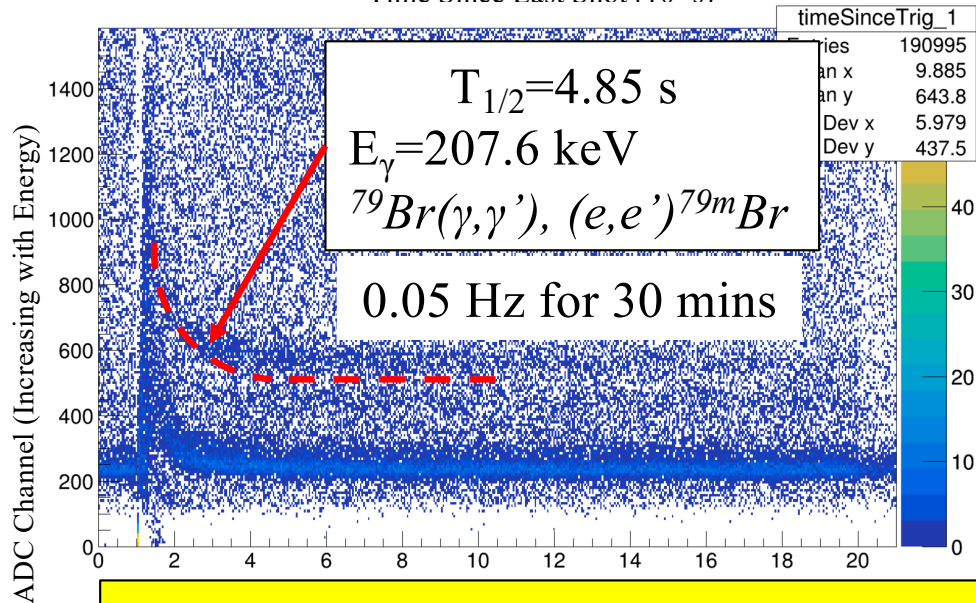
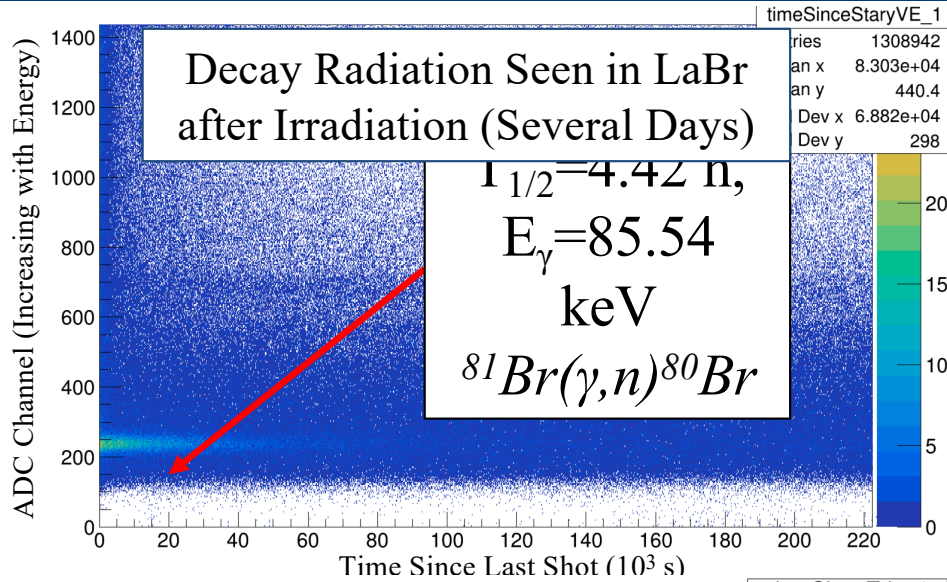
$e^-$  removal magnet

LaBr Detector



Robert Jacob  
(UCB GS)

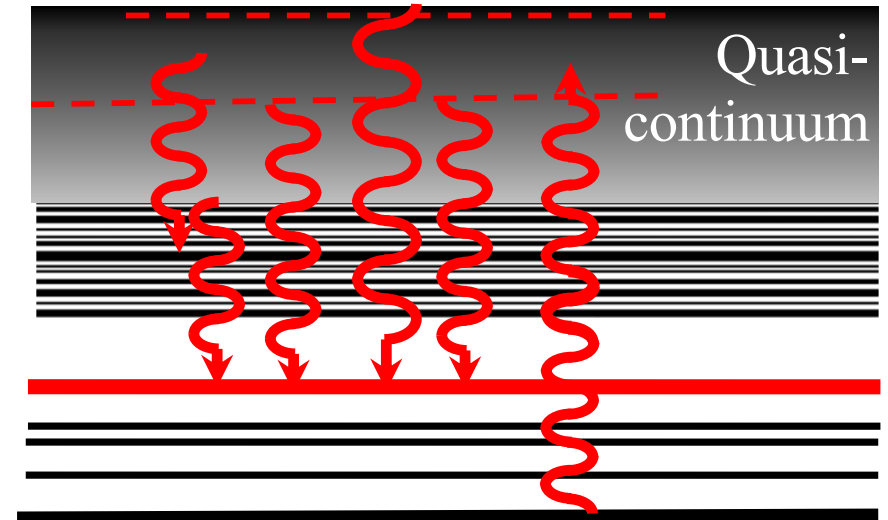
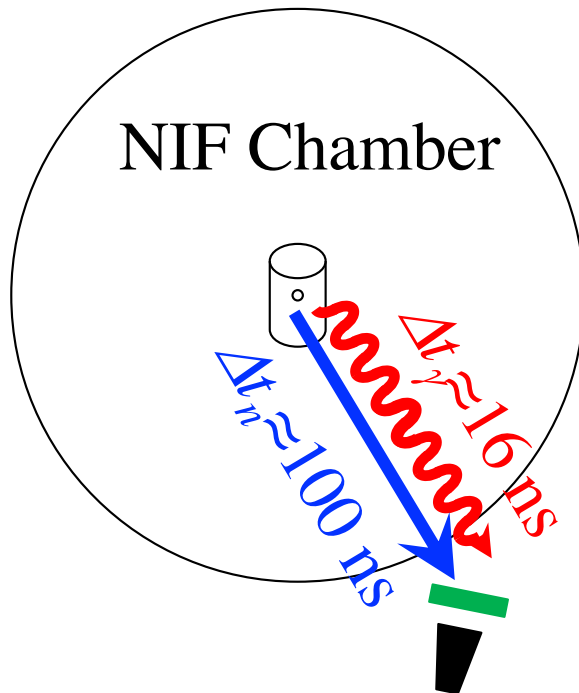
# Data analysis



Irradiation Cases	Ratios (#/Shot)
Brem. Photons 79/80	0.023±0.0024
Electrons 79/80	0.033±0.0015
Brem. 79/Electrons 79	0.503±0.0515
Brem. 80/Electrons 80	0.719±0.0326

*Results point towards a radically different QC J value*

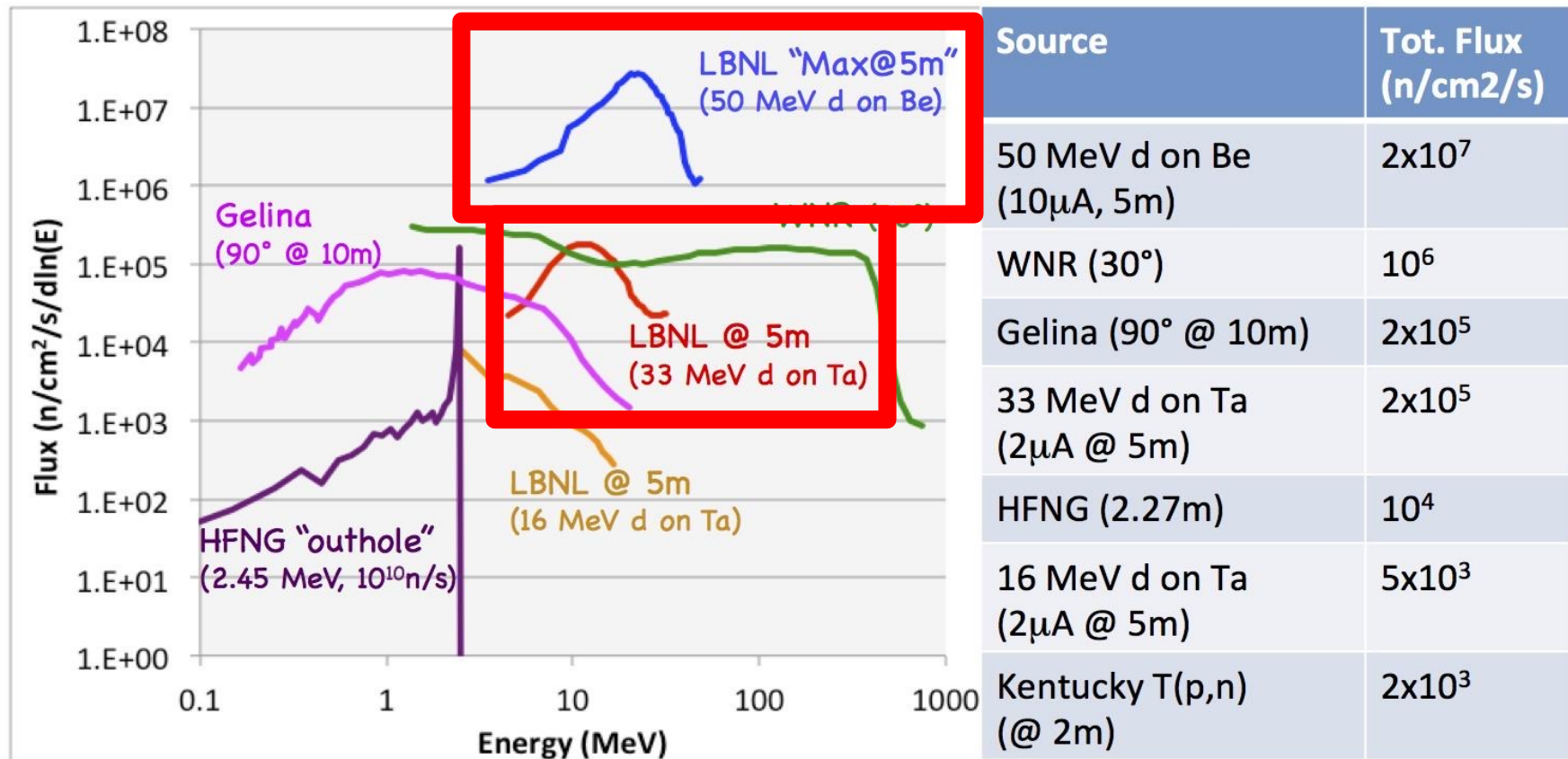
# The population of isomers is a potential new NIF $\gamma$ -diagnostic with a lower threshold than GRH



- Photons reach the chamber wall 84 ns before the 14 MeV neutrons.
- A nuclear excited state with  $\tau \leq 50+ ns$  will give a clear temporal signal.
- Candidates include, Nb, Ta, Y and Zr.
- A greater distance would provide a longer list of candidates

Element	$E_x$ (keV)	$J^\pi$	$t_{1/2}$ (ns)
$^{181}\text{Ta}$	482	$5/2^+$	10.8
$^{93m1}\text{Nb}$	1335	$17/2^+$	14
$^{93m2}\text{Nb}$	1491	$15/2^+$	14
$^{98}\text{Mo}$	735	$0^+$	21.8
$^{91}\text{Zr}$	2288	$(15/2)^-$	29
$^{96}\text{Zr}$	1582	$0^+$	38
$^{90}\text{Zr}$	1761	$0^+$	61.3
$^{94}\text{Mo}$	2956	$8^+$	98

# LBNL hosts one of the most intense 14 MeV-like neutron sources in the country based on thick target deuteron breakup

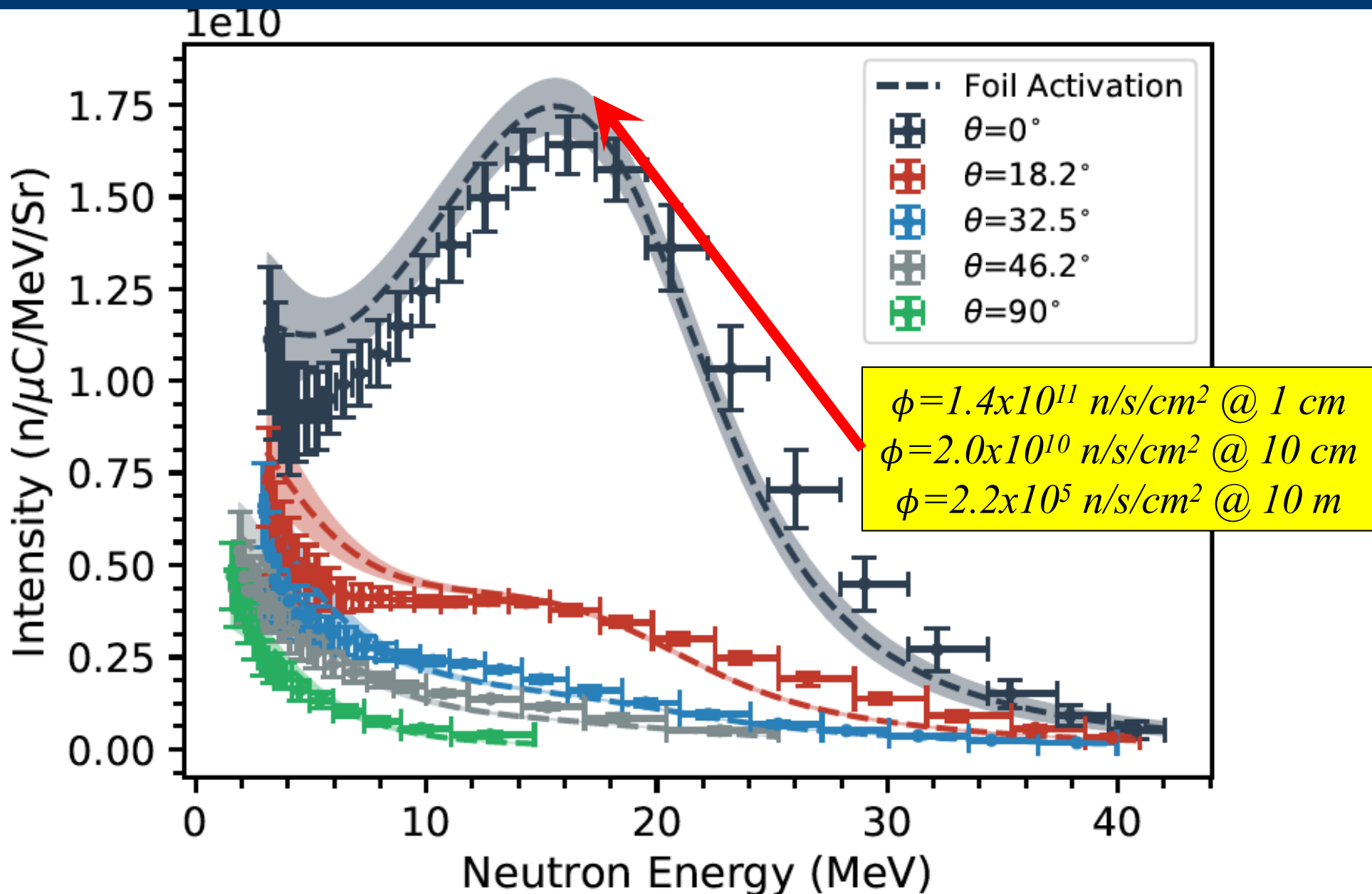


*This neutron source could be used to test and benchmark new neutron diagnostics and to assess neutron damage on optics etc.*

<https://bang.berkeley.edu/events/ndnca/whitepaper/>

*Thanks to Darren Bleuel (LLNL)!*

The TTDB neutron spectrum is adjustable depending on  $E_d$  and the breakup target material (Example: 1 cm beam,  $I_D=20 \mu\text{A}$ )\*



# Some of my students interested in working @ LLNL



**Joey Gordon**

- 4<sup>th</sup> year UCB Ph.D. Student.
- 12/23 Graduation  
Ph.D. project:  
 $^{56}\text{Fe}(n,n'\gamma)$
- Expertise in  
GEANT, neutron/ $\gamma$ -  
ray spectroscopy



**Tyler Nagel**

- 5<sup>th</sup> year UCB Ph.D. Student
- 12/23 Graduation
- Ph.D. project:  
 $^{35}\text{Cl}(n,x)$  using  
CLYC, activation  
and prompt  
neutron/ $\gamma$ -  
spectroscopy



**C. Joe Henderson**

- UCB senior in Nuclear Engineering
- 5/23 Graduation
- B.S. projects: (p,xn $\gamma$ )  
cross sections, Neutron  
air-scatter modeling  
*Coming to NIF this  
summer!*

# Summary

- What NIF can do for Nuclear Physics
  - Nuclear reactions in HEDP environments
  - Using DD shots at NIF to study s-process neutron capture nucleosynthesis in an HEDP
  - The wheel experiments using NIF+ARC to study fission in High Energy Density Plasmas
- What Berkeley can do for NIF & LLNL
  - The 88-Inch cyclotron high-intensity neutron source
  - The BELLA petawatt laser-driven HED facility
    - New gamma diagnostic for  $E_\gamma < 3$  MeV
  - Workforce Development

*Closing note: The White House is holding a one-day meeting on nuclear data for fusion on 5/4 with LLNL representation. Fusion is on the rise 😊*

# Backup slides (not shown)